## Ondřejov Southern Sky Photometry Survey Dynamic Light curves

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### Introduction

- Data
- Continuous identification

## Identification

- Constructing identifiers
- Crossmatching
- Light curve publication
  - SSAP + Datalink

## Practical



- Data from several working groups
  - e.g. small planet light curves
- 99 % of data is not used
- Image coverage
  - Project objectives can move
  - No fixed survey grid
  - Image coverage can shift randomly
- ~ 300 mil measurements (several bands)
- ~ 8 mil sources and growing

# Standard approach

#### **Overlap images (standard)**

Need previous images

Virtual Observatory

- Exactly same regions
- Working with images, not objects
- Differential photometry, astrometry
- Specific for one survey

#### **Process individually (ours)**

- Needs only identifiers
- Images can shift
- Working with actual data on the images
- Aperture photometry, astrometry (individual for each frame)
- Reusable for any image ever taken



# Pipeline

- I. Identify light dots on images with Munipack
  - Astrometry + Photometry for individual images
- 2. Generate common identifiers for actual objects
  - Assigning the light dots to identifiers
- 3. Dynamically construct light curve
  - Possible cutout or any other post-processing



## Problem

Identified objects on one image (Munipack result)





## Problem

Identified objects overlapped (error in astrometry)

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## Problem

Identifying clusters of positions of light curves

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## Example

#### Identifying cluster



# Czech Virtual Observatory

## **Continuous** identification

- Strategy (version 2)
  - I. Generate catalog
  - 2. Cross-match observations
- Continuous identification (updating catalog)
  - I. Is there catalog identifier?
    - No => Add identifier
    - Yes => Cross-match
  - 2. Possible Alerts
    - Mocks is this object a nova?

## Identificators generation

Technology for catalog generation

Virtual Observatory

- PostgreSQL with PgSphere + Q3C (indexing)
- Python for parallelizing CPU-heavy parts

## Process

- I. Pick observation in center of the cluster
- 2. Generate average coordinates from neighbors
- Sophistication question? (need optimalization)
  - Lightweight iterative (95 % accuracy) ~ linear complexity
  - Self-join all (>99,9 % accuracy?) ~ quadratic complexity
  - Need combination

## Lightcurve publication

## SSAP service

**Czech Virtual Observatory** 

- Slightly different from specral data
- ssap.xml?REQUEST=queryData&POS=13.15,-72.86&Band=1
- Result: list of lightcurves

### Datalink post-processing

- Identified by catalog ID (ipix) + Bandpass (U,B,V,...)
- SDM model for light curve transfer
- Cutout on MJD
- dlget?ID=6667183574623977470R
- Result: Cutouted light curve in SDM

## CZZVO Czech Virtual Observatory

## Light curve construction

## Absolutely dynamical (SQL + Python)

SELECT	
	hjd, mag
FROM	
	\schema.objobs_complete
WHERE	
	ipix=
AND	
	band=
ORDER	BY
	hjd



#### SPLAT SSAP 00 52 42.75 -72 48 04.6

Search parameters: Simple Query		
Object: smc	Lookup	
RA: 00 52 42.75	Dec: -72 48 04.6	
Radius: 0.02		
Band:		
Time:		
Query Format:	COMPLIANT	
Wavelength calibration:	None	•
Flux calibration:	None	•



#### SPLAT SSAP result

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#### SPLAT SSAP length

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Whole light curve



2-d compound coordinate system



#### SPLAT datalink

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#### Light curve cutout



2-d compound coordinate system



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#### Aladin SIAP

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Practical



#### Aladin cluster





## Dynamic light curves

Summary

- PostgreSQL access
- SSAP
- Datalink
- SIAP
- SCS



## Acknowledgement

#### Special thanks to Markus



# Technology

- Munipack, Filip Hroch
  - http://munipack.physics.muni.cz/
- Q3C paper, Sergey Koposov
  - https://code.google.com/p/q3c/
- GAVO Dachs, Markus Demleitner
  - http://vo.ari.uni-heidelberg.de/soft/



## Questions

#### Room for discussion