

SED Data Model Status

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Nara 2010

Status:

SED model development descope'd in ~2005 to focus on single Spectrum model, approved in 2007

Ideas for Photometry model presented in DM working group and in US NVO meetings from 2007-2009

Proposal is to reuse Spectrum model for Photometry with minor additions

Proposal is to define the SED as an aggregation of Photometry and Spectrum segments

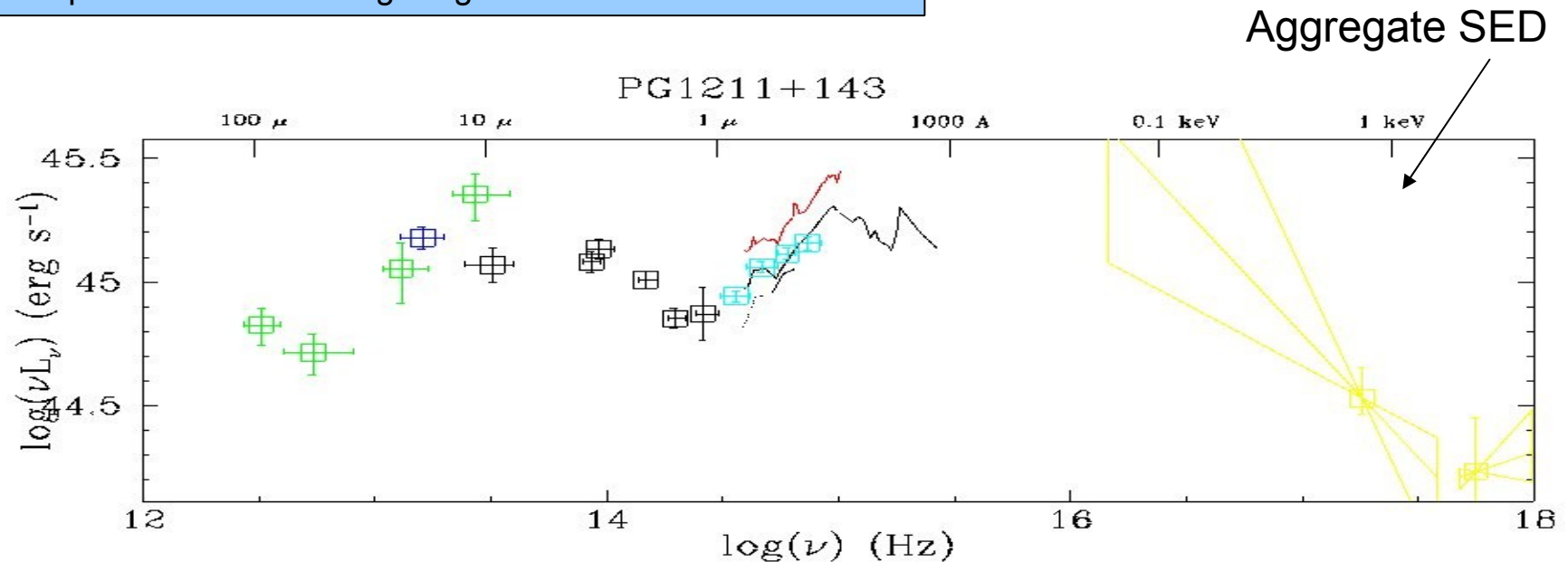
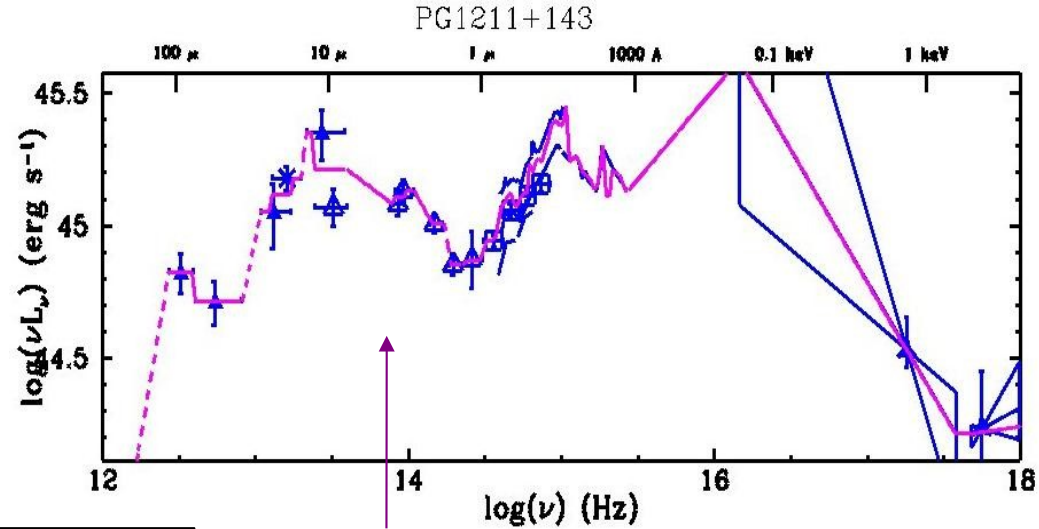
Building the SED

- Assemble **aggregate SED**

- convert each segment to VO format
- collect in one file with all metadata

- Create **uniform SED**

- time average/blend phot. and spectra in each wavelength bin
- interpolate across missing ranges?



Name: Q1211+143
 QSO Name: PG1211+143
 Position: 12 11 44.8 14 19 53.0
 Redshift: 0.0850 V mag: 14.63
 NH value: 283.0
 Sample: UV/WE/HL/EXO

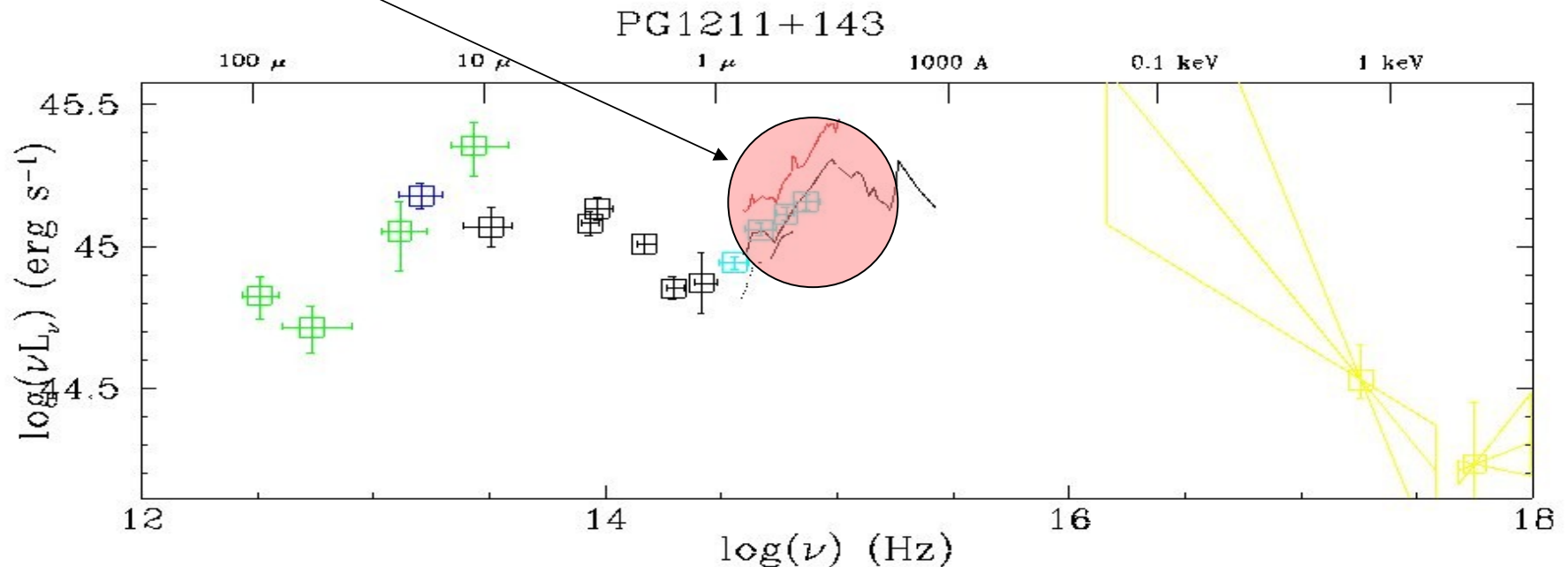
—	--- Averaged	—
—	1980 Jun Bechtold et al. (1986)	Hale
—	1980 Jun Bechtold et al. (1986)	IPC
—	1983 --- Elvis et al. (1990)	IRAS
—	1985 May Elvis et al. (1990)	MtHopk(24in)
—	1986 Feb Elvis et al. (1990)	IRTF
—	1987 Feb RDAF	LWP10115
—	1988 Sep Elvis et al. (1990)	MMT/FOGS

The hard part: combining overlapping data

Overlapping spectra and photometry is a common situation

What is the SED flux value here?

Implication: uniform SED data points don't have a 1-1 mapping back to aggregate SED segments.



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Photometry Bands

- * Many definitions of “effective wavelength”
 - transmission weighted
 - reference spectrum weighted
 - isophotal
 - pivot
 - etc, etc...
- * Often just use a fiducial wavelength
 - Degeneracy between choice of wavelength and value of zero point and color correction.
 - e.g. use same wavelength for all V filters: easier to compare derived fluxes

- * Magnitude equation:

$$F(\lambda_0) = F_0 * 10^{**(-0.4(m-m_0-C))}$$

where

F₀ is the flux at $\lambda=\lambda_0$ for an object with the reference spectrum $f_R(\lambda)$ and magnitude m_0 .

C is the color correction

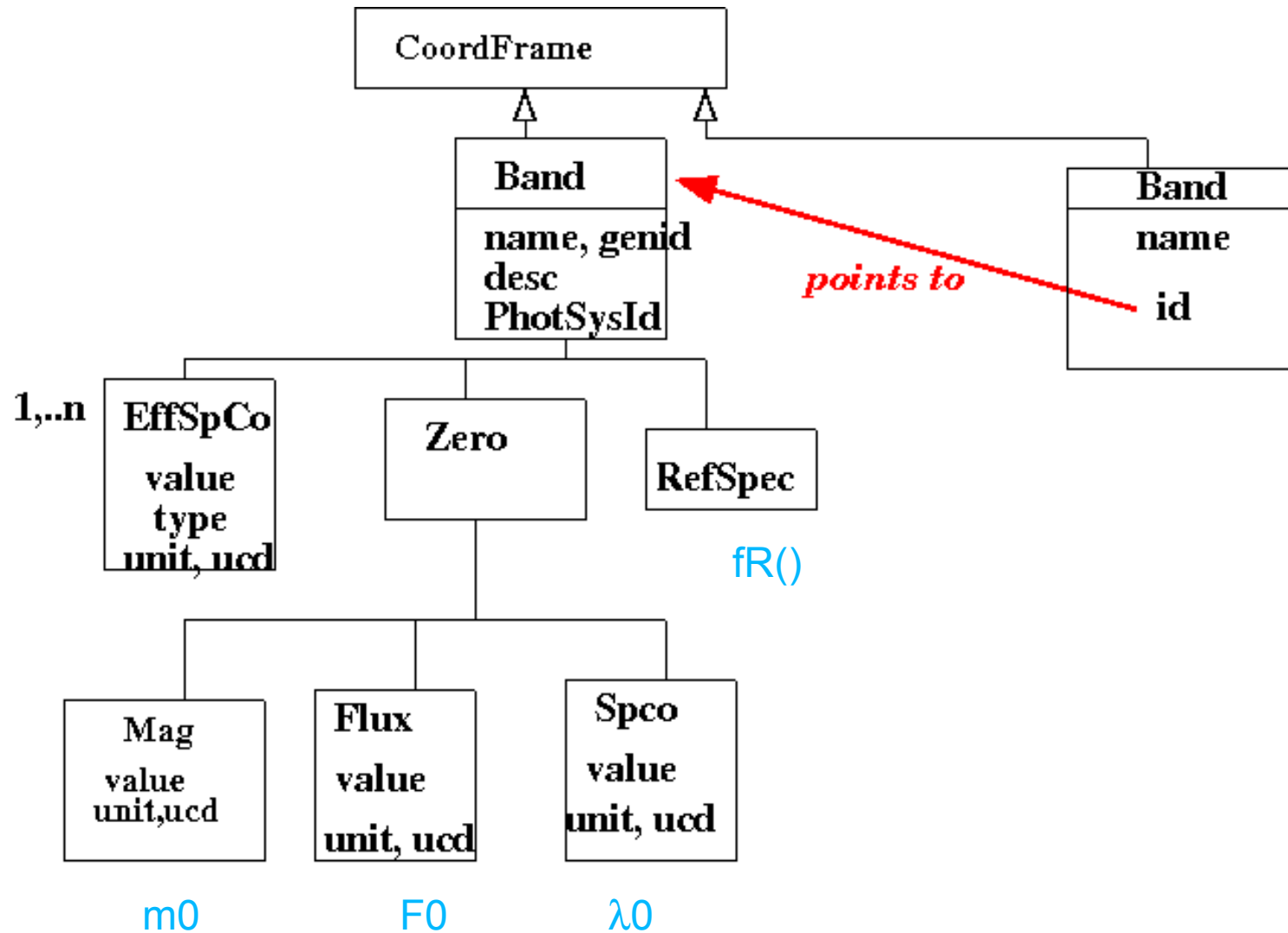
$$C = -2.5 \log c, \quad c = \frac{\int f(\lambda)T(\lambda) d\lambda}{\int f_R(\lambda)T(\lambda)d\lambda}$$

comparing the object spectrum with the reference spectrum.

Band parameters:

λ_0 fiducial wavelength
F₀ zero point flux at fid. wave.
m₀ zero point mag (usually 0.0)
T(λ) transmission curve
f_R(λ) reference spectrum


Proposed Band model: use as a coordinate system on the Flux axis
 Can exist in a Spectrum CoordSys, in full or as a reference to an external database of bands by (name, id)



Allow data providers to also supply their favorite effective wavelength definition (optionally) – or even more than one

```
<GROUP utype="CoordSys">  
  <GROUP utype="Band">  
    <GROUP utype="Band.EffSpCo">  
      <PARAM utype="Value" value="5400.0"/>  
      <PARAM utype="Type" value="MLAM"/>  
      <PARAM utype="Unit" value="Angstrom"/>  
      <PARAM utype="UCD" value="em.wl"/>  
    </GROUP>  
  </GROUP>  
</GROUP>
```

What flavor of effective wavelength is this? Allowed values:
EFLAM, EPLAM, MLAM, ISO, EFNU, EPNU, MNU, ISONU, PIVOT



Aperture corrections

Propose a new element within the CharacterizationAxis object:
Correction

First use:

FluxAxis.Correction.ApFrac.Value
FluxAxis.Correction.ApFrac.Applied

- Define aperture fraction from 0 to 1 as fraction of source total flux contributing to measurement
- NOT necessarily point source fraction (e.g. measure part of spiral galaxy, use spatial model to extrapolate total flux)
- Measured Flux = ApFrac * Total Flux
- Include flag to say whether the number in Data.FluxAxis.Value is the Measured Flux or the Total Flux, i.e.: has the correction been applied? Both cases occur

Can envisage other cases of Correction to various Char axes

Proposal is to include the transmission as part of the Characterization. It is simply the 4th Coverage level on the spectral axis as discussed in the IVOA Char standard: degree of exposure as a function of spectral coordinate. It would be a URI pointing to the band def:

