## CoRoT and Kepler time series Focus session on time domain astronomy

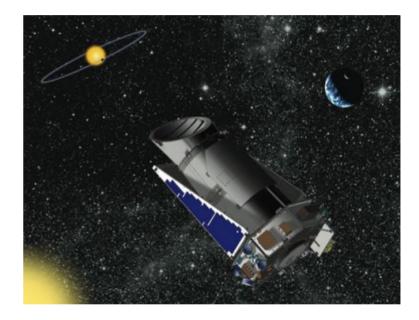
J. Debosscher Instituut voor Sterrenkunde, KULeuven, Belgium

IVOA Interoperability meeting, Heidelberg, 12-17 May, 2013

## CoRoT



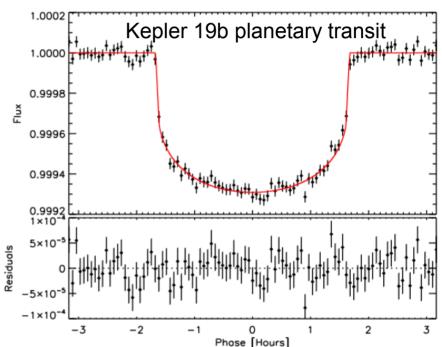


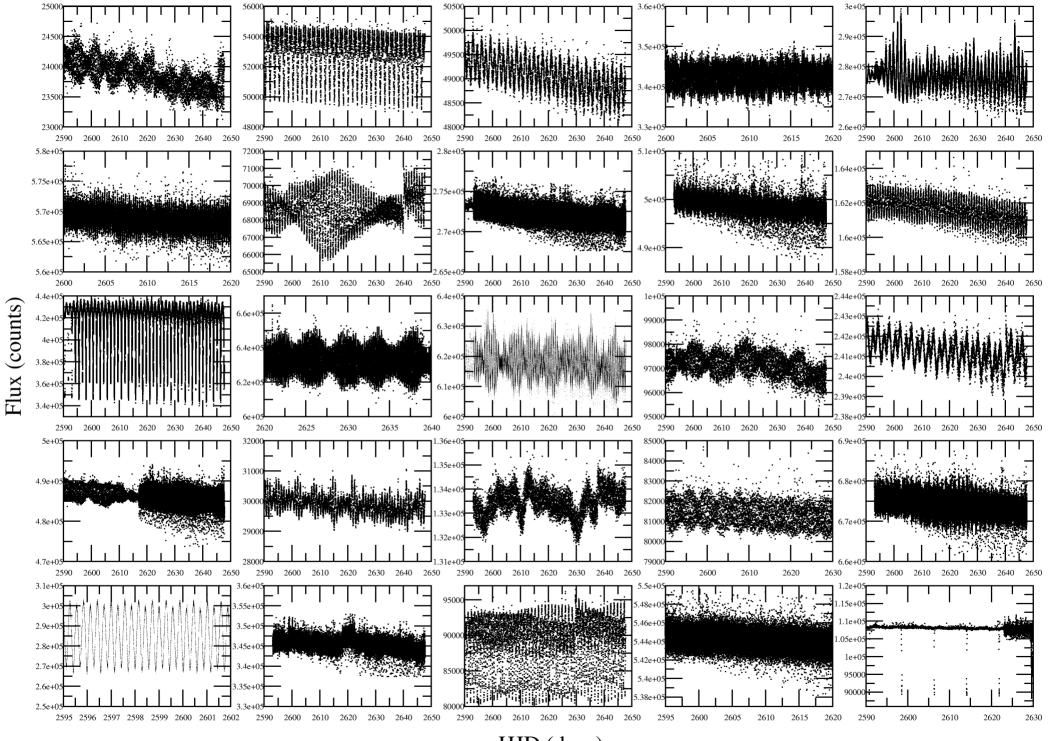


- Polar orbit
- 27 cm mirror telescope
- visual light
- 4 2k x 2k CCD imagers
- Several observed fields (each about 3 x 3 deg)

- Earth trailing orbit
- 0.95 m mirror telescope
- visual light
- 42 2k x 1k CCD imagers
- 1 observed field (105 deg<sup>2</sup>)

- CoRoT
  - Launched Dec. 2006, no longer producing science data since Nov. 2012, recovery efforts ongoing
  - <u>CO</u>nvection <u>RO</u>tation and planetary <u>Transits</u>
  - Asteroseismology, search for exoplanets
  - ~160.000 stars
  - 32 or 512 sec. sampling, max 150 days
- Kepler
  - Launched March 2009 (until 2016 at least)
  - Exoplanet detection using transit method
  - ~156.000 stars
  - 1 or 30 min sampling, observing quarters of 90 days (up to 7 years total!)
  - Both missions deliver densely sampled light curves with high photometric precision ---> unique data sets for stellar variability studies!





HJD (days)

## Time series (light curve) data

- Data volumes (minimum): ~1TB for CoRoT, ~ 10TB for Kepler (will be much more for e.g. LSST and Gaia!)
- Minimal data delivery for users: target ID, coordinates, stellar flux time series (ASCII or FITS format)
- Often more information is needed (usually present in FITS files headers)
  - Instrument parameters (e.g. to identify artifacts in the data)
  - Data reduction information, e.g. background flux, noise levels, quality flags...
  - Additional time columns, e.g. barycentric or heliocentric Julian dates
  - Additional flux columns (different data reduction or extraction)
  - Complementary data from other databases & higher level metadata

#### CoRoT Data Center at IAS: http://idoc-corot.ias.u-psud.fr/ Downloading CoRoT light curves per observed run

You can retrieve the N2 data by downloading a whole set for a given run (last version available) :

Runs	Light curves of the Astero channel Monochromatic light cur (AN2_STAR) Exo channel (EN2_STA				Chromatic light curves of channel (EN2_STAR_C		-WINDESCRIPTOR Exo channel	the Exe	Light curves from imagettes files of the Exo channel (EN2_STAR_IMAG and associated WINDESCRIPTOR)				
IRa01	here     [MD5]     here     [MD5]       (94MB)     [SHA-1]     (1,5GB)     [SHA-1]		here [MD5] (5,9GB) [SHA-1]		here [MD5] (40MB) [SHA-1]		here [MD5 (289MB) [SHA-						
SRc01	here (44MB)	[MD5] [SHA-1]	he		here [MDS	- 1		D5]	here [MD5				
LRc01	(251MB) [SHA-1]				s with FITS files, format depends on type of light								
LRa01	1 here [MD5] (217MB) [SHA-1] CURVES (some are multi-colour)												
SRa01	Select	_ DAT 23A	□ DATE □ 0 23A		DATEHEL	STATUS	REDFLUX E	REDFLUXDEV E	GREENFLUX E	GREENFLUXDEV	BLUEFLUX		
LRc02	🗌 Ali Invert	yyyy-mm-jjTh	h:mi:ss	COROT JULIAN DAY	COROT JULIAN DAY		ELECTRONS	ELECTRONS	ELECTRONS	ELECTRONS	ELECTRONS		
SRc02		1 <u>]</u>											
	1	2007-02-03T13:		2.590045509073 <b>E</b> +03	2.590049982487E+03	9		9.711546E+03	1.025038E+05	3.642345E+03	1.581949E+05		
SRa02	2	2007-02-03T13:		2.590051435029E+03	2.590055908168E+03	0		7.956298E+03	1.022476E+05	3.063443E+03	1.570602E+05		
	3	2007-02-03T13:		2.590057360986E+03	2.590061833848E+03	16		9.009035E+03	1.020234E+05	3.545518E+03	1.569270E+05		
LRa02	4	2007-02-03T13:		2.590063286942E+03	2.590067759528E+03	0		1.485986E+04	1.022411E+05	5.587126E+03	1.569962E+05		
LRc03	5	2007-02-03T13:		2.590069212898E+03	2.590073685208E+03	0		9.390163E+03	1.022943E+05	3.574360E+03	1.571913E+05		
Encos	6	2007-02-03T13:		2.590075138855E+03	2.590079610888E+03	0		1.019362E+04	1.024649E+05	3.994630E+03	1.571392E+05		
LRc04	7	2007-02-03T13:		2.590081064812E+03	2.590085536568E+03	0		9.981426E+03	1.024448E+05	3.798916E+03	1.572613E+05		
L	8	2007-02-03T14:		2.590086990768E+03	2.590091462249E+03	0		6.345261E+03	1.024935E+05	2.634932E+03	1.573250E+05		
LRa03	9	2007-02-03T14:		2.590092916725E+03	2.590097387929E+03	36		9.293336E+03	1.022291£+05	3.844240E+03	1.569177E+05		
	10	2007-02-03T14:		2.590098842682E+03	2.590103313609E+03	0		9.614719E+03	1.022442E+05	3.821578E+03	1.573274E+05		
SRa03	11	2007-02-03T14:		2.590104768639E+03	2.590109239290E+03	0		7.878013E+03	1.023414E+05	3.187052E+03	1.574162E+05		
LRc05	12	2007-02-03T14:		2.590110694595E+03	2.590115164969E+03	0		8.726795E+03	1.023315E+05	3.298300E+03	1.573521E+05		
	13	2007-02-03T14 :-		2.590116620552E+03	2.590121090649E+03	8		1.577869E+04	1.023986E+05	5.469698E+03	1.571704E+05		
LRc06	14	2007-02-03T14:	122 / - C - C - C - C - C - C - C - C - C -	2.590122546508E+03 2.590128472464E+03	2.590127016329E+03 2.590132942008E+03	16		7.198164E+03	1.023130E+05 1.021604E+05	3.044901E+03 2.989277E+03	1.570422E+05 1.568566E+05		
	15	2007-02-03T15: 2007-02-03T15:		2.590128472464£+03 2.590134398421£+03	2.590132942008E+03	16		7.461863E+03 9.245952E+03	1.021604£+05	2.989277E+03 3.232375E+03	1.572059E+05		
LRa04	17	2007-02-03115: 2007-02-03T15:		2.590134398421E+03 2.590140324377E+03	2.590138887888E+03			9.243952£+03 8.840103£+03	1.022329E+05	3. 370405E+03	1.572425E+05		
LRa05	18	2007-02-03115: 2007-02-03T15:		2.590146250334E+03	2.590144793388E+03			7.733802E+03	1.022215E+05	3.011939E+03	1.570797E+05		
LKaUJ	19	2007-02-03115: 2007-02-03T15:		2.590152176291E+03	2.590156644727E+03			1.069012E+04	1.023017E+05	3.961668E+03	1.571807E+05		
LRc07	20	2007-02-03T15:		2.590158102247E+03	2.590162570407E+03	0		5.486179E+03	1.022811E+05	2.051909E+03	1.573224E+05		
SRc03	Not A	vailable		Not Available	19		838		(4.6MB) [SHA-	1]			
LRc08									here [MD5 (419MB) [SHA-				
SRc03	]												
SRa05			<b>he</b> (1,3)		here [MD5 (3,3GB) [SHA-			D5] A-1]	here [MDS (180MB) [SHA-				

## Selecting light curves based on their properties and/or target properties obtained from external observations

Available products		
Chromatic light curves 🛛	AND/OR	Selection based on target
Oversampled light curves (32s)	AND/OR	properties such as spectral t
Light curves from imagettes (32s)		from other than CoRoT data
N2 context		*
Observational informations		
Start date	(DD-MM-YYYY)	
Corot ID	Use semi-colon, comma or blank cha	ration as a
Corot ID	× All	Selection based on classific
Magnitude from (>=)		results obtained from the lig
Right ascension from		curves
Declination from		<b>A</b>
Magnitude difference (B-V) from		
Spectral type from	The spectral type combined in the headers concerning luminosity classes.	is based on a SED or isochrone analysis of multi-col
Spectral type to 💷 🗘	For more information please read Delexil	et al. and if needed please contact Claire Moutou.
Luminosity class		
Long Run(and initial Run) 🗹		
Short run 🖌		
Run	0	
Selection based on the class of variability of the star		

#### Higher level data products: light curve classification

	CoRoT Public N3 Products
Gat	COROT FUDIIC NO FIODUCIS
-	BACK TO HOME
~	BACKTOHOME

#### THE COROT VARIABLE CLASSIFIER

The COROT Variable Classification is based on an analysis of the COROT light curves described in Debosscher et al. The resulting classification is

Туре	Run IRa01	Run SRc01	Run LRc01	Run LRa01	Run SRa01	Run LRc02	Run SRc02
	(Updated : 19/02/2013)	(Updated : 19/02					
N3 CVC	here	here	here	here	<b>here</b>	here	here
	(9,3MB)	(6,6MB)	(11MB)	(11MB)	(7,7MB)	(11MB)	(11MB)

Column description:

-----

1) objectname=CoRoT ID

- 2) normalized Mahalanobis distance to the centre of the most probable class (class 1)
- 3) normalized Mahalanobis distance to the centre of the second most probable class (class 2)
- 4) normalized Mahalanobis distance to the centre of the third most probable class (class 3)

5) classprob1=relative probability for class 1

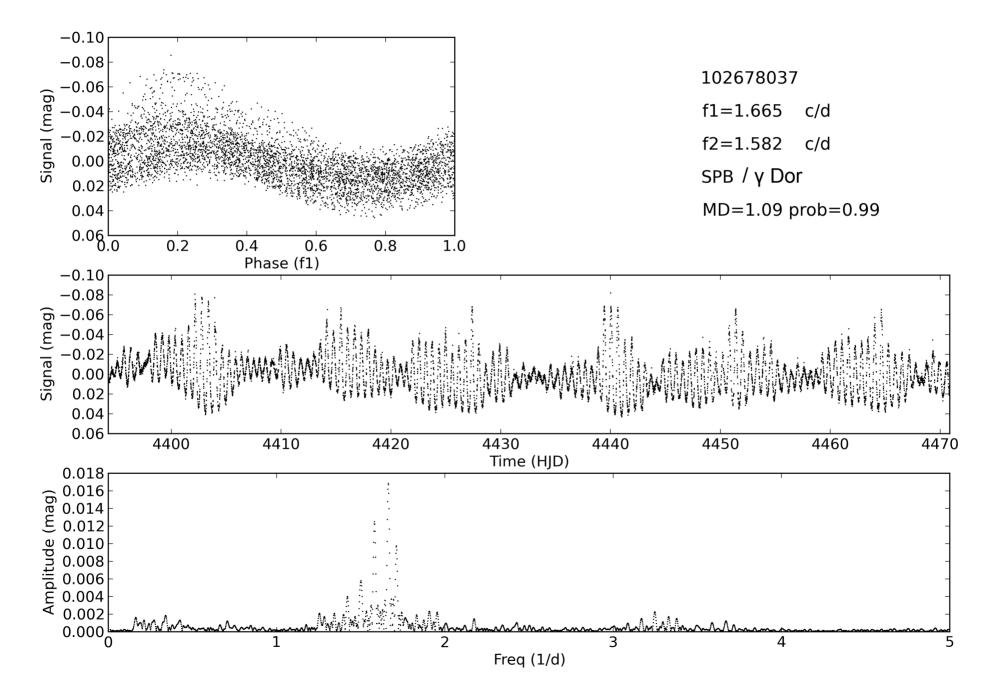
6) classprob2=relative probability for class 2

- 7) classprob3=relative probability for class 3
- 8) classcode1=variability class 1
- 9) classcode2=variability class 2
- 10) classcode3=variability class 3
- 11) Pf1=significance parameter frequency 1 (probability)
- 12)Pf2=significance parameter frequency 2 (probability)
- 13)Pf3=significance parameter frequency 3 (probability)

Results from an automated supervised light curve classifier, very useful for selecting science targets!

Debosscher et al., 2007, 2009

#### Light curve classification example



#### CoRoT data tools at the Spanish VO http://sdc.cab.inta-csic.es/corotfa/

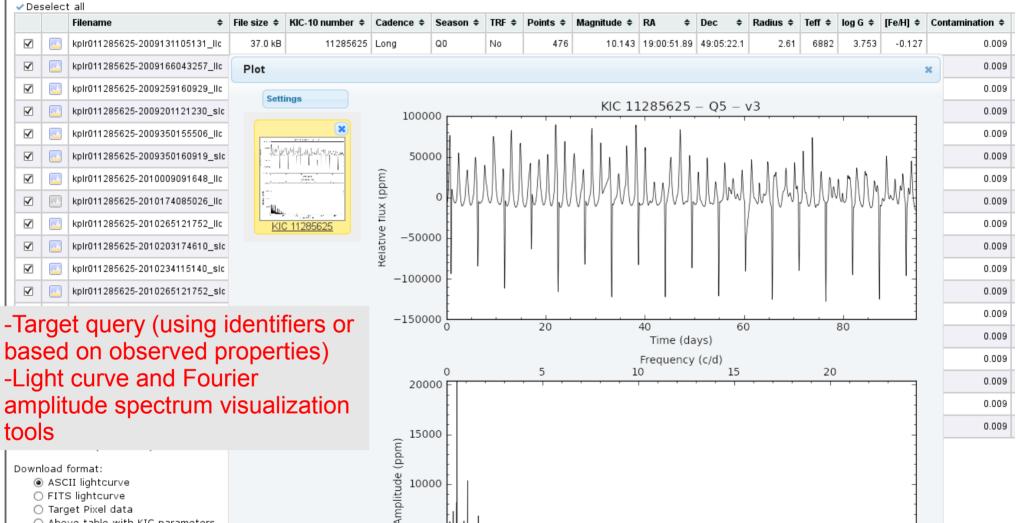
	THE COROT PUBLIC A	RCHIVE AT CAB
Run: IRa01 V Data Type • Asteroseismology V Light curves V	■ Exoplanet ✓ Monochromatic light Curves ✓ Chromatic Light Curves ✓ Variability Type: Low Frecuency □ (If not checked, by default: 'All Frecuency')	
Search Corot ID:	Object ID:	Coordinates List:
<u>Light Curve Filter Criteria</u> Obs Date:	From: 31 - January V - 2007 V Vmag: -	To: 09 - January V - 2012 V B-V: -
Spectype:	Lumclass:	Variability class (Important note): (Only for exoplanet data) ACT [Activity] BCEP [Beta-Cephei stars] BE [Variable Be-stars] CLCEP [Classical Cepheids] CP [Chemically peculiar stars]

														THE COROT ARCH			
Ret	Retrieval Format: zip 🗸 Mark Fits: 🗌 Mark ASCII: RAW: 🗌 HEL: 🗌 HELREG: 🗌																
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IRa	1 20	HD 49933	102.708	-0.54088	2007-01-31 11:06:34.0	2007-04-02 07:12:15.0	F	v									
IRa	1 83	HD 50773	103.654	-0.45264	2007-02-03 13:04:59.0	2007-04-02 07:12:15.0	A			298201,865	2590,0	9 2595,85	2601,62	2607,39	2613,16 2618,92	2624,69 2630,46	2000,23 2042 2047,70
IRa	1 116	HD 50747	103.603	-1.12698	2007-01-31 11:06:34.0	2007-04-02 07:12:15.0	A	IV		SIMPLE = T ///ritten by IDL: Thu Dec 6 17:31:53 2007 BITPIX = 8/ NAXIS = 0/ EXTEND = T //File contains extensions TELESCOP- 'COROT ' //Felescope name ORIGIN = 'CDC ' /processing site CREA_DAT='2007-12-06T17:31:53'/file generation date FILENAME= 'N/2_STAR_000000083_20070203T130459_20070402T071215 PIPE_VER='0.2/1.0' /name/version of the process STARTDAT='2007-02-03T13:04:59'/UT date of the first meas. END_DATE='2007-04-02T07:12:15'/UT date of the last meas. COROTID =' 83' /ID of the CoRoT target RUN_CODE='RA01' / /code of the run							
IRa	1 123	HD 50844	103.709	-1.07078	2007-02-03 13:04:59.0	2007-04-02 07:12:15.0	А										
IRa	1 156	HD 50846	103.728	-1.37579	2007-02-03 13:04:59.0	2007-04-02 07:12:15.0	в			HLFCCDID='A1R ' /ref. of the half ccd ALPHA = 103.654 /Right ascension of the star DELTA = -0.452640 /Decination of the star STARNAME='HD 50773' /Usual name of the star (COROTSKY) MAGNIT_V= ABSM_V = COL_B_V = SPECTYPE= Interactive light curve Show Data							Clean Raw Hel Helreg
IRa	1 187	HD 50405	103.262	-1.88363	2007-02-06 13:34:53.0	2007-04-02 07:12:15.0	A										
IRa	1 214	HD 51106	103.976	-1.58537	2007-01-31 11:06:34.0	2007-04-02 07:12:15.0	A		7.36	0.23	8027.0	3.88	0.37	FITS	FITS	raw	
IRa	1 223	HD 50170	102.966	-2.17604	2007-01-31 11:06:34.0	2007-04-02 07:12:15.0	F		6.82	0.46	6560.0		-0.24	FITS	FITS	raw	
IRa	1 263	HD 292790	103.686	-2.12311	2007-02-06 13:34:53.0	2007-04-02 07:12:15.0	F		9.48	0.45	6477.0			FITS	FITS	raw	
IRa	1 400	HD 50890	103.745	-2.80359	2007-02-06 13:34:53.0	2007-04-02 07:12:15.0	G	ш	6.04	1.1	4651.0		-0.18	FITS	FITS	raw	

EXOPLANET

Found 4990 records, displaying page 1 of 100

#### Kepler data at KASOC (Kepler Asteroseismic Science Operations Center) http://kasoc.phys.au.dk/



50

100

150

Frequency (µHz)

200

250

5000

ASCII lightcurve

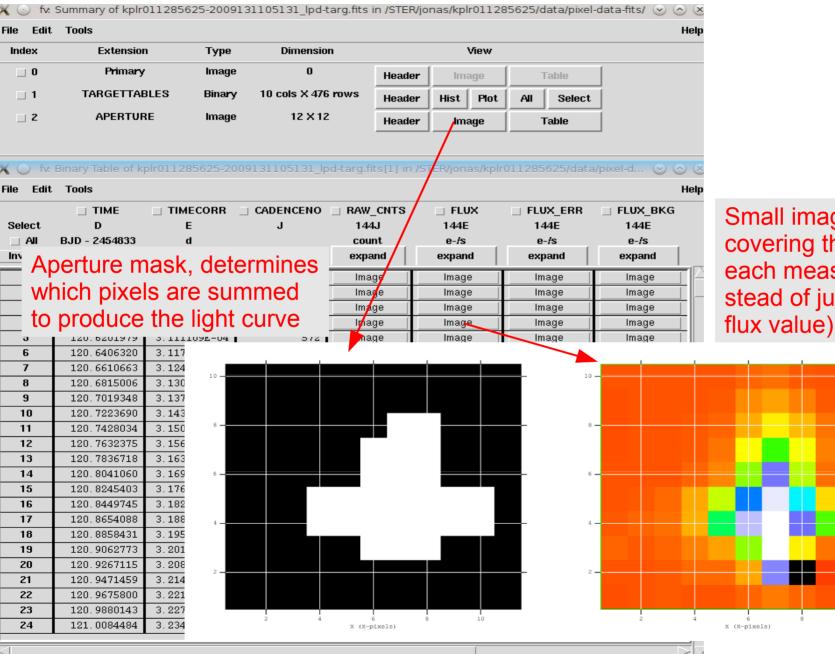
- O FITS lightcurve
- O Target Pixel data
- O Above table with KIC parameters
- O Plots of timeseries and amp. spectr
- wget script (ASCII)
- wget script (FITS)
- wqet script (TPD)

🖢 Download Selected

# Kepler data at the MAST archive http://archive.stsci.edu/kepler/data search/search.php

Archive Status	Kepler Data	Search & Re	trieval <sub>Fi</sub>	<u>Help</u> eld Descriptions
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Output				

### Extended light curve data products: Kepler pixel data

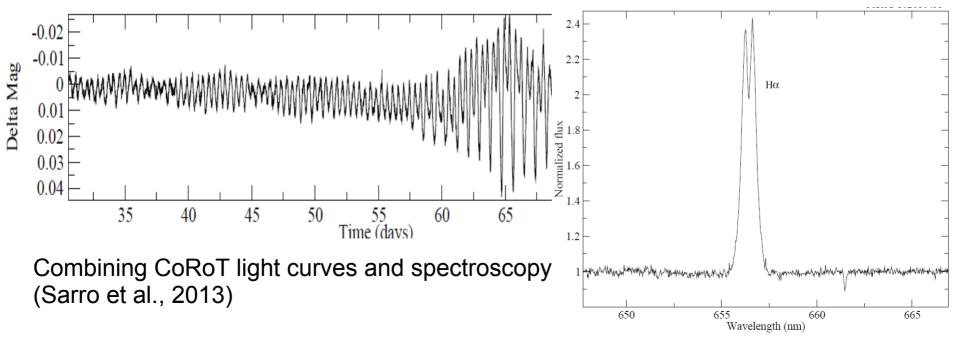


Small image of CCD pixels covering the target, for each measurement (in stead of just one summed flux value)

10

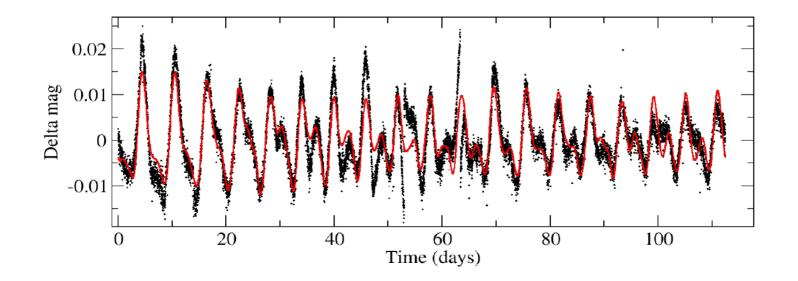
## How can VO help for stellar variability studies?

- Time series data (single bandpass) are not sufficient to fully characterize and study stars!
  - Juse VO to complement time series with other observational and higher level data, such as highresolution spectroscopy, photometry in other bands, temperatures, gravities, parallaxes...



## Light curve analysis tools in the VO?

- Some commonly used light curve analysis tools:
  - Period determination by means of Fourier amplitude spectrum (e.g. for pulsating stars) or box-least-squares periodogram (e.g. for planetary transits)
  - Filtering and detrending methods to reduce instrumental artifacts or disentangle variability at different time scales
  - Linear and non-linear least-squares fitting (e.g. of sine functions)



## Stellar variability data mining and the VO

- Combining large observational data sets for data mining purposes, constructing training sets for light curve classifiers
- Data mining tools such as dimensionality reduction and clustering techniques in the VO?
  - scalability of methods with future data amounts in mind

