

Complex Datasets: Use-cases for Provenance DM

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Provenance DM

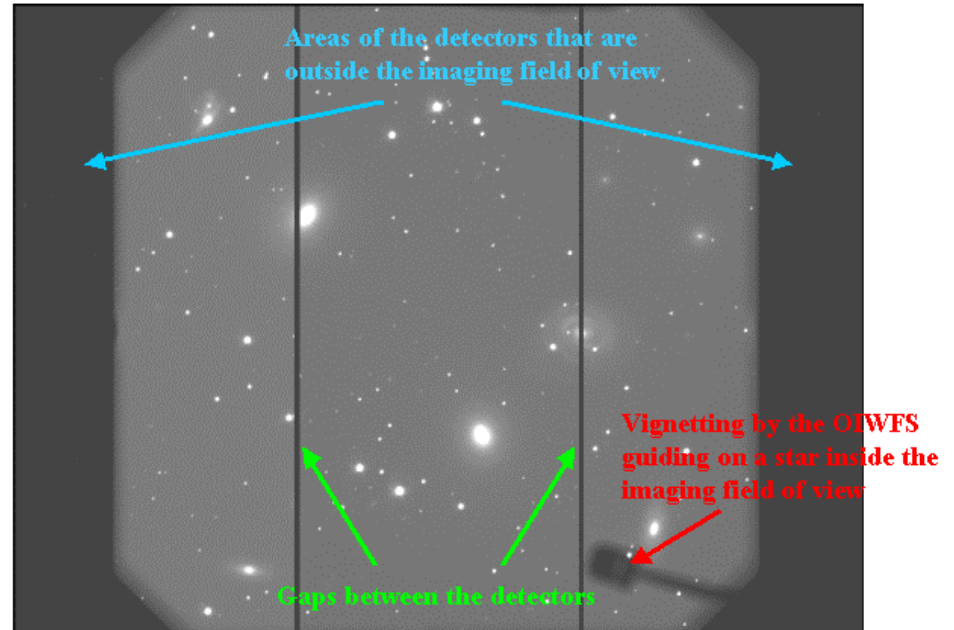
- An important block of the Observation DM
- Is supposed to describe the way an original dataset produced by the instrument was transformed/processed/reduced into the science ready form
- This may be very important, i.e. to be able to compute correctly the statistical properties of the data: error-bars, distributions, etc.

Complex datasets: examples

- GMOS (Gemini Multi-Object Spectrograph) – a multi-mode instrument:
 - imaging
 - long-slit spectroscopy
 - multi-slit spectroscopy
 - IFU spectroscopy
- Complex configuration:
 - 3 separate CCD chips forming a mosaic
 - **all figures on the following slides were stolen from the instrument web-pages at <http://www.gemini.edu/>**

Example 1

- Mosaic imaging
- 3 CCD chips:
 - mosaicing
 - offsets
 - rotations
 - different bias levels
 - different gain values
 - different spectral sensitivity curves => slightly different calibration solutions



Example 2

- Long-slit spectra

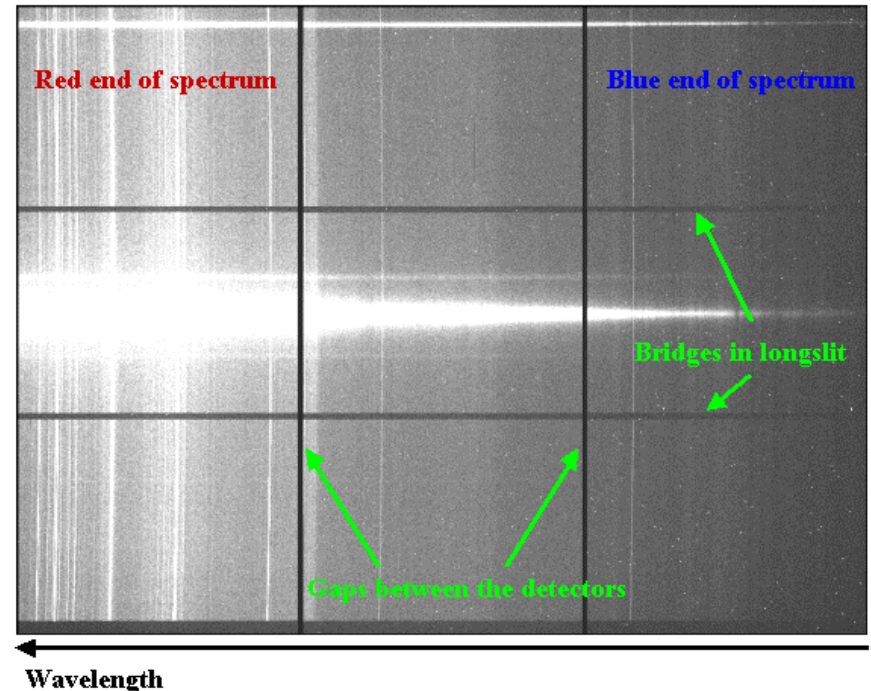
- In addition

- slit is curved

- CCD gaps are projected to different wavelengths at different slit positions

- different response along the slit

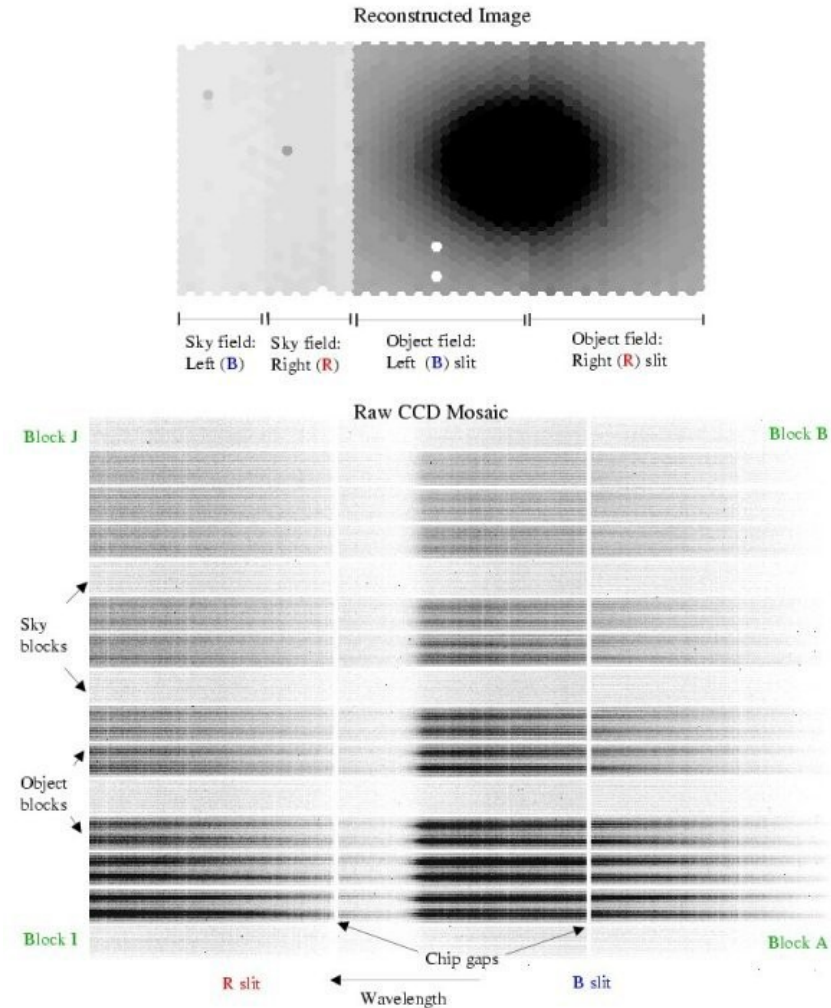
- dust, slightly variable slit width => affects the resolution



Example 3

- IFU spectroscopy
 - fiber traces as 1D-spectra
 - field-to-slit mapping
 - 1-slit mode
 - 2-slit mode: 2 pseudo-slits projected onto 3 CCD chips
 - different parts of the FoV in the reconstructed datacube contain data originally coming from different CCDs
 - non-trivial flux calibration

GMOS IFU Example Data: NGC 221



Other examples:

- Multi-CCD HST Imagers (WFPC2, ACS) and mosaic wide-field imagers (e.g. **MEGACAM@CFHT**) widely available
- **VIMOS@VLT**
 - 4 independent spectral units (different gratings, CCDs, optical distortions, etc.) for 4 quadrants of the FoV. Will be even worse for MUSE: 24 units
- **X-shooter@VLT**
 - Different types of detectors for different wavelength from optical to NIR