

# ISMDB

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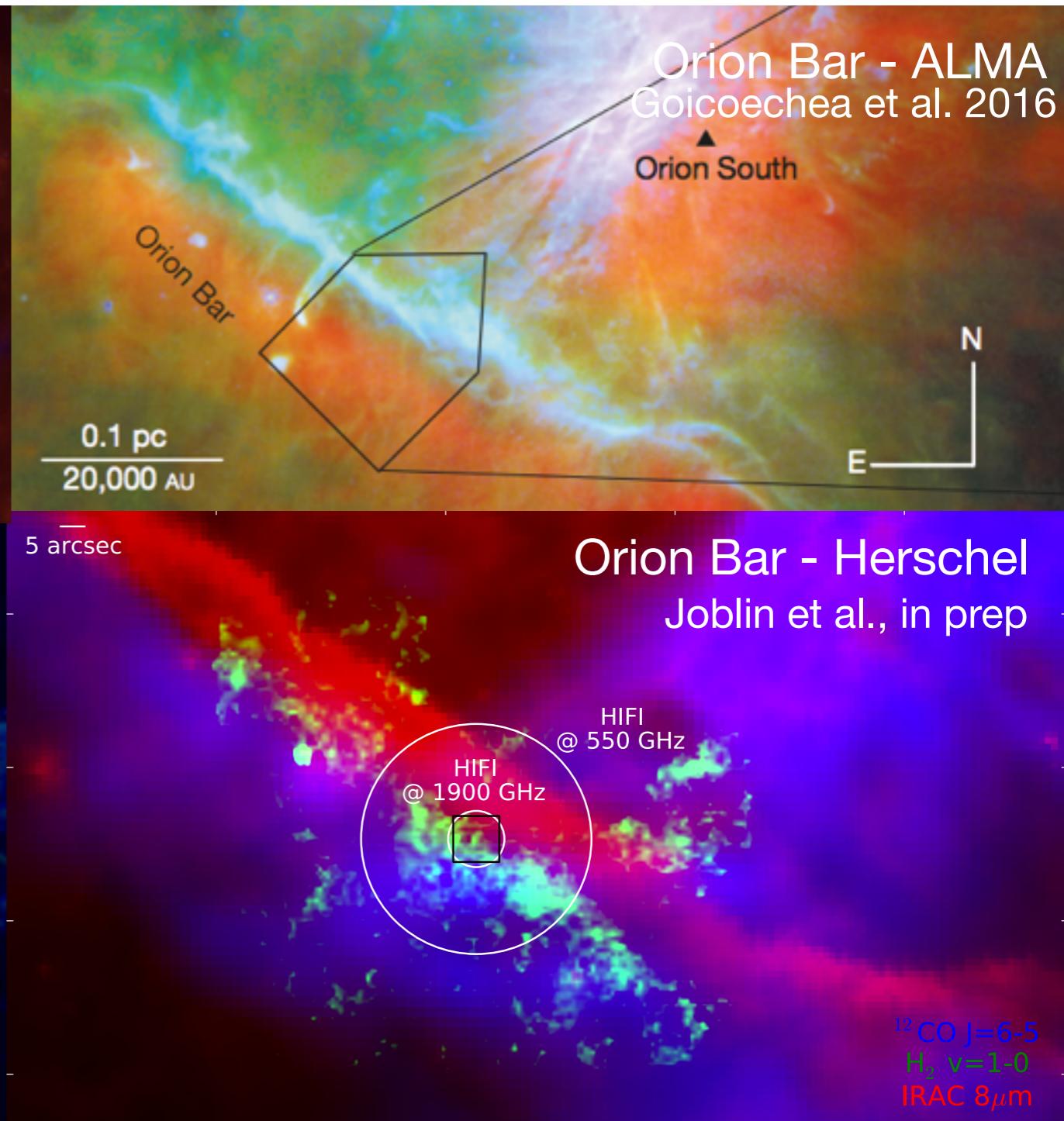
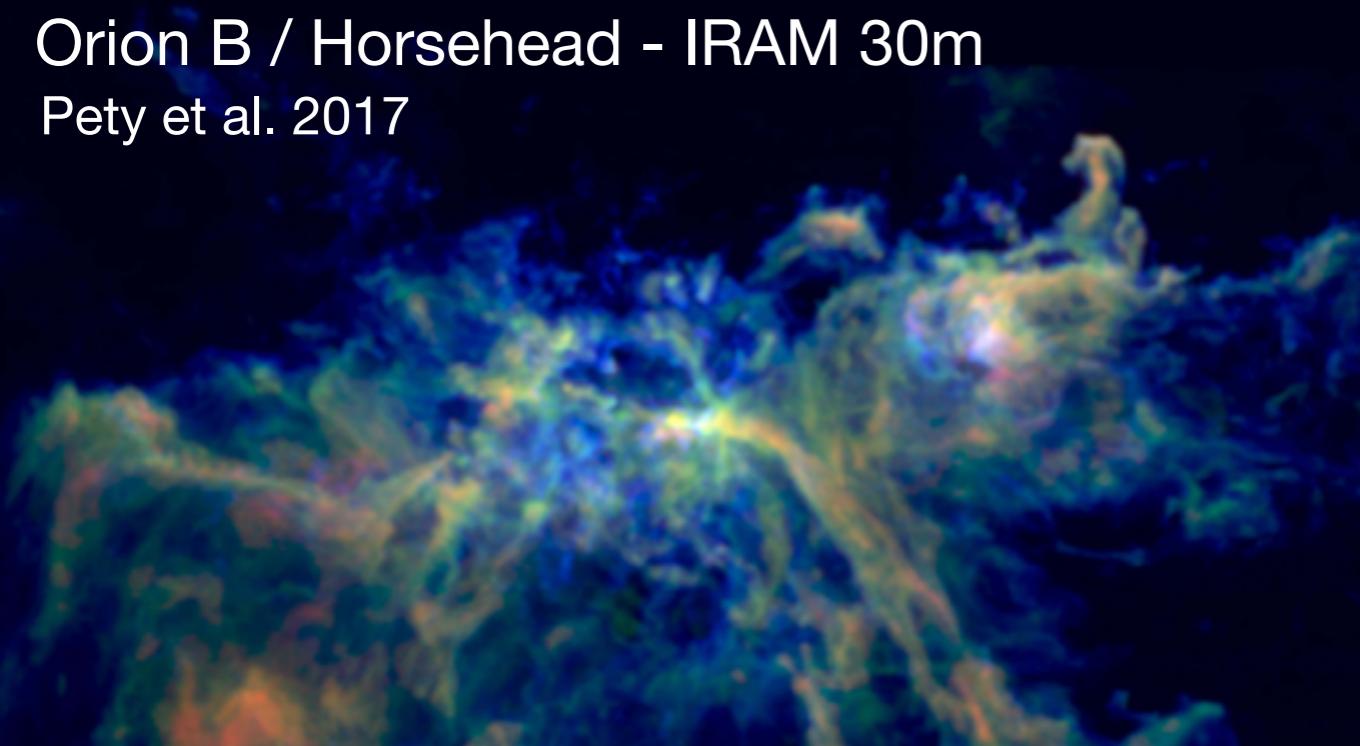
# InterStellar Medium DataBase

<http://ism.obspm.fr>

Franck Le Petit  
David Languignon  
Emeric Bron  
Nicolas Moreau

# ISMDB: Interstellar Medium DataBase

Goal: Provide numerical tools to interpret observations in the interstellar medium.

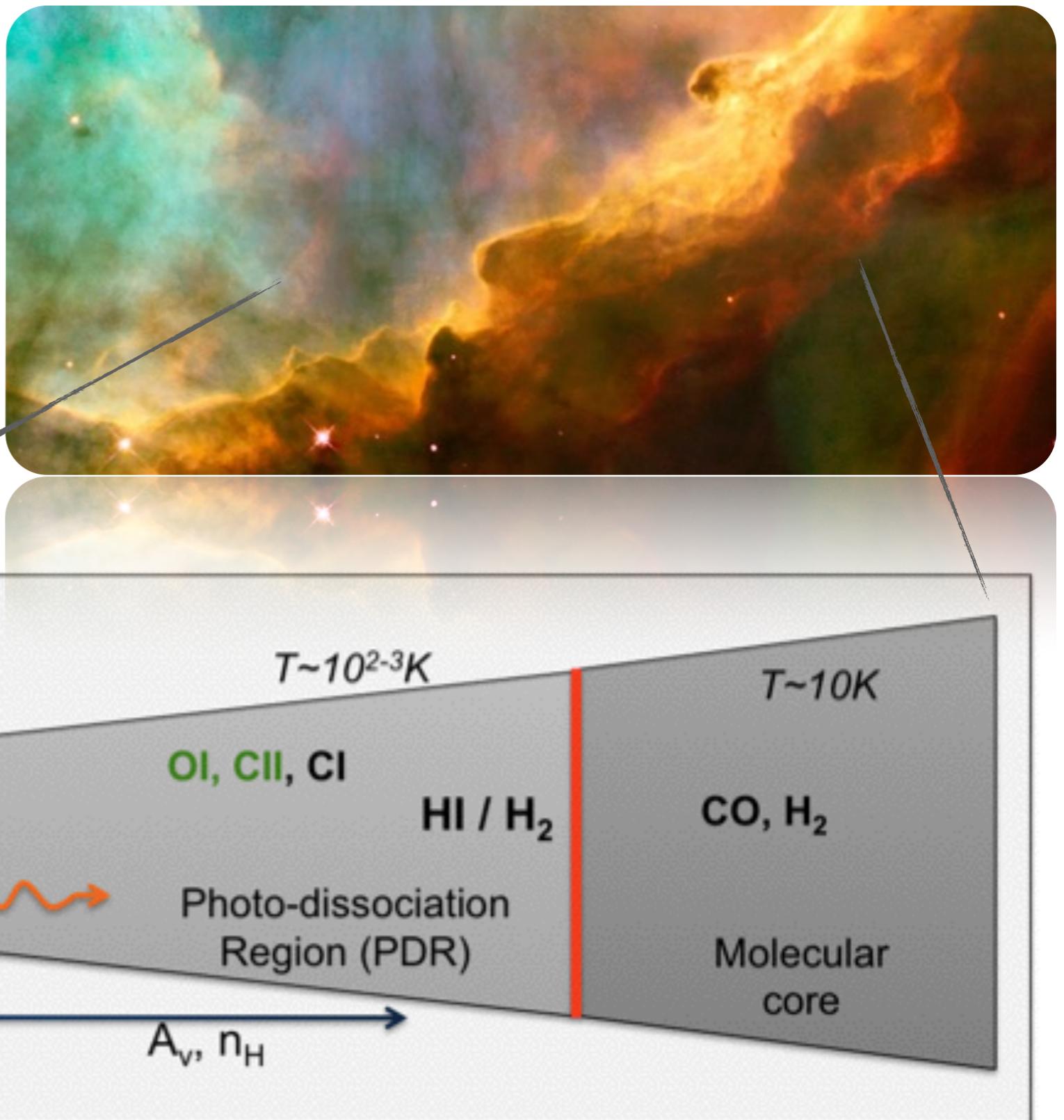


With the new instruments, observations of the ISM are:

- more and more detailed
- more and more difficult to interpret

# Schema of the transition atoms / molecules

PDR: Photo-dissociation region  
Transitions between H / H<sub>2</sub> regions

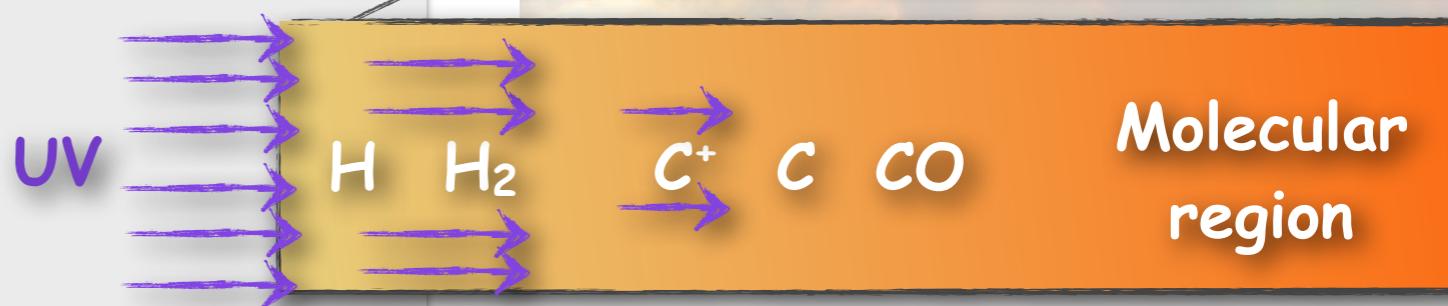


# Simulation code: Meudon PDR code

Le Petit et al. (2006), Gonzalez-Garcia et al. (2008), Le Petit et al. (2009), Le Bourlot et al. (2012)

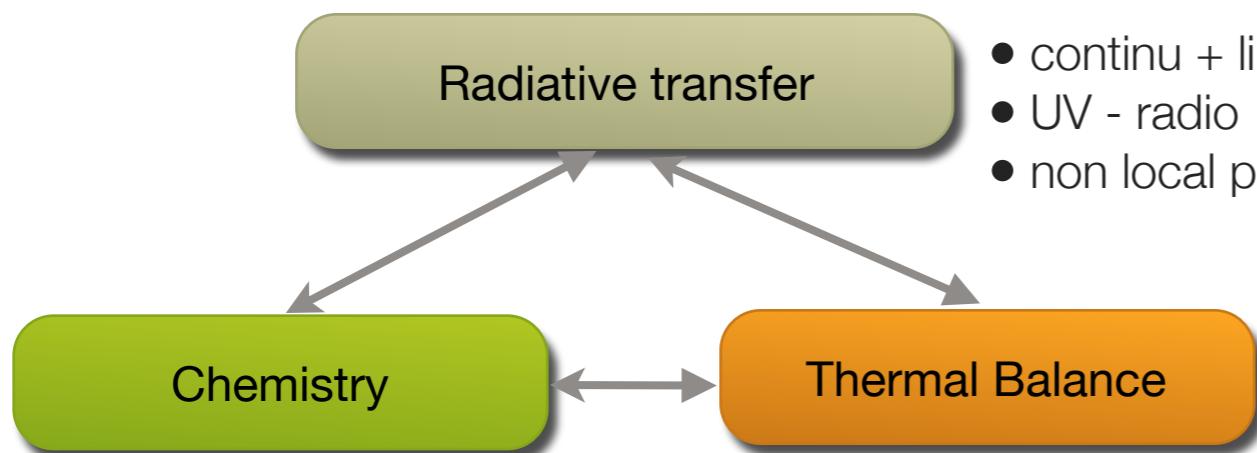
**PDR Code :** computation of the atomic and molecular structure of clouds and analysis of physical processes

- abundances
- excitation states
- temperatures (gas & grains)
- Intensities ( $H_2$ , CO,  $H_2O$ , ...)
- Column densities



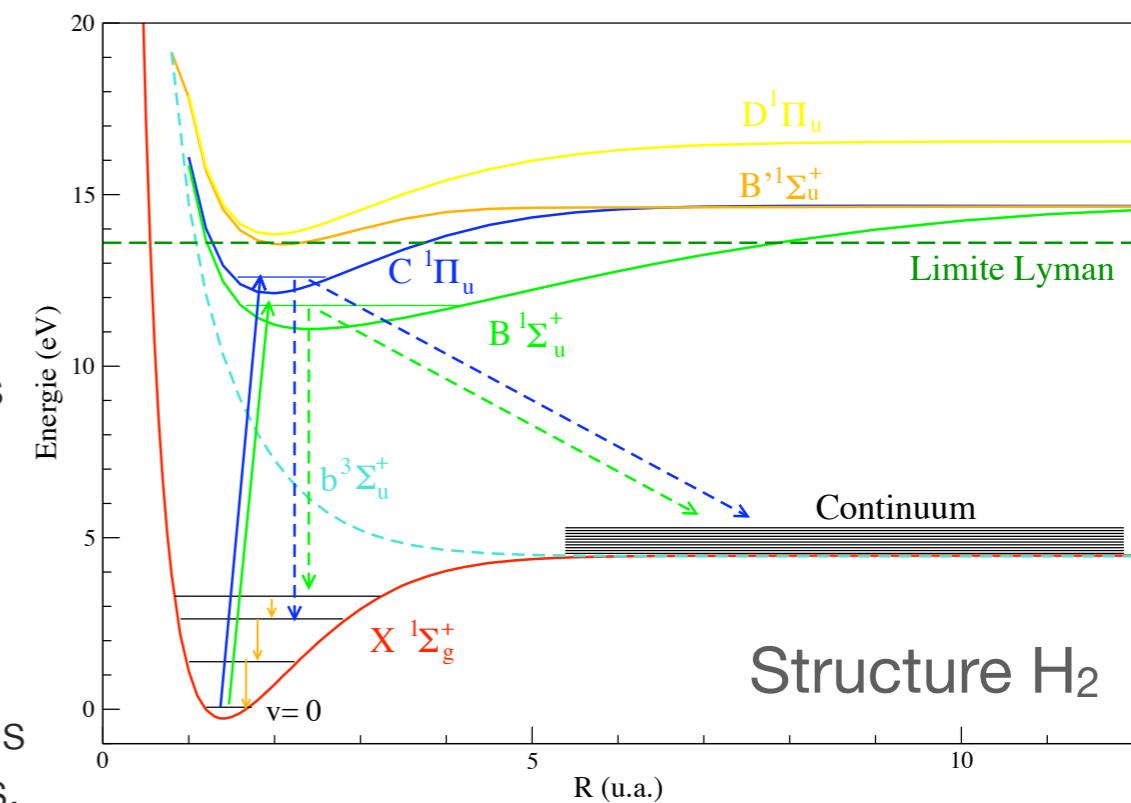
Interpretation of large instruments observations

ALMA, SOFIA, IRAM, HERSCHEL, Spitzer, VLT, HST, FUSE, ...



- gas & grains chemistry
- several hundred species
- thousands reactions

- Non LTE Detailed balance in levels
- Photo-electric effect, cosmic rays, chemistry, ...



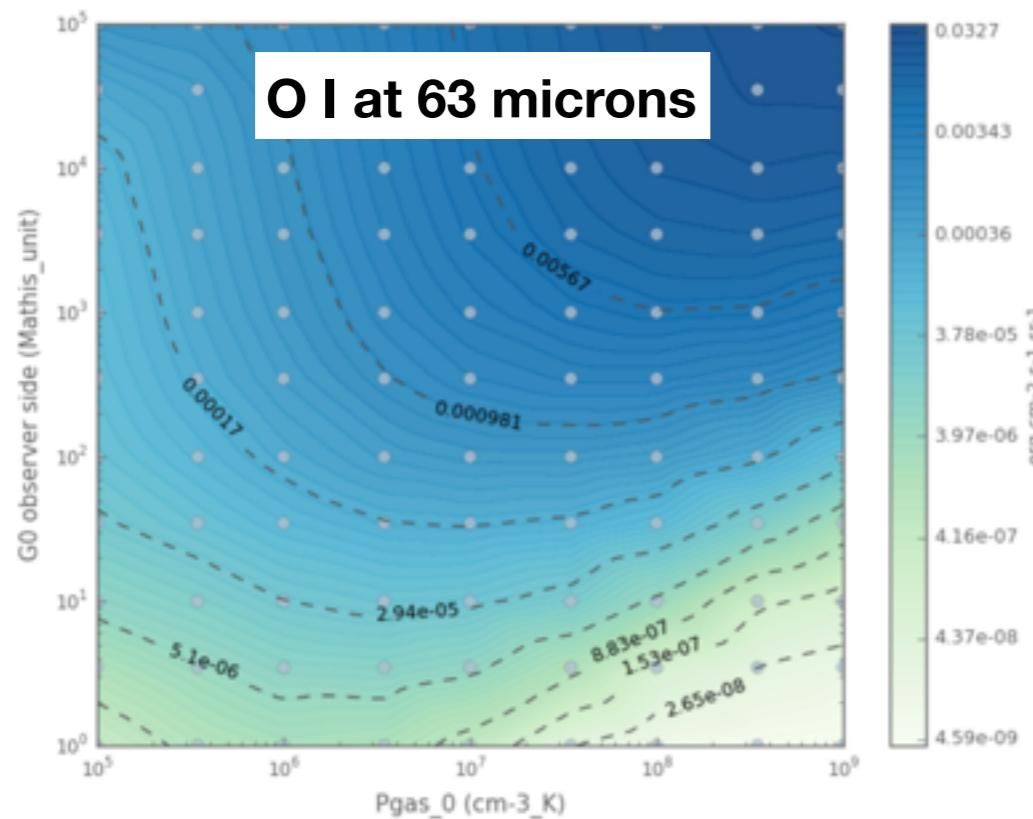
# Interpretation of line intensities

Comparison of observations & grids of models

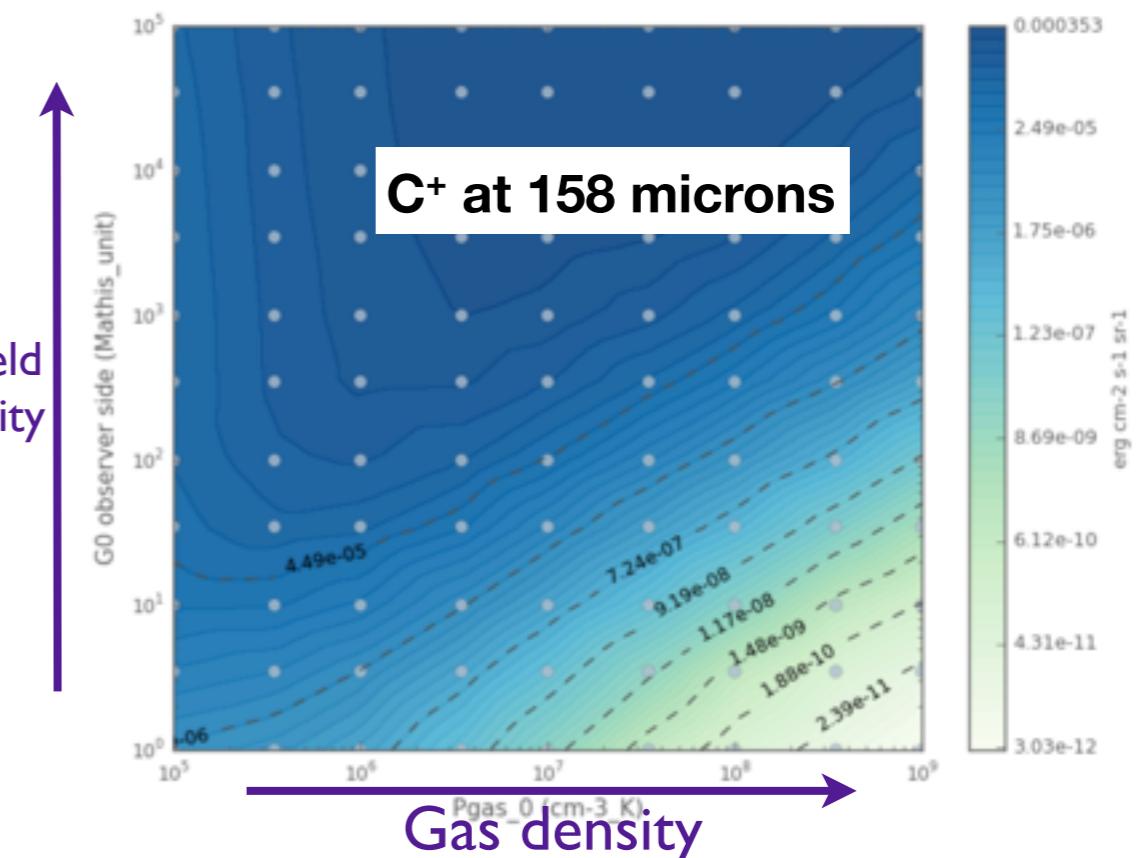
Determine physical conditions:

- density / pressure of clouds
- size of clouds
- intensity of radiation field : UV and X-rays
- metallicity
- ...

Grids of models:



UV field  
intensity



Gas density

1 model : 6 hours to several days CPU time

Interpretation requires up to thousands models

→ weeks or months of work

How to reduce this work to a few minutes ?

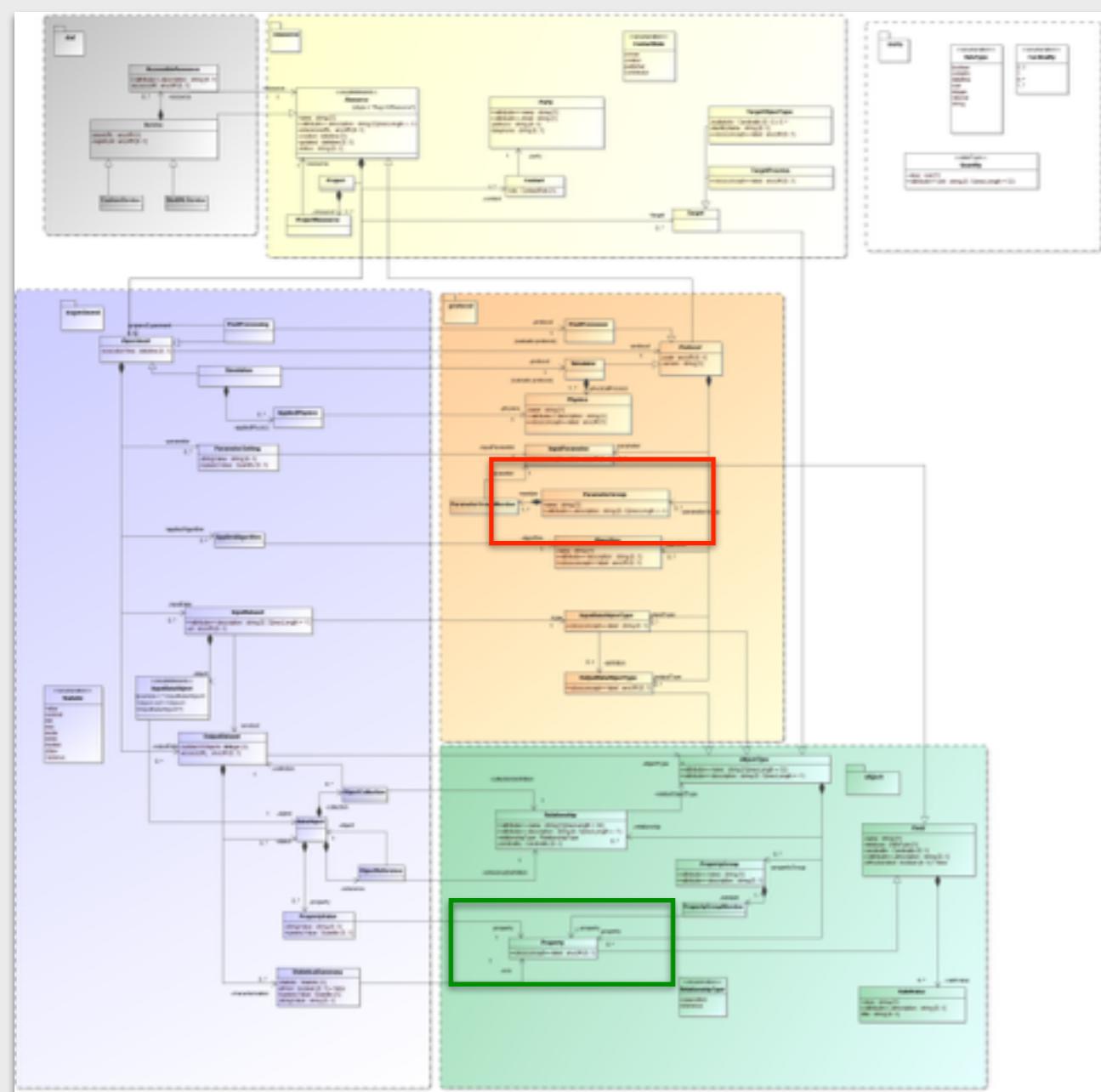


Production of grids of PDR models

- ~ 3000 PDR models in ISMDB for **standard galactic conditions**
- ~ 10 To of data

Cover: **PDR models for Herschel, IRAM/Noema, JWST, ALMA ... observations**

Characterisation of PDR models with the **Simulation DataModel** :



For each model:

### **InputParameters**

- pressure, UV field, ...
- ~ 20 quantities

### **Properties / statistics**

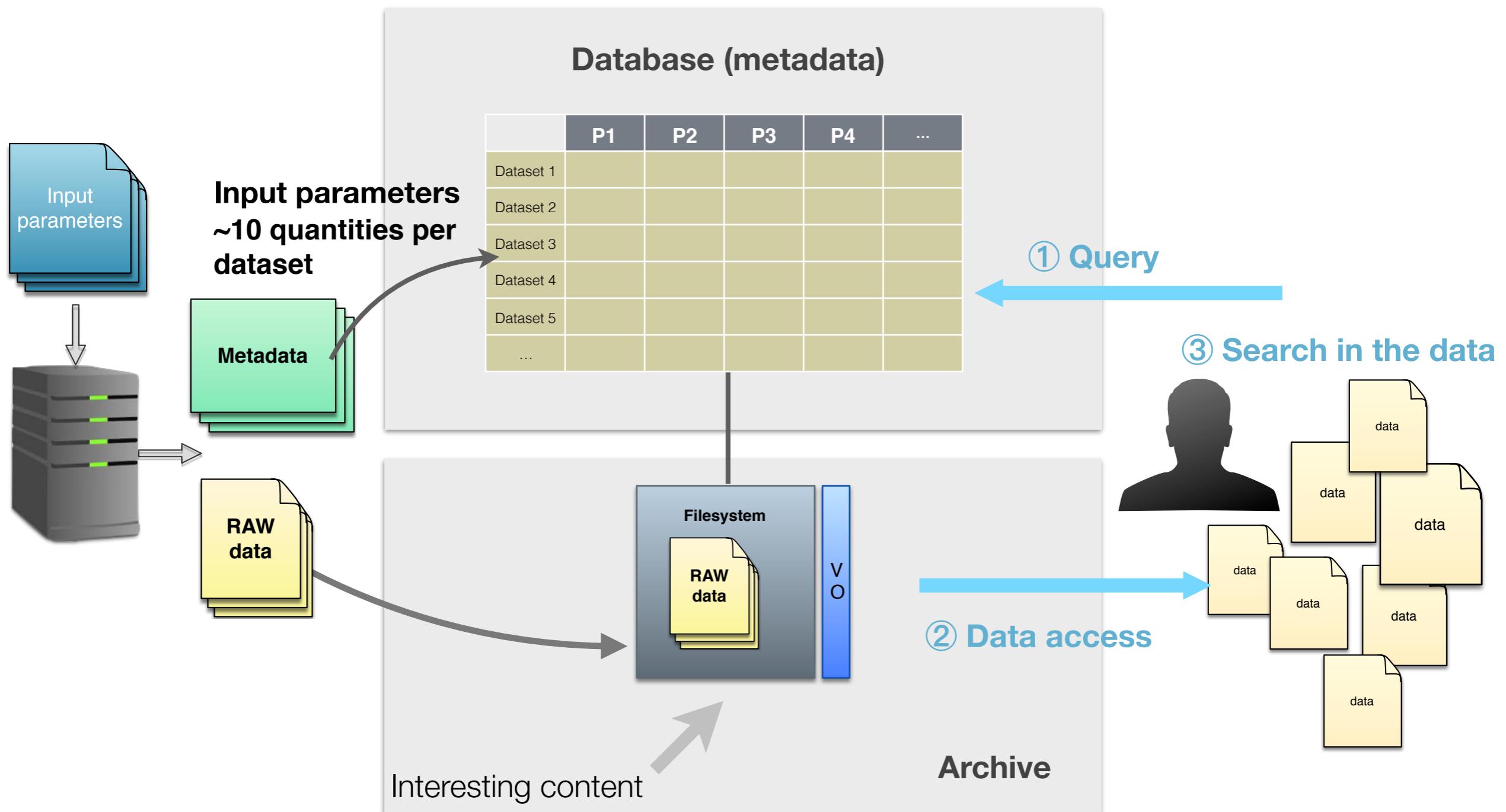
- line intensities: C+, C, H<sub>2</sub>, CO, H<sub>2</sub>O, ...
- +150 000 metadata

# ISMDB

## ISMDB: InterStellar Medium DataBase

- not only a classical database to find pre-computed models
- but **also a tool that can *interpret* observations**

Standard databases:

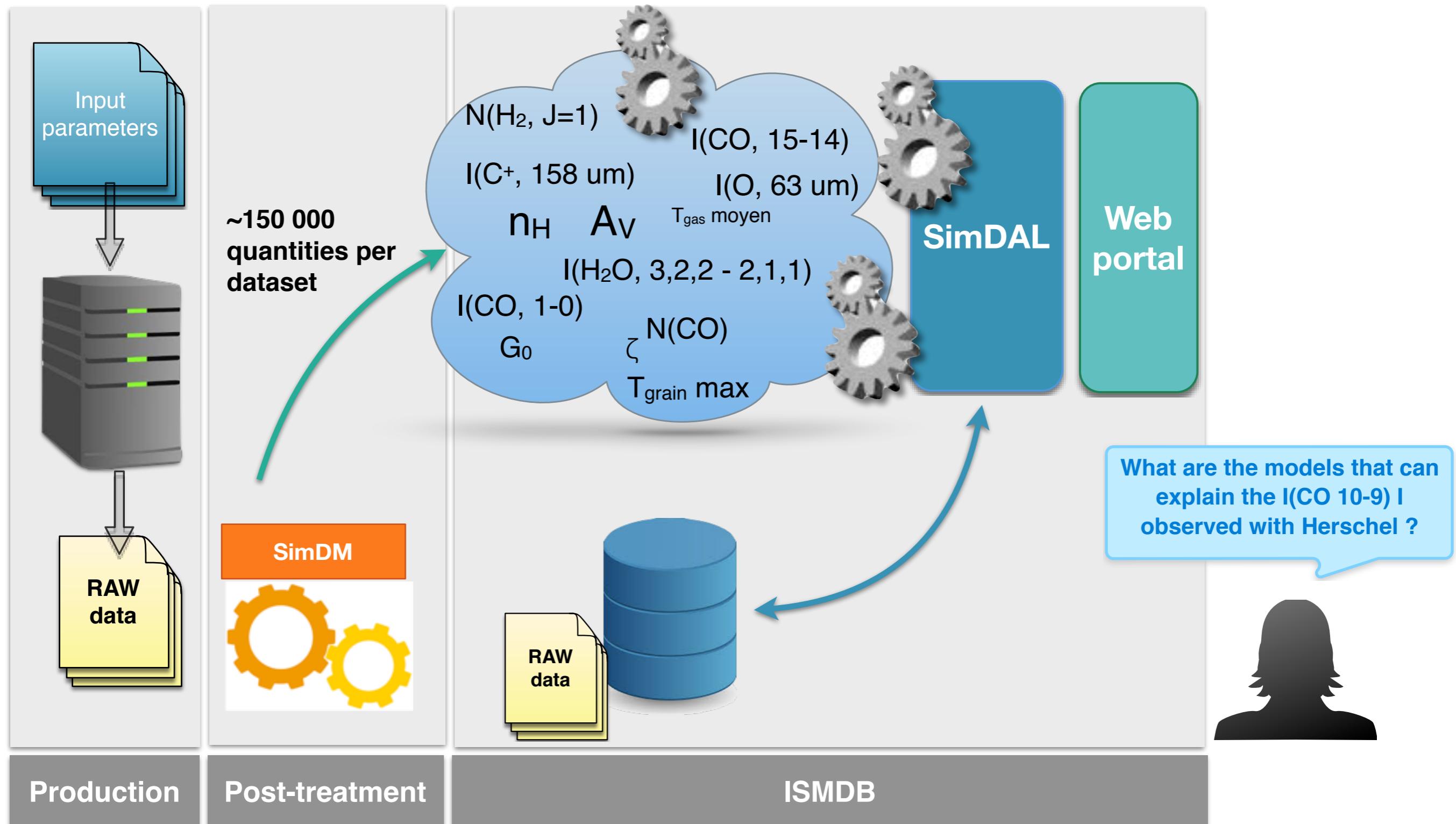


# ISMDB

- not only a classical database to find pre-computed models
- but also a tool that can *interpret* observations

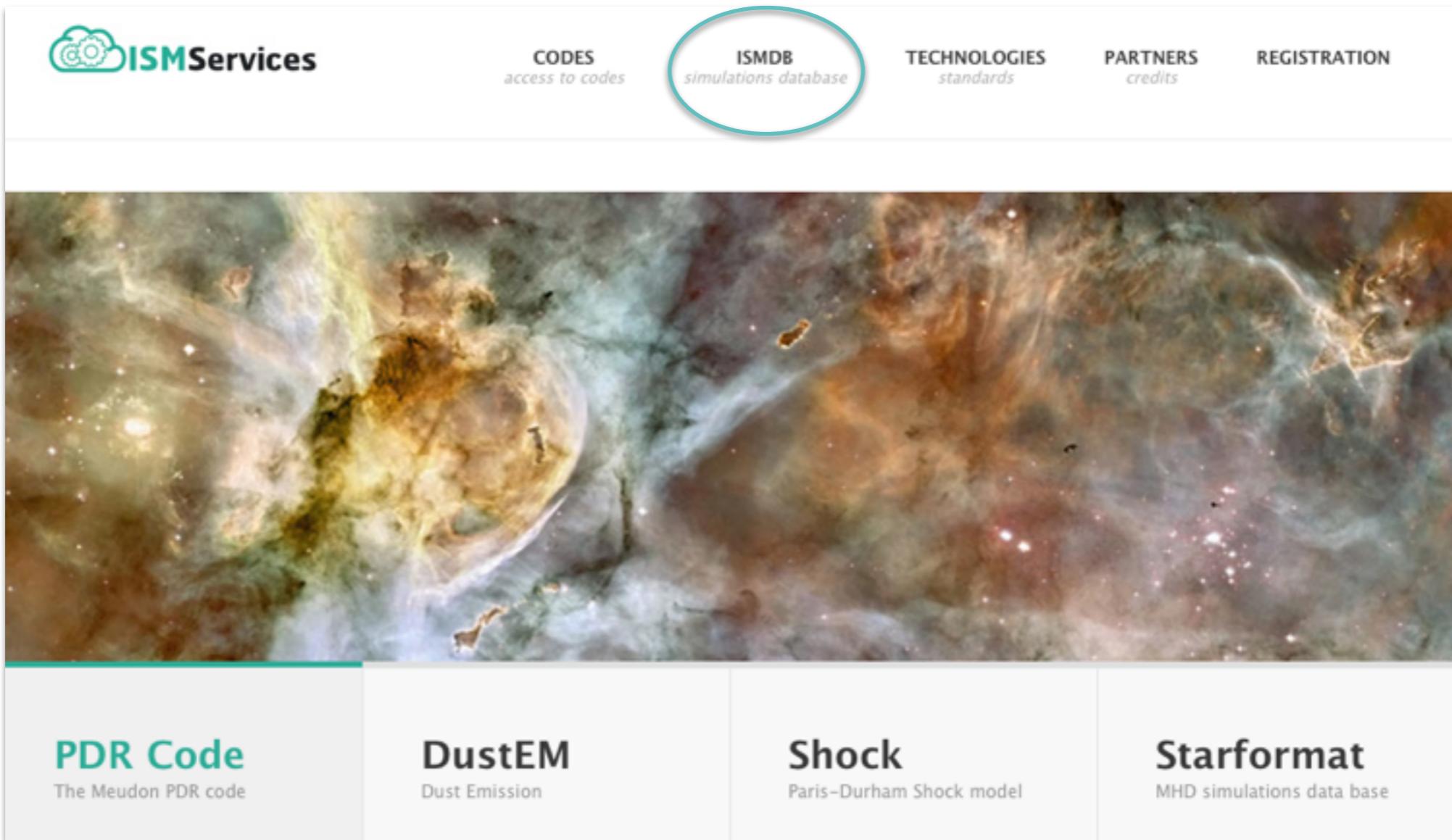
**High dimensionnality**

=> specific technologies required



# ISMDB

<http://ism.obspm.fr>



The screenshot shows the homepage of the ISM Services website. At the top, there is a navigation bar with links: 'ISM Services' (with a cloud icon), 'CODES access to codes', 'ISMDB simulations database' (which is circled in green), 'TECHNOLOGIES standards', 'PARTNERS credits', and 'REGISTRATION'. Below the navigation bar is a large, colorful image of a nebula or galaxy. At the bottom of the page, there are four boxes with links: 'PDR Code' (The Meudon PDR code), 'DustEM' (Dust Emission), 'Shock' (Paris-Durham Shock model), and 'Starformat' (MHD simulations data base).

**ISM Services**

CODES  
access to codes

ISMDB  
simulations database

TECHNOLOGIES  
standards

PARTNERS  
credits

REGISTRATION

**PDR Code**  
The Meudon PDR code

**DustEM**  
Dust Emission

**Shock**  
Paris-Durham Shock model

**Starformat**  
MHD simulations data base

- Source codes & specific developments
- Online codes
- Tools to analyze results
  - Extractor & Chemistry Analyzer
- ISMDB

## ISM DataBase – Inverse Search service Beta

Grid of isobaric PDR 1.5.2 models

2016.12.03

### 1 – search among two parameters

x	Pgas_0	(cm <sup>-3</sup> _K)	<input checked="" type="checkbox"/> log scale
y	G0 observer side	(Mathis_unit)	<input checked="" type="checkbox"/> log scale

### 2 – fix all the other parameters

AVmax	(mag)	10
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### 3 – observational constraints

Search for available quantities... Ex: N(H)	<a href="#">Use</a>
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"I(CO v=0J=1->v=0J=0 angle 00 deg)" > 1.8E-7  
"I(CO v=0J=1->v=0J=0 angle 00 deg)" < 2.4E-7  
"I(H2 v=0J=2->v=0J=0 angle 60 deg)" > 1E-8  
"I(H2 v=0J=2->v=0J=0 angle 60 deg)" < 5E-7

[Search](#)

### ① Select the searched input parameters

Example of a search:

- gas pressure
- UV intensity

### ② Fix the other input parameters

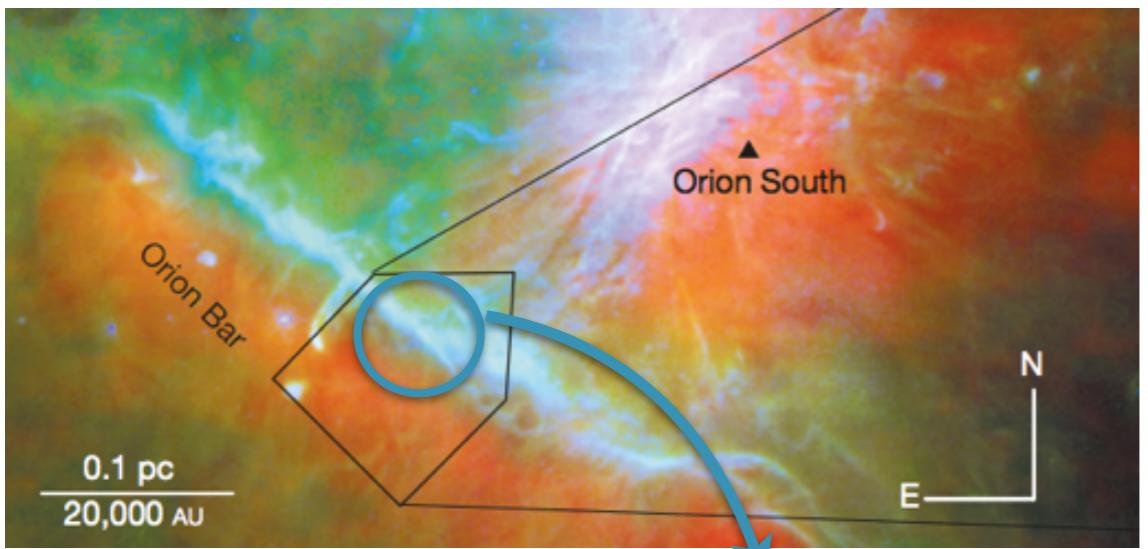
Example: size of the cloud

### ③ Enter the observations

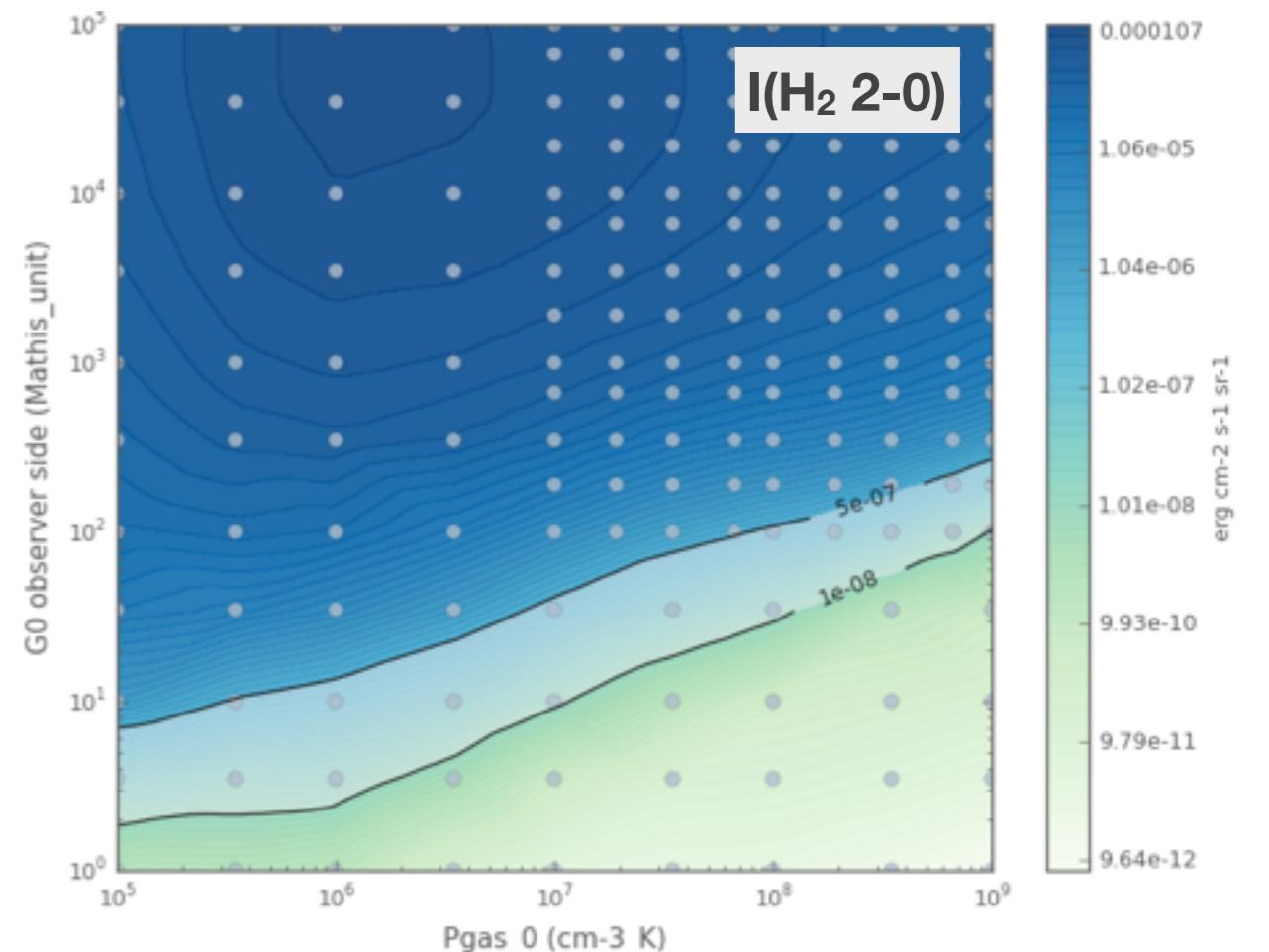
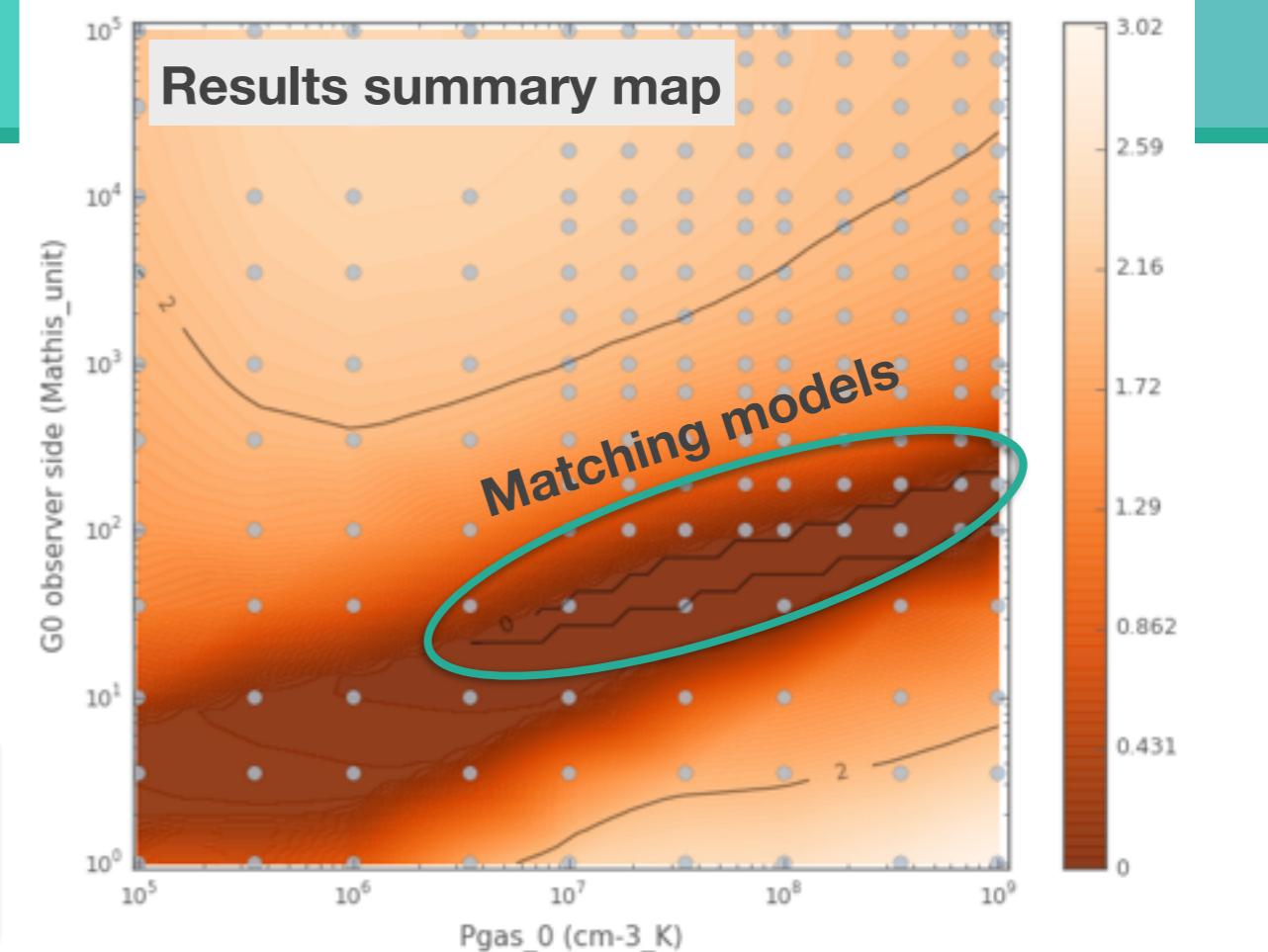
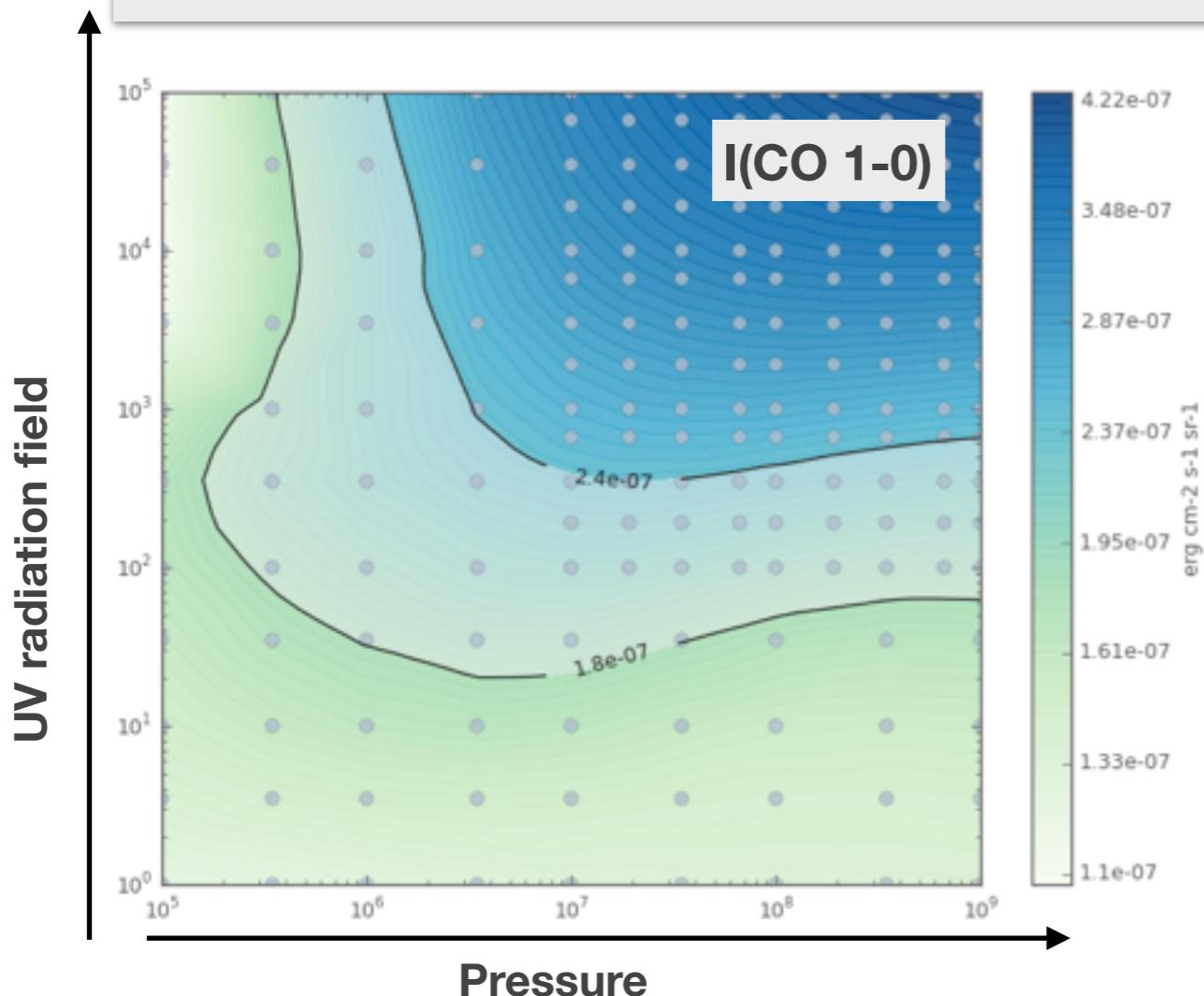
Example: observations CO and H<sub>2</sub> intensities

$1.8 \cdot 10^{-7} < I(\text{CO } 1\text{-}0) < 2.4 \cdot 10^{-7}$  erg cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>  
 $1.0 \cdot 10^{-8} < I(\text{H}_2 \text{ } 2\text{-}0) < 5.0 \cdot 10^{-7}$  erg cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>

# Interpretation of observations



$1.8 \cdot 10^{-7} < I(\text{CO } 1-0) < 2.4 \cdot 10^{-7} \text{ erg cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$   
 $1.0 \cdot 10^{-8} < I(\text{H}_2 \text{ 2-0}) < 5.0 \cdot 10^{-7} \text{ erg cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$



# VO Integration

Metadata definition & organization

Data Access

Raw data (extractor tool)

Extractor Tool

PDR Extractor

n(CO) Confirm Remove All

DM54NoPAH\_A1e1p6p7e7r1e4\_s\_20.hdf5

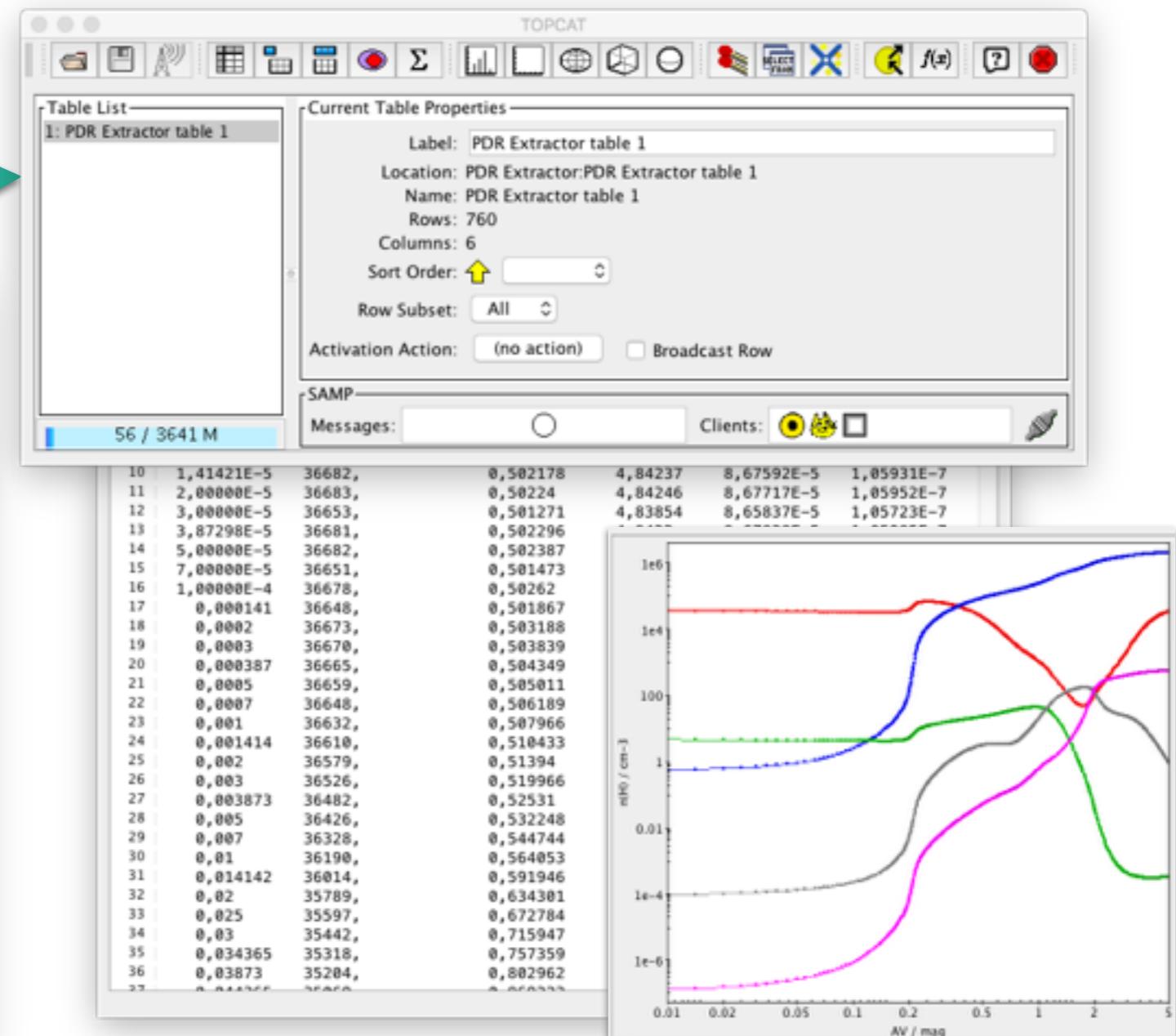
- Integrated quantities
- Local quantities
  - Auxiliary
    - Excitation
    - H<sub>2</sub> chemistry
    - Ice layers
    - Molecular fraction
    - Photo reactions
    - Radiation
    - Thermal balance
  - Densities
    - Column densities
    - Densities
  - Dust
  - Gas state
  - Positions
    - AV
    - Distance
    - tauV
  - Parameters

Export as Text Export as VOTable Send Table

Simulation Data Model (SimDM)

Simulation Data Access Layer (SimDAL)

VOTable  
SAMP connector



# Interpretation of observations

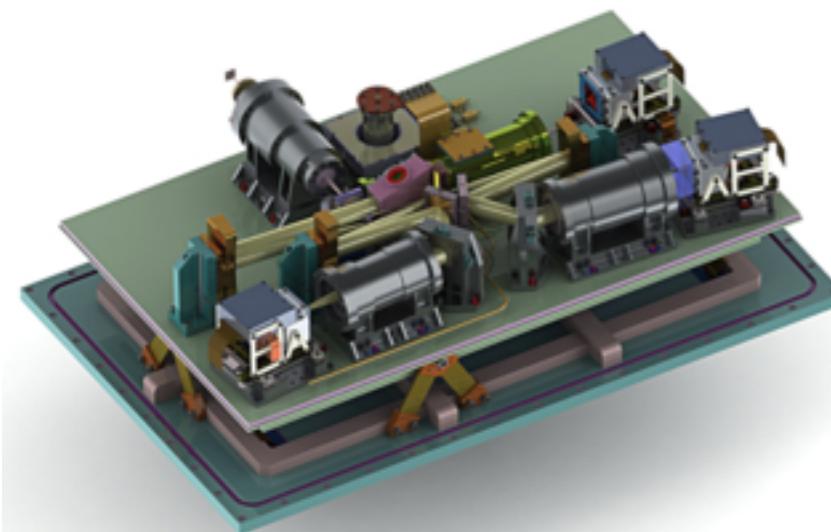
## IGRINS observations

Instrument - Univ. Texas, Austin

Mc Donald observatory

Bands: H et K (1.5 to 2.5 microns)

R = 45 000

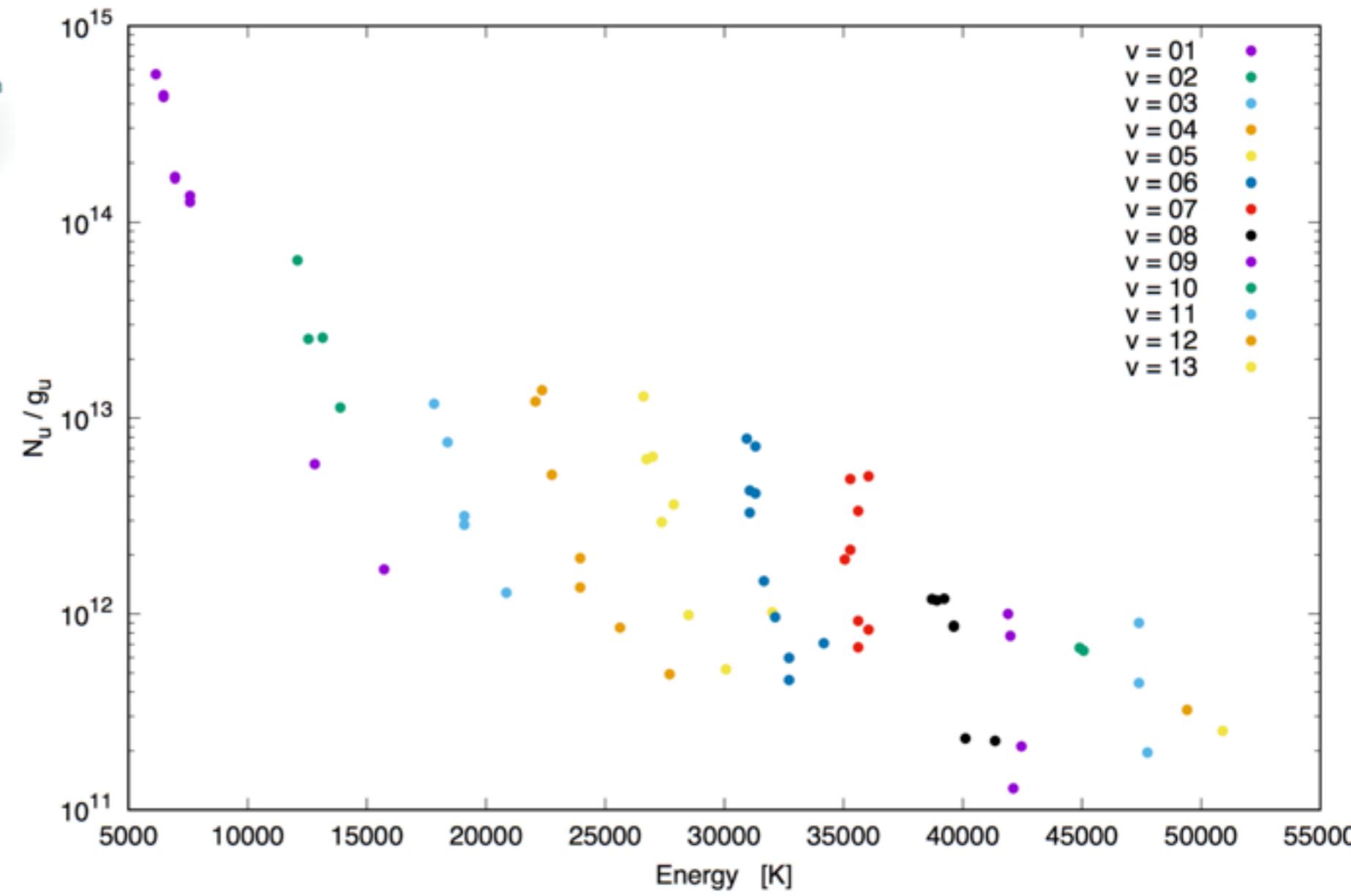


## Observations of NGC 7023

(Le et al. - 2016 / ArXiV)

- Detection of **70 H<sub>2</sub> lines in NGC 7023**
- Conclude to a clumpy medium

H<sub>2</sub> excitation diagram at position A

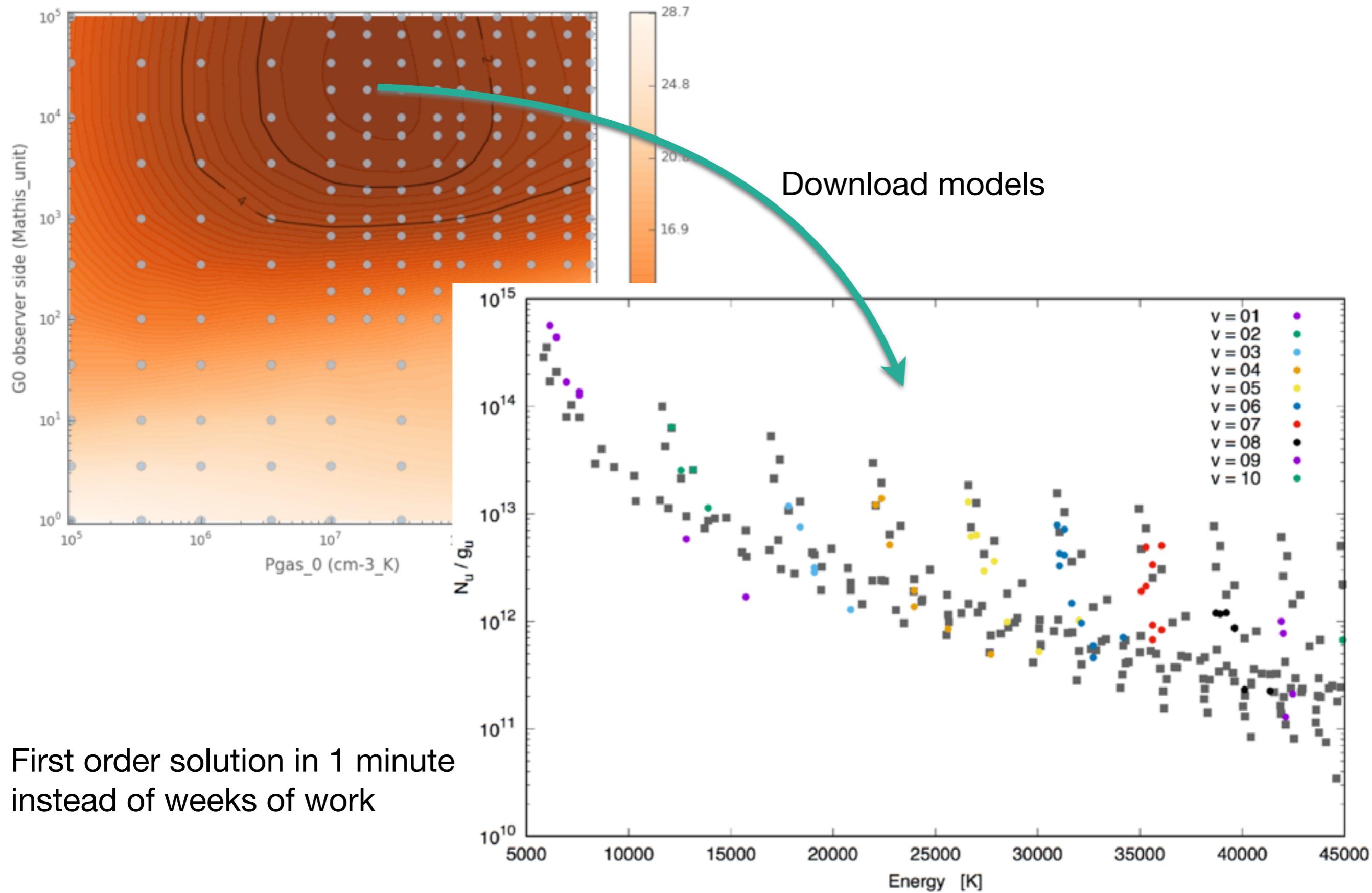


# Interpretation of observations

Build the query for the 70 H<sub>2</sub> lines (140 constraints)

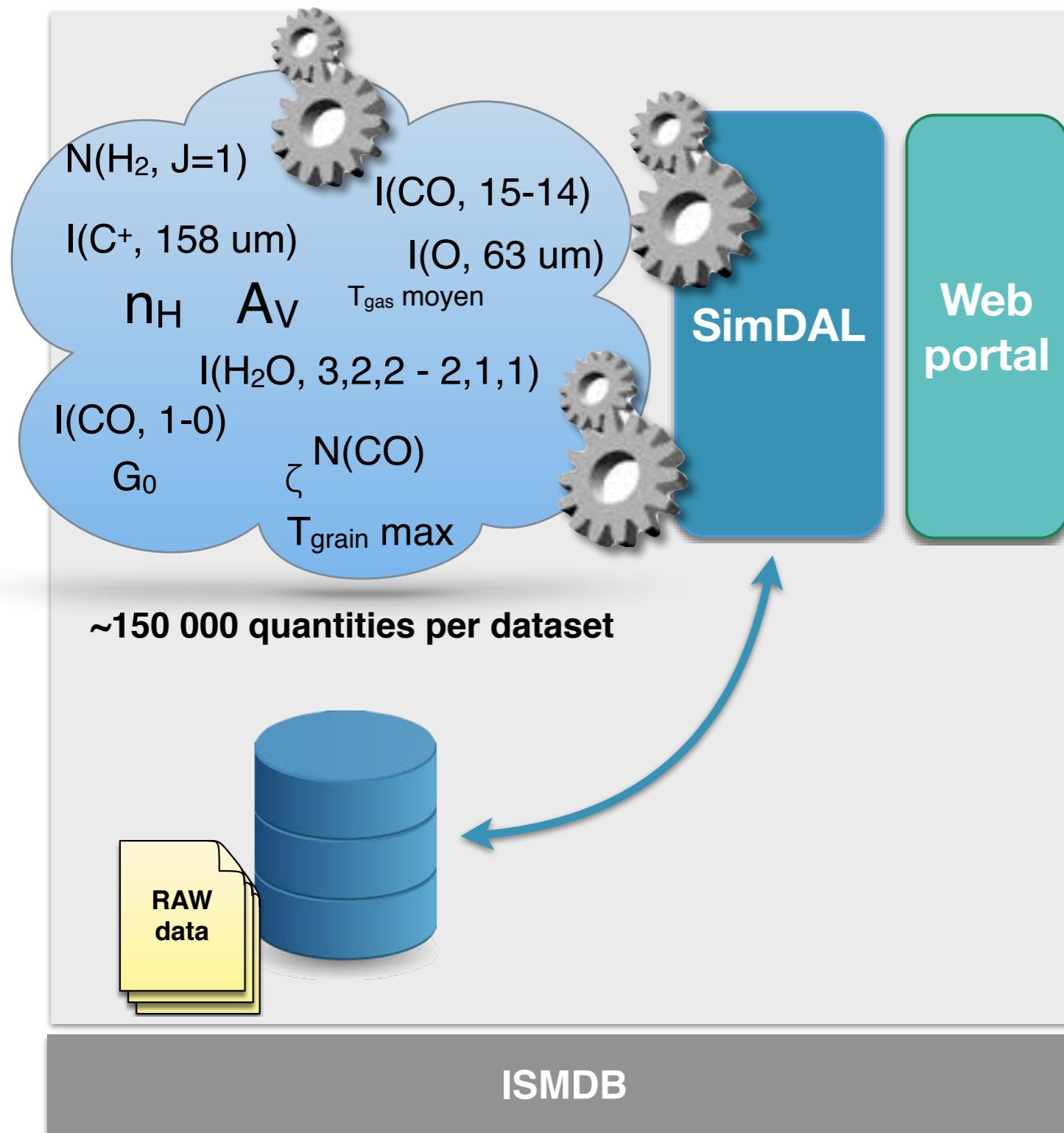
```
"I(H2 v=4,J=9->v=2,J=9 angle 60 deg)" < 5.408e-06
"I(H2 v=4,J=9->v=2,J=9 angle 60 deg)" > 2.912e-06
"I(H2 v=4,J=1->v=2,J=3 angle 60 deg)" < 2.665e-05
"I(H2 v=4,J=1->v=2,J=3 angle 60 deg)" > 1.435e-05
"I(H2 v=5,J=4->v=3,J=4 angle 60 deg)" < 9.659e-06
"I(H2 v=5,J=4->v=3,J=4 angle 60 deg)" > 5.201e-06
"I(H2 v=3,J=3->v=1,J=5 angle 60 deg)" < 1.2532e-05
"I(H2 v=3,J=3->v=1,J=5 angle 60 deg)" > 6.748e-06
"I(H2 v=5,J=5->v=3,J=5 angle 60 deg)" < 9.607e-06
"I(H2 v=5,J=5->v=3,J=5 angle 60 deg)" > 5.173e-06
"I(H2 v=6,J=2->v=4,J=0 angle 60 deg)" < 7.67e-06
"I(H2 v=6,J=2->v=4,J=0 angle 60 deg)" > 4.13e-06
"I(H2 v=10,J=1->v=7,J=3 angle 60 deg)" < 4.121e-06
"I(H2 v=10,J=1->v=7,J=3 angle 60 deg)" > 2.219e-06
"I(H2 v=5,J=0->v=3,J=2 angle 60 deg)" < 1.118e-05
"I(H2 v=5,J=0->v=3,J=2 angle 60 deg)" > 6.02e-06
"I(H2 v=5,J=7->v=3,J=7 angle 60 deg)" < 6.955e-06
"I(H2 v=5,J=7->v=3,J=7 angle 60 deg)" > 3.745e-06
"I(H2 v=4,J=2->v=2,J=4 angle 60 deg)" < 1.1531e-05
"I(H2 v=4,J=2->v=2,J=4 angle 60 deg)" > 6.209e-06
"I(H2 v=7,J=4->v=5,J=2 angle 60 deg)" < 5.109e-06
"I(H2 v=7,J=4->v=5,J=2 angle 60 deg)" > 2.751e-06
"I(H2 v=6,J=1->v=4,J=1 angle 60 deg)" < 1.846e-05
"I(H2 v=6,J=1->v=4,J=1 angle 60 deg)" > 9.94e-06
"I(H2 v=6,J=2->v=4,J=2 angle 60 deg)" < 1.599e-05
"I(H2 v=6,J=2->v=4,J=2 angle 60 deg)" > 8.61e-06
"I(H2 v=5,J=9->v=3,J=9 angle 60 deg)" < 1.729e-05
"I(H2 v=5,J=9->v=3,J=9 angle 60 deg)" > 9.31e-06
"I(H2 v=5,J=1->v=3,J=3 angle 60 deg)" < 2.379e-05
"I(H2 v=5,J=1->v=3,J=3 angle 60 deg)" > 1.281e-05
"I(H2 v=13,J=1->v=9,J=1 angle 60 deg)" < 9.334e-07
"I(H2 v=13,J=1->v=9,J=1 angle 60 deg)" > 5.026e-07
"I(H2 v=6,J=3->v=4,J=3 angle 60 deg)" < 1.287e-05
"I(H2 v=6,J=3->v=4,J=3 angle 60 deg)" > 6.93e-06
I(H2 v=7,J=3->v=5,J=1 angle 60 deg)" < 1.1531e-05
I(H2 v=7,J=3->v=5,J=1 angle 60 deg)" > 6.209e-06
I(H2 v=4,J=3->v=2,J=5 angle 60 deg)" < 1.2961e-05
I(H2 v=4,J=3->v=2,J=5 angle 60 deg)" > 6.979e-06
I(H2 v=6,J=4->v=4,J=4 angle 60 deg)" < 3.523e-06
I(H2 v=6,J=4->v=4,J=4 angle 60 deg)" > 1.897e-06
I(H2 v=6,J=5->v=4,J=5 angle 60 deg)" < 7.878e-06
I(H2 v=6,J=5->v=4,J=5 angle 60 deg)" > 4.242e-06
I(H2 v=3,J=5->v=1,J=7 angle 60 deg)" < 2.457e-06
I(H2 v=3,J=5->v=1,J=7 angle 60 deg)" > 1.323e-06
I(H2 v=11,J=1->v=8,J=1 angle 60 deg)" < 2.899e-06
I(H2 v=11,J=1->v=8,J=1 angle 60 deg)" > 1.561e-06
I(H2 v=7,J=2->v=5,J=0 angle 60 deg)" < 4.628e-06
I(H2 v=7,J=2->v=5,J=0 angle 60 deg)" > 2.492e-06
I(H2 v=8,J=7->v=6,J=5 angle 60 deg)" < 5.824e-06
I(H2 v=8,J=7->v=6,J=5 angle 60 deg)" > 3.136e-06
I(H2 v=5,J=2->v=3,J=4 angle 60 deg)" < 9.425e-06
I(H2 v=5,J=2->v=3,J=4 angle 60 deg)" > 5.075e-06
I(H2 v=6,J=0->v=4,J=2 angle 60 deg)" < 9.776e-06
I(H2 v=6,J=0->v=4,J=2 angle 60 deg)" > 5.264e-06
I(H2 v=6,J=7->v=4,J=7 angle 60 deg)" < 1.2532e-05
I(H2 v=6,J=7->v=4,J=7 angle 60 deg)" > 6.748e-06
I(H2 v=11,J=3->v=8,J=3 angle 60 deg)" < 1.924e-06
I(H2 v=11,J=3->v=8,J=3 angle 60 deg)" > 1.036e-06
I(H2 v=1,J=11->v=0,J=9 angle 60 deg)" < 4.407e-06
I(H2 v=1,J=11->v=0,J=9 angle 60 deg)" > 2.373e-06
I(H2 v=8,J=5->v=6,J=3 angle 60 deg)" < 5.122e-06
I(H2 v=8,J=5->v=6,J=3 angle 60 deg)" > 2.758e-06
I(H2 v=7,J=1->v=5,J=1 angle 60 deg)" < 1.2922e-05
I(H2 v=7,J=1->v=5,J=1 angle 60 deg)" > 6.958e-06
I(H2 v=8,J=4->v=6,J=2 angle 60 deg)" < 5.109e-06
I(H2 v=8,J=4->v=6,J=2 angle 60 deg)" > 2.751e-06
I(H2 v=6,J=1->v=4,J=3 angle 60 deg)" < 2.405e-05
...
...
...
...
```

# Interpretation of observations



# Semantics

## Interaction between a human and the system



What are the models that can explain the  $I(CO\ 10-9)$  I observed with Herschel ?



How the user can know the name of the available quantities ?

# Semantics

ISM Services    CODES    ISMDB

## ISM DataBase – Inverse Search service Beta

Grid of isobaric PDR 1.5.2 models  
2016.12.03

**1 – search among two parameters**

x Pg<sub>as</sub>\_0 (cm<sup>-3</sup>\_K)  log scale

y G<sub>0</sub> observer side (Mathis\_unit)  log scale

**2 – fix all the other parameters**

AVmax (mag) 10

**3 – observational constraints**

Search for available quantities... Ex: N(H) Use

"I(CO v=0J=1->v=0J=0 angle 00 deg)" > 1.8E-7  
"I(CO v=0J=1->v=0J=0 angle 00 deg)" < 2.4E-7  
"I(H<sub>2</sub> v=0J=2->v=0J=0 angle 60 deg)" > 1E-8  
"I(H<sub>2</sub> v=0J=2->v=0J=0 angle 60 deg)" < 5E-7

Search

More than 150 000 queriable quantities in the database

- users do not know all the available list
- users do not know how these quantities are named

Example:

- H<sub>2</sub> line at 12 microns
- H<sub>2</sub> 0-0 S(2)
- H<sub>2</sub> v=0, J=4 -> v=0, J=2

### 3 – observational constraints

Search for available quantities... Ex: N(H) Use



Semantics interpreter

# Semantics : SKOS vocabularies

Each metadata is taggued by:

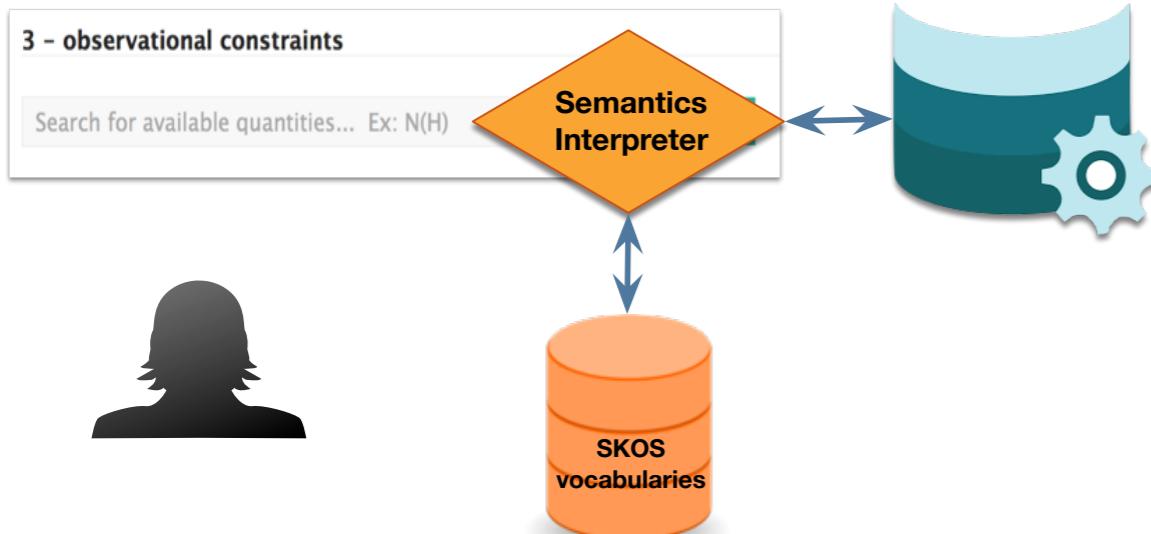
- ID
- name
- unit
- utype
- description
- **label (UCD / SKOS)**
- ...



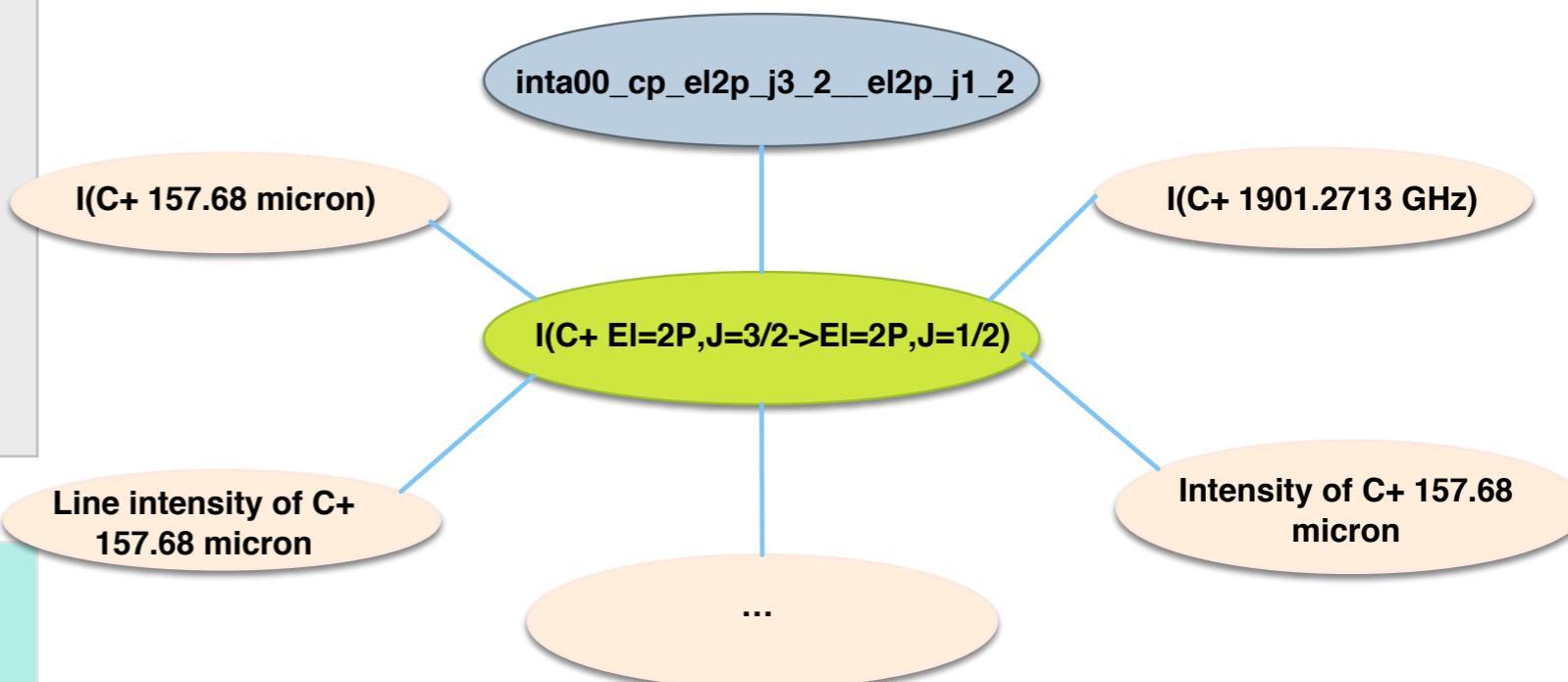
## SKOS vocabulary

For each quantity several synonyms  
(name, units, ...)

~ 300 000 terms for the PDR code



Example of the 157.7 micron C+ line intensity

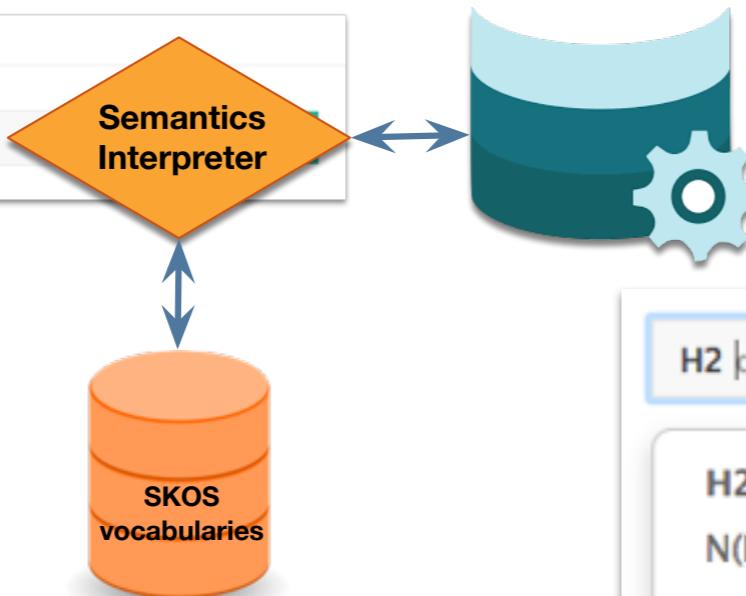


**ID:** inta00\_cp\_el2p\_j3\_2\_el2p\_j1\_2  
**PREF:** I(C+ EI=2P,J=3/2->EI=2P,J=1/2)  
**ALT:** I(C+ EI=2P,J=3/2->EI=2P,J=1/2) face on  
**ALT:** I(C+ 157.68 micron) face on  
**ALT:** Intensity of C+ 157.68 micron face on  
**ALT:** Line intensity of C+ 157.68 micron face on  
**ALT:** I(C+ 1901.2713 GHz) face on  
**ALT:** Intensity of C+ 1901.2713 GHz face on  
**ALT:** Line intensity of C+ 1901.2713 GHz face on  
...

# Semantics : SKOS vocabularies

3 - observational constraints

Search for available quantities... Ex: N(H)



## Semantics Interpreter

### Semantics

SKOS: PREF + ALT  
→ synonyms

+

### Ranking system

(learn from users)

H<sub>2</sub> column density

Use

H<sub>2</sub> column density

N(H<sub>2</sub>)

N(C<sub>2</sub>H<sub>2</sub>)

N(c-C<sub>3</sub>H<sub>2</sub>)

N(C<sub>\_13</sub>CH<sub>2</sub>)

N(C<sub>\_13</sub>CH<sub>2</sub> +)

C<sub>2</sub>H<sub>2</sub> column density

Column density of H<sub>2</sub>

straints.

I(H<sub>2</sub> | 0-0 S(0)) angle 00 degrees

Use

I(H<sub>2</sub> 0-0 S(0)) angle 00 degrees

I(H<sub>2</sub> 10-10 S(0)) angle 00 degrees

I(H<sub>2</sub> 9.6645 micrometres) angle 00 degrees

I(H<sub>2</sub> 28.2196 micrometres) angle 00 degrees

I(H<sub>2</sub> 156.4883 micrometres) angle 00 degrees

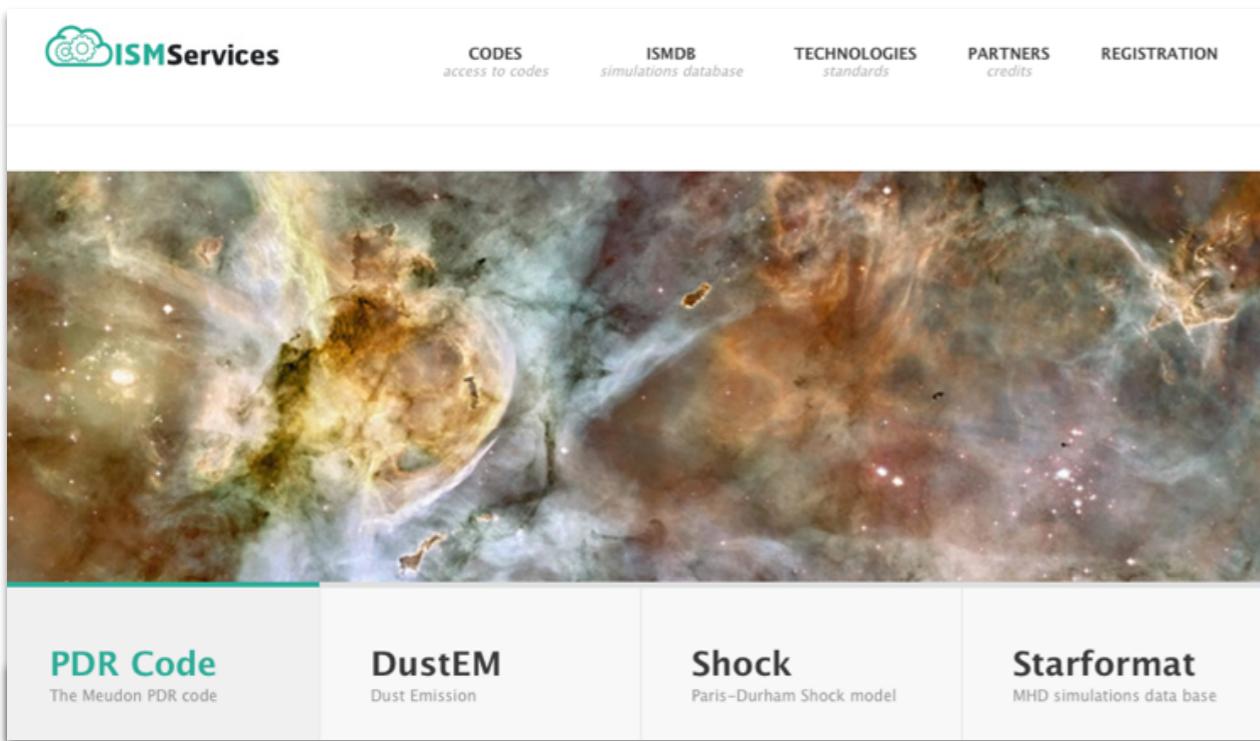
I(H<sub>2</sub> v=0,J=2->v=0,J=0) angle 00 degrees

I(H<sub>2</sub>O 6.1140 cm<sup>-1</sup>) angle 00 degrees

I(H<sub>2</sub>O J=1,ka=1,kc=1->J=0,ka=0,kc=0) angle 00 degrees

straints.

# ISMDB



The screenshot shows the ISM Services website. At the top, there's a navigation bar with links for ISM Services, CODES (access to codes), ISMDB (simulations database), TECHNOLOGIES (standards), PARTNERS (credits), and REGISTRATION. Below the navigation is a large image of a nebula. Underneath the image are four cards: PDR Code (The Meudon PDR code), DustEM (Dust Emission), Shock (Paris-Durham Shock model), and Starformat (MHD simulations data base). A red "Beta" badge is visible next to the Shock card. At the bottom of the page, there's a section titled "ISM DataBase – Inverse Search service" with a "Beta" badge. It includes fields for "x" (Pgas\_0, cm<sup>-3</sup>, K) and "y" (G0 observer side, Mathis\_unit), both with "log scale" checkboxes. There are also fields for "AVmax" (mag) and "1". A search bar at the bottom left says "Search for available quantities... Ex: N(H)" with a "Use" button. A yellow box contains instructions: "Type quantities to plot in the input below, with optional constraints. (click Search to view the result of the example query below)". Below this, a search result for "I(CO v=0,J=1->v=0,J=0 angle 00 deg)" is shown, including its name, doc, range, and unit.

## Status

- Available at <http://ism.obspm.fr>
- Grids of PDR models

## Starts to be used:

- Individual teams: ALMA, IRAM, NOEMA, ...
- Projects as SPICA, GUSTO (NASA/CNES)
- JWST ERS

## Plans - short term

- Grids of shocks models
- Operations on quantities

## Plans - medium term

- Quickviews on models
- Interpretation of images (pixels)

## Plans - long term

- Search in N-dimension space
- Machine Learning techniques
- Suggestions of observations

