#### **SSA Revisited**

# (Also SDM and SODA suggestions)

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And

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IVOA Interoperability meeting , DAL Session 1 Shanghai, China, 15th May 2017

### **Motivation - LPV**





#### V436 Per Janík 2003

Rho Pup, Aerts et al.

# Variability of Line Profile on Be star



# **Spectra Postprocessing Service**

- Normalization rectification (FT, Crosscor, RV, LPV)
- Cutouts
  - Data volume, several lines short ranges
- Future development
  - Rebinning
  - Instrument profile convolution
  - Broadening functions (rotation, limb dark) for TSAP
  - RV/z shift
  - Combining diiferent spectra resolution power Using SPECRP

# Pros and Cons of current DaCHS DataLink

- Two step process lost query params in accref
- Separated query and process params

- You cannot simply request only Halpha line from many services - complication in
  - Client must remember big list in VOTable
  - Can it work for 8 mil records in 1 VOTable ?
  - How query sky instead of all DR4 FITS files
- IMPOSSIBLY SLOW for VOCLOUD download

# LAMOST (Guoshoujing)

Xinglong- China 4m mirror (30 deg meridian) 4000 fibers 10 mil spectra / 5 yr Automatic RV-z





### VO-compatible LAMOST Archive in Ondřejov







Thanks: Chenzhou Cui + all China-VO and LAMOST team Jiří Nádvorník and Markus Demleitner

# LAMOST with VO in Ondřejov

- Download all FITS in multiple subdirs
- Prepare DaCHS RD
- Tricks during import
  - Convert from log lamda to lambda
  - Normalized accessed by FITS extension
- DL + SODA
  - Cutout in lambda (need auto conversion to vot)
  - Redshift (using computed z what with bad ?)
- Tested on SPLAT-VO + DaCHS

### **LAMOST emission candidate**



10000 Elux (count) 

# **LAMOST Query Ha region**

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# **LAMOST** emission candidate



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# **SSA** issues

- Query by POS is unsufficient
  - Theoretical spectra need different approach
  - Observed no coordinates available
    - Secret exoplanets Kepler 15, TRESS-1
    - Multiple stars on slit (extraction of two files)

- HR1847A,B - bad seeing - A+B

- Imprecise telescope coordinates
- Exotic objects solar system or clusters (WEBDA 17) - finding charts in galaxies
- Must support TARGET NAME (but WILDCARD, Sgr A \*)
- WHEN Name resolver not available ????
- How query by BAND echelle with multiple orders (e.g. BeSS,STIS) and how to CUT by SODA ?

# **SSA** issues

- Original format NATIVE (IRAF WCS -image/fits)
- LAMOST, SDSS in log lambda (better for precision)
- VO supported format is VOTABLE or application/fits
  - Binary table
  - Loosing dispersion function (processing)
- To return results must create virtual accref + conversion
  - VOTable link + image/fits link
  - Query by FORMAT = COMPLIANT only VOT
  - Default (implied FORMAT) or FORMAT=ALL both

#### **SSA + DL + SODA issues**

- If FORMAT=image/fits (Native)
  How to convert by DL SODA FORMAT processing ?
- If native should return native ?
  - Image/fits understood by client SPLAT-VO only
- Should convert to votable to process
- If application/fits should process ? It is VO supported !
- If Result is votable to SODA FORMAT app/fits convert ?
- DL + SODA does not give names and extensions to files
  - Problem if using DL to create files for ML (lists )

# **SSA + SDM issues**

- SSAP requires VACUUM wavelenght (easy vac-->air)
- But original files from ground in AIR

Conversion not precise but approximate formulae

www.as.utexas.edu/~hebe/apogee/docs/air\_vacuum.pdf

- No metadata to express the nature (air/vacuum)
- No param to query by this (is vacuum?)
- Air/vacuum lambda who has to convert it

- client or DL service ? - massive scripting

- The precision on echelle spectra + I2 cell requires
  - Knowledge of dispersion (nonlinear)
  - WCS III support not common arXiv:astro-ph/0507293

# **SSA + SDM issues**

- How to handle two BAND ?
  - Query by BAND=I1/I2 in SSA to SELECT
  - SODA BAND = minl1 "space" minl2 to CUT
- What if data not contained in cutout request
  - Error or auto convert the SSA query and redo?
    - User wants to get only given region to compare
- Continuum normalization
  - Query by (virtual generation ?)
  - Process (on the fly, preselect files ?)
  - In theoretical spectra (TSAP) another collumn

# Conclusions

- The community needs VO way for precise spectroscopy task for CSP !
- Current SSA + SDM not fulfills the requirements of precision on subpixel level (bintable insufficient – dispersion interpolation)
- SSA seems to be confusing with SODA
- It is time to replace SSA but many spectra HARDLY in VO still !!!!

Should TAP + SODA work better ?