Connecting EPN-TAP and PDS4

S. Erard and the VESPA / Europlanet team

Observatoire de Paris-PSL

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Space missions archives

- Stored / preserved by space agencies
- Common archive format = PDS4 (new missions ≥ 2015)
 - Data products are grouped in datasets (~ collections)
 - Associated documentation included
 - (most archives are still in PDS3, migrating)
- Variations in level of accessibility, in particular granularity
 - ESA at file level (TAP & EPN-TAP interfaces)
 - NASA PDS at collection (dataset) level
 - currently no PDS registry, and no complete catalogue in a dataset

=> how can we search for a particular target / configuration in PDS?

Space missions archives

• PDS4:

- Data file + separated xml label
- Labels include keywords from dictionaries
- Many (\approx 60) dictionaries related to science field / PDS node, etc
- Data file may be a PDS object or a standard format (including fits)
- Q: can we provide an interface to these collections at datafile level via an EPN-TAP table?
- i.e.: can we build an EPN-TAP table from PDS4 information? *(of course, ESA is doing that somehow…)*

Would allow detailed searches for configurations, and cross-searches

Assessment study

- Naive approach, based on 2 examples at NASA PDS Small Bodies Node
 - Two limited collections of asteroid spectra, with slightly different formatting: ascii tables + xml labels + global tables with labels
 - Does not cover all possible situations

=> feasible, but not straightforward

<u>https://vespa.obspm.fr</u> (or via astropy) services spectro_m_ast spectro_trojans service files available at:

https://voparis-gitlab.obspm.fr/vespa/dachs/services/padc/voparis-tap-planeto/spectro_m_ast (via EduTEAMS or ORCID)

First try: M-type asteroids

Initial difficulties:

- The inventory table provides little information => need to parse individual file labels
- Dataset structure doesn't seem to be defined
 parsing the /data dir Need to handle specific subdirectories
- labels do not contain all observational / instrumental parameters (as opposed to fits), which can be provided in a separated table
 other ascii files may contain relevant information ;(

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2-step process

1) Read/parse individual labels

Not using PDS4_tools (xml.etree instead)

2) Add extra metadata from additional catalogues - which are collection-specific

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A /Take	

First try: M-type asteroids

Formal difficulties when parsing labels:

- Only ~10 global EPNCore parameters can be mapped from labels Trial and error process: dictionary may grow with enlarged experience

- target names apparently encoded as "(##) name" but formatting is implicit
- dates: ~ ISO strings but incomplete / variable formatting
- **string values**: case is variable (lower in EPNCore, except names)
- instrument & instrument host names: not standardized, variable (as usual)
- coverages:

Only time is included in the labels => need to parse files to get spectral range

Important quantities such as phase angle, Earth or Sun distances not found ;(
 => call IMCCE ephemcc to complement the EPNCore table

First try: M-type asteroids

access_url: to xml label at SBN - PDS4 tables are supported by TOPCAT

Will be different for other datatypes: add SAMP to PDS4_viewer?



PDS4_viewer

VESPA portal

TOPCAT

Second try: Trojan asteroids

Difficulties:

- Data subdirectory tree is different
- need to grab data from additional data tables (photometry)
 May include non-standard names, to be handled manually
- need to cross-match tables:

With other observations: the only match is (observation time + target) Can be different from spectra (by several h if they are different measurements: spectra and broad filters)

With other documents: instrument/host names etc, are different

- Third try with an image collection: mid-values are found for angles, coordinates, distances in labels — not ranges

Conclusion

Conversion is very feasible - and done

- Data structure is not normalized in PDS4
- Parameter names and values are not entirely standardized
- Only some parameters currently mapped, need to check with new trials
- Coverages are not easily found, except time
- Information is often spread on several files, difficult to merge. Reference to granules not always clear (may be through observing time, with tolerance)

=> Each dataset is a new project, requires hand tuning

• Need further checks with images — Should be easier with embedded fits files