

Time Series in Exoplanet Research



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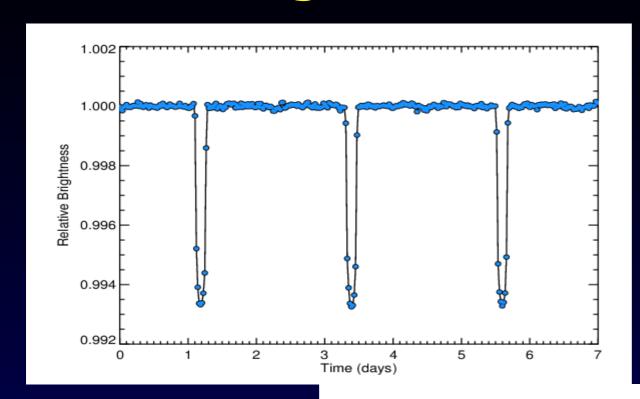
IVOA Interoperability Meeting
Time Domain Interest Group Session
Goerlitz, Germany
November 14th 2025

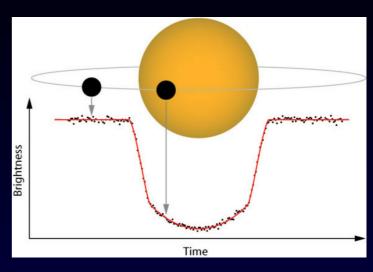
Credits

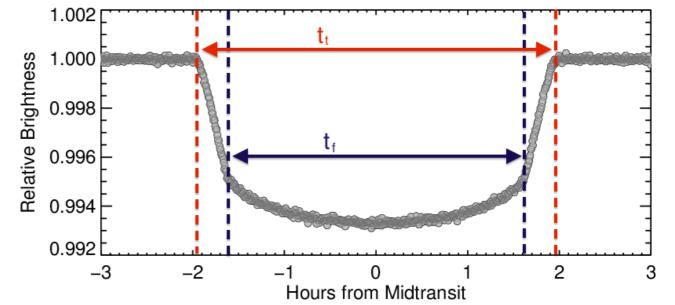
The presentation is based on many different sources – mainly the on-line published slides from IVOA meetings, EPRV group slides and docs, ADASS meetings, tutorials, manuals or pictures found on Internet.

Light Curve

Light Curve HAT-P-7b



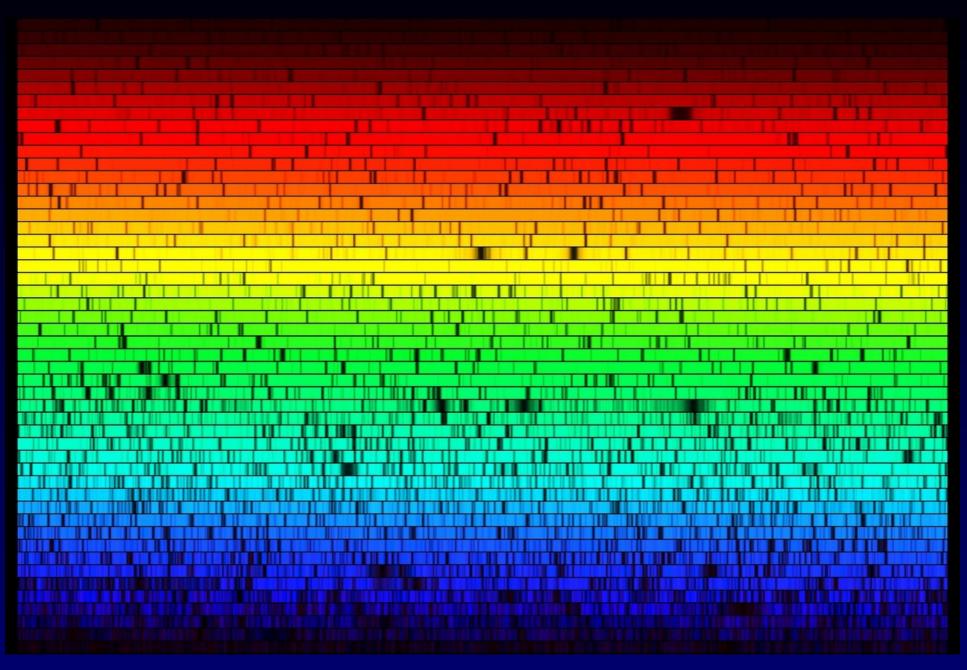




Vanderburg tutorial

Echelle Spectra

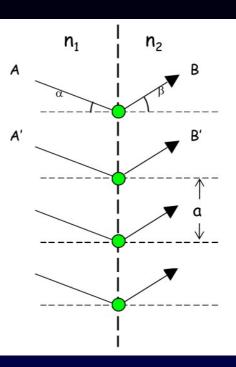
Spectrum of Arcturus (pixel-order)



Grating Equation

- Interference condition:
- \Rightarrow path difference between AB and A'B'
- Grating equation:

$$m\rho\lambda = n_1 \sin\alpha + n_2 \sin\beta$$
 where $\rho = \frac{1}{a}$



Note the *m* is the spectrum (interference) order

High order – high dispersion $R \sim m^*N$ (no of rules in beam)

Èchelle – stairway in French - coarse grooves (big steps) HARPS 31.6/mm

Echellette – normal grating in 1st, 2nd order e.g. 833 /mm

Orders Define Geometry

Multiple orders

 Many orders to cover desired λλ:

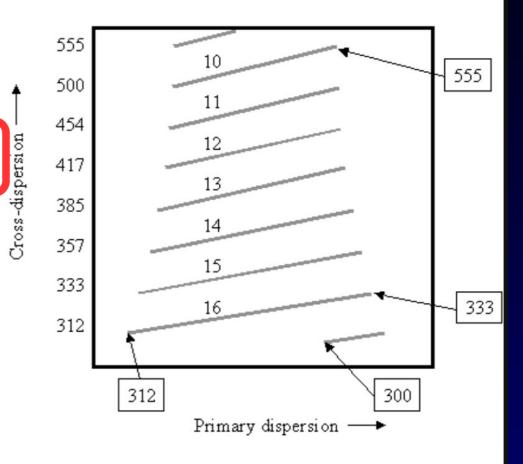
Free spectral range

$$\Delta \lambda = \lambda/m$$

Orders lie on top of each other:

$$\lambda(m) = \lambda(n) \times (n/m)$$

- Solution:
 - use narrow passband filter to isolate one order at a time
 - cross-disperse to fill detector with many orders at once



Cross dispersion may use prisms or low dispersion grating

ESO X-Shooter

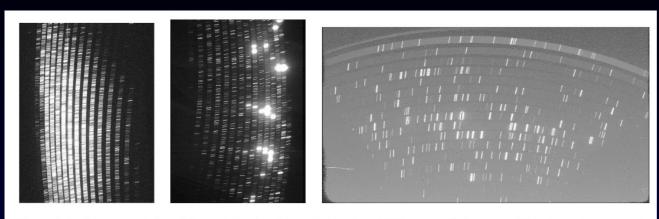
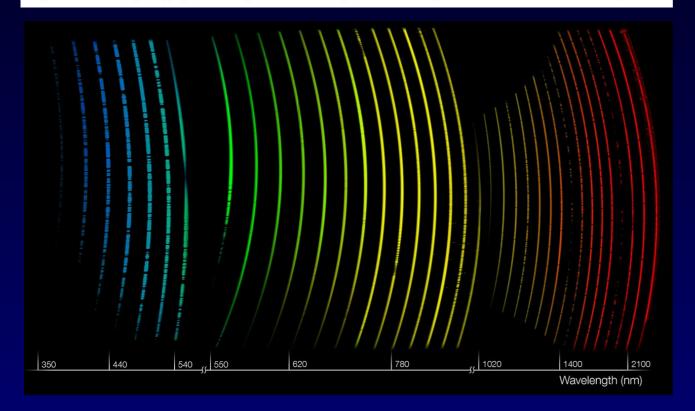
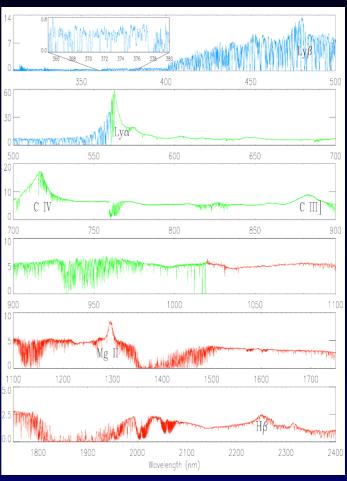


Figure 1. This image displays left to right the XSH UVB, VIS and NIR spectral formats with different image scales.



2936-24807 A, R ~9100,17400,11300 0.5arcsec slit



High-z QSO Vernet et al. 2011

HARPS-S

71 orders (89-161) – 115 not

3780-6910 A

R~115000

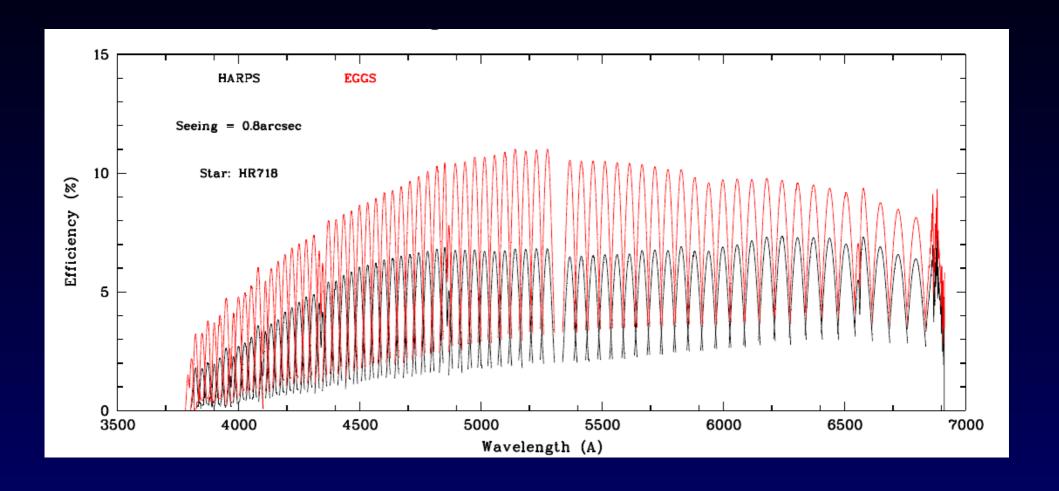
dRV~3m/s

Simultaneus ThAr/Laser comb/Fabry Perot

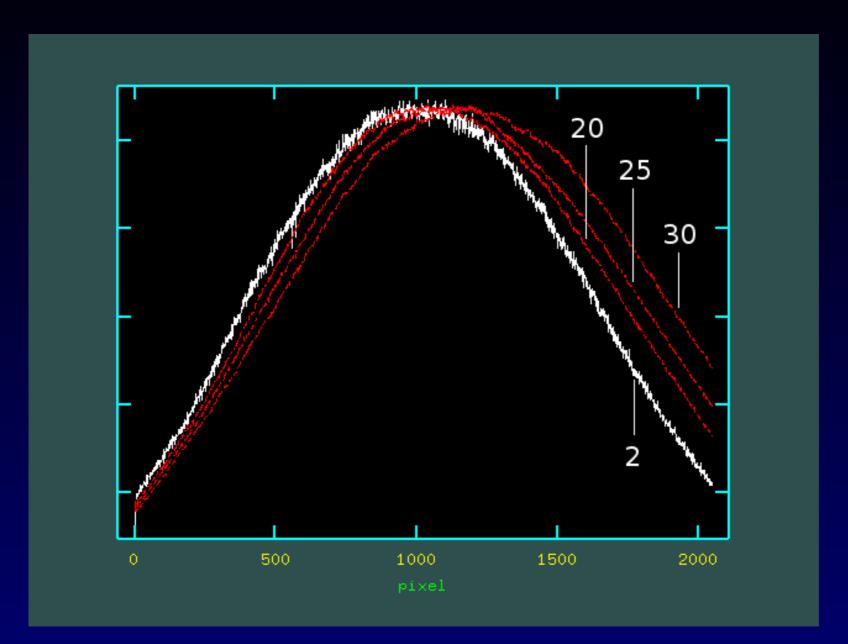
https://cdn.eso.org/images/large/eso0308c.jpg

PlatoSpec simcal ThAr

HARPS-S

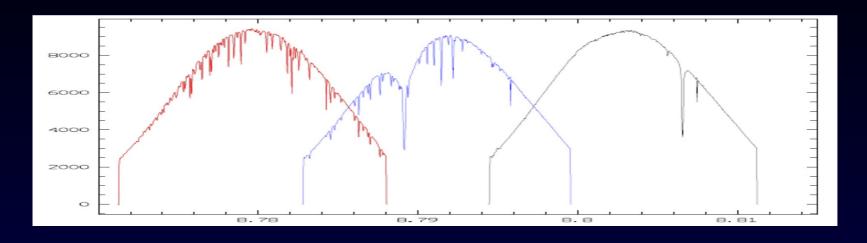


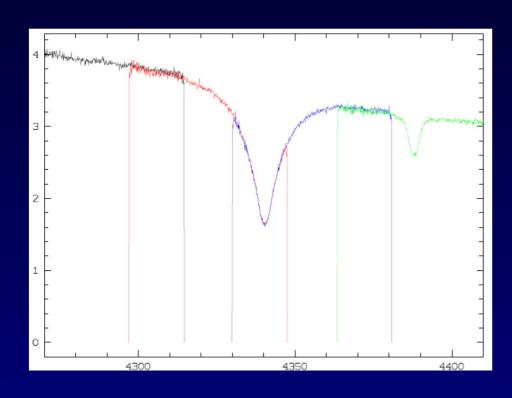
OES Blaze shift

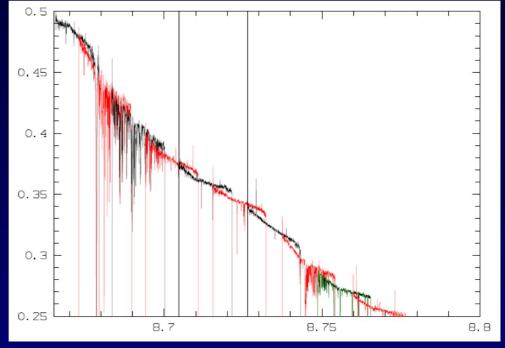




Echelle Spectra Problems in Hot Stars







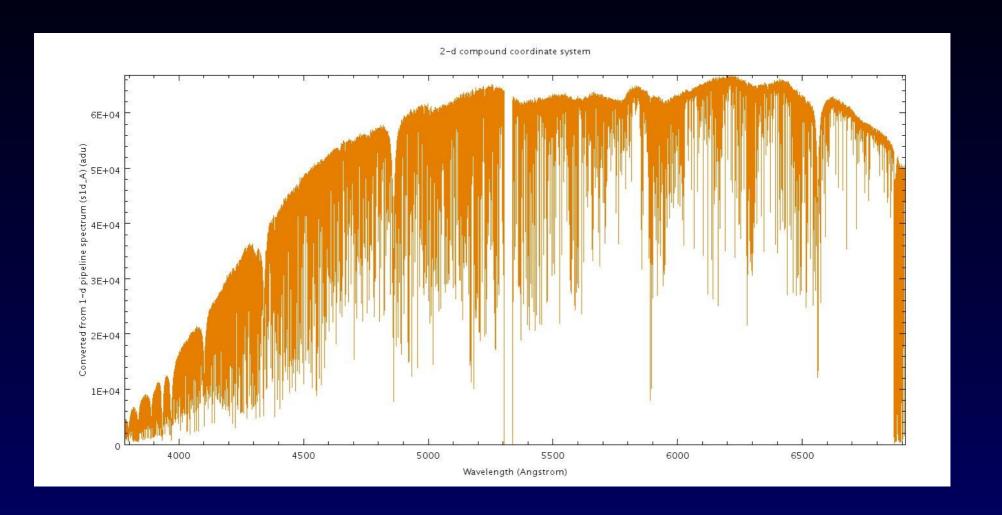
(Simple) Spectra in VO

VO Tools

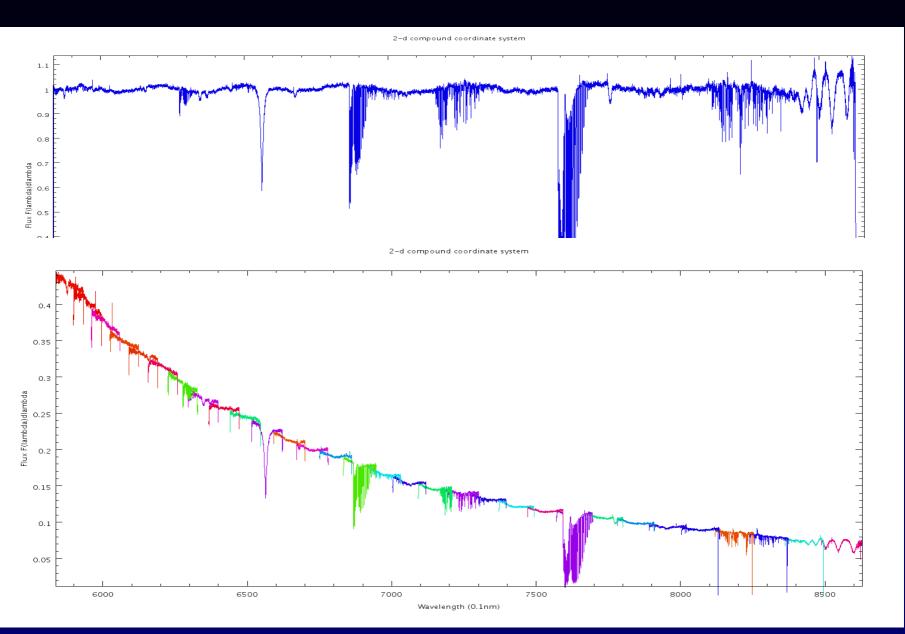
- SPLAT-VO (GAVO Ondrejov)
- CASSIS (OMP)
- IPAC Firefly (IRSAViewer)
 - All NASA data, LSST (Rubin)
 - Spectra in echelle started
 - https://github.com/Caltech-IPAC/firefly

- Not Supported:
 - SpecView (HST)
 - VOSpec (ESA)

HARPS-S in SPLAT



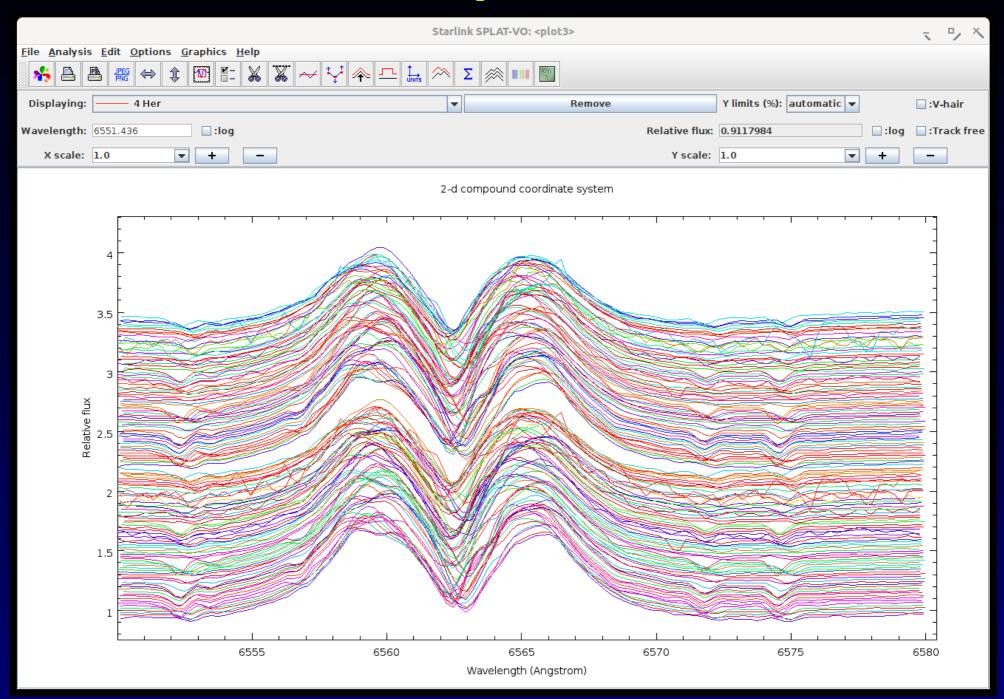
DaCHS Testbed Split-Order SSAP



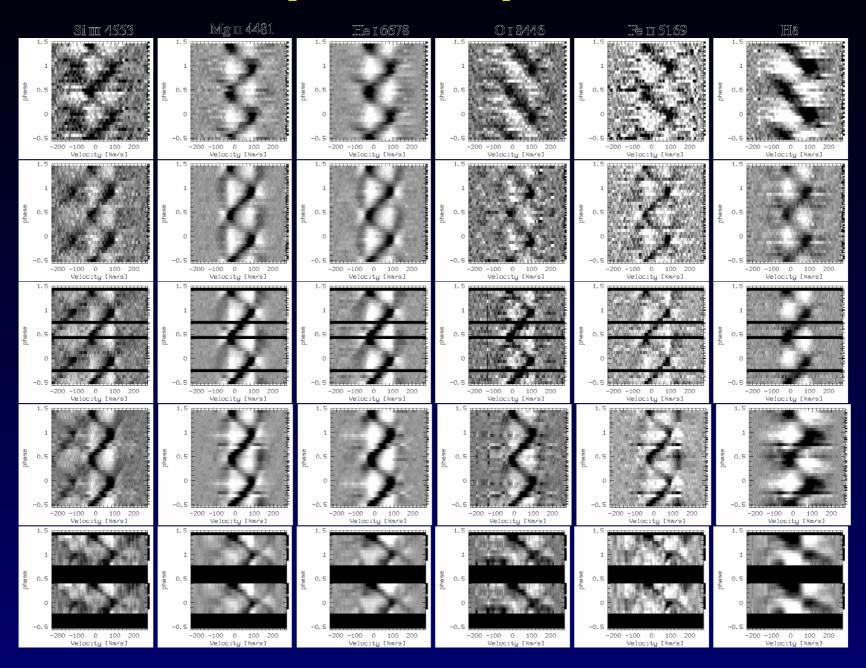
Products derived from spectra

Time Series of Spectra

Series of Spectra – LPV

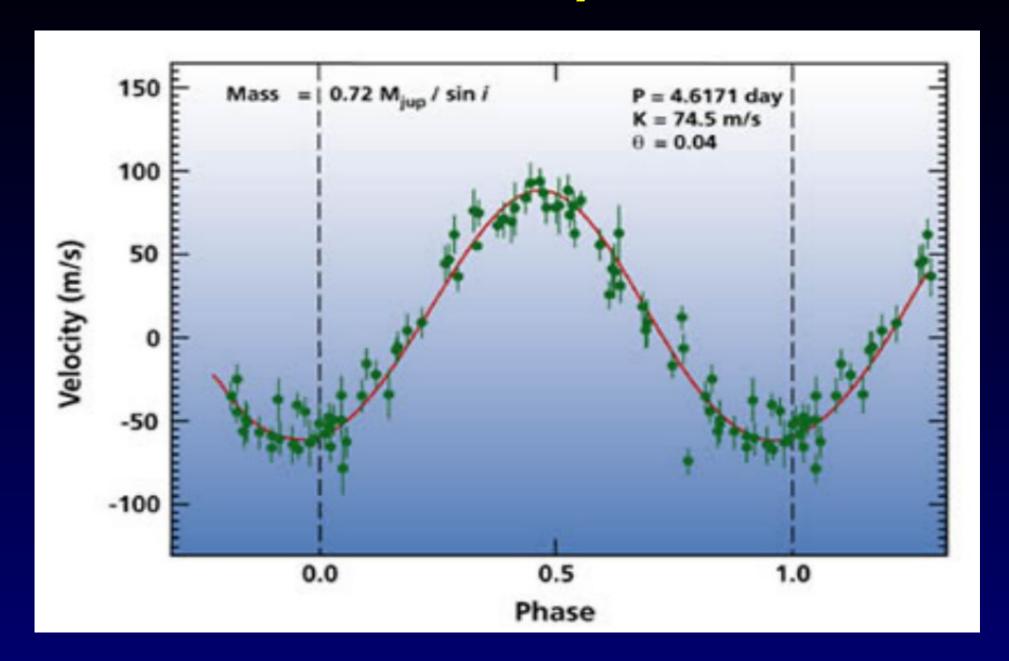


Dynamic Spectra

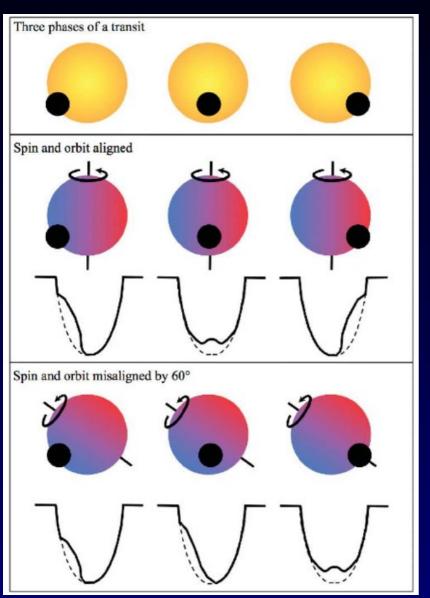


Radial Velocity Curve

Radial Velocity Curve

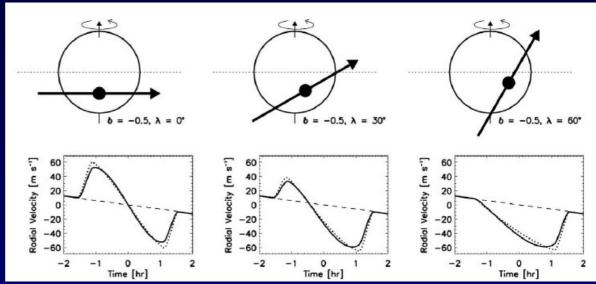


Rossiter-McLaughlin effect



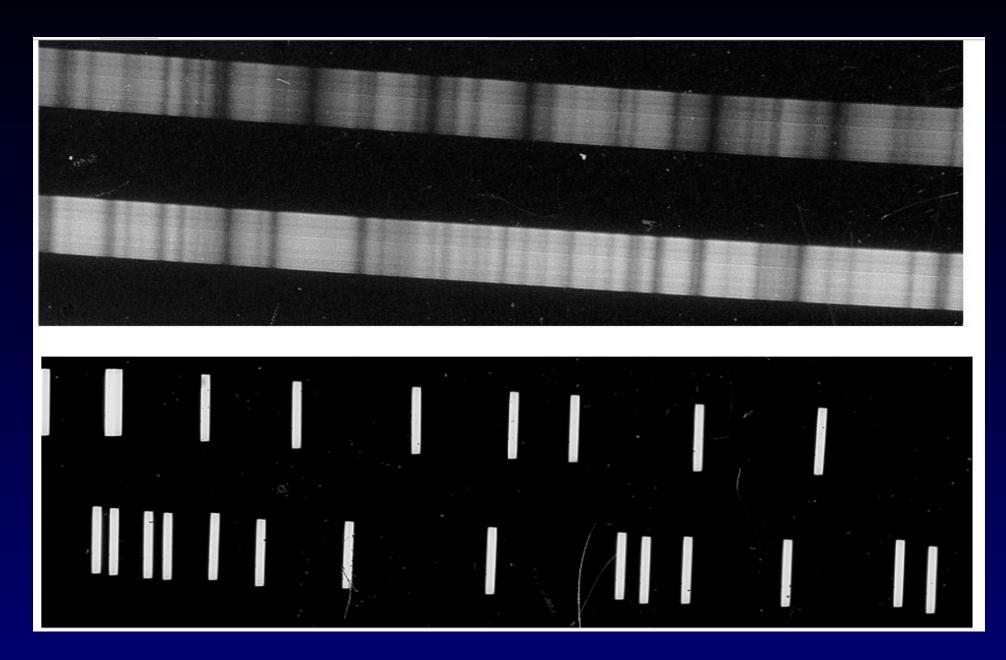
RV Curve – time series of RV

Only for direct occultation

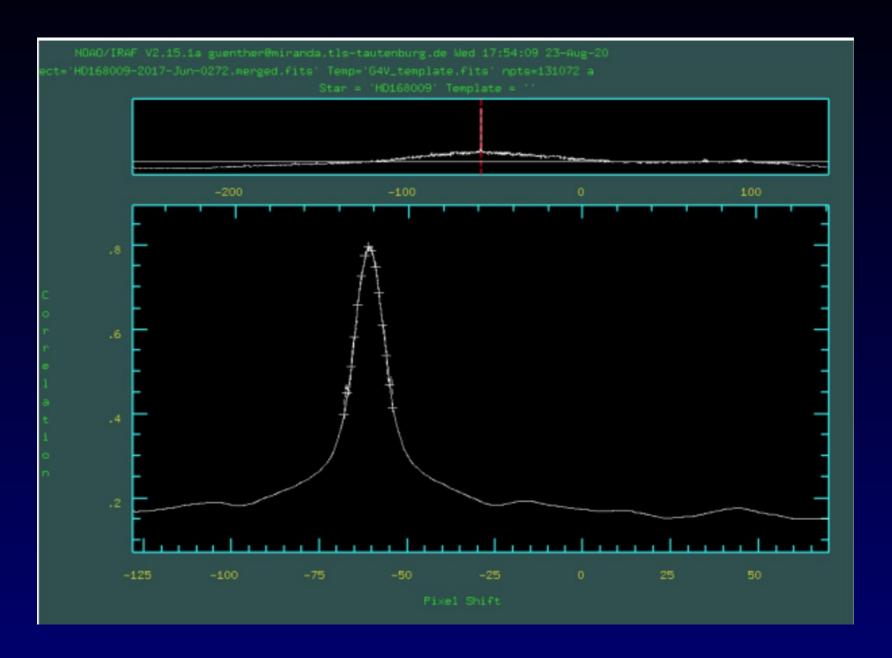


Cross Correlation Function

Analogue Technique

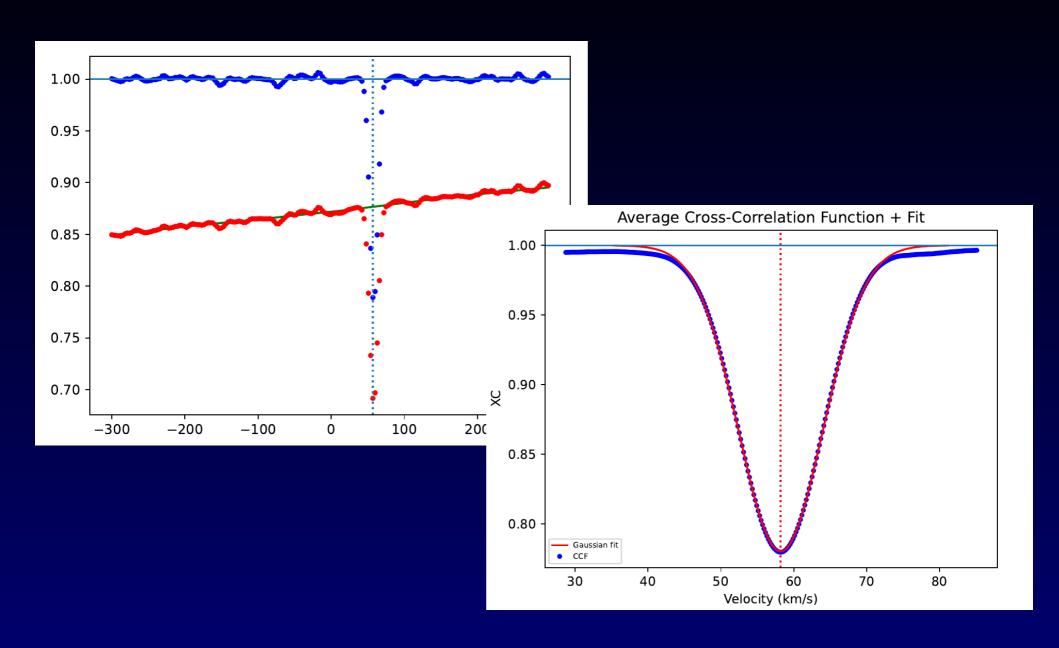


Cross Correlation Function

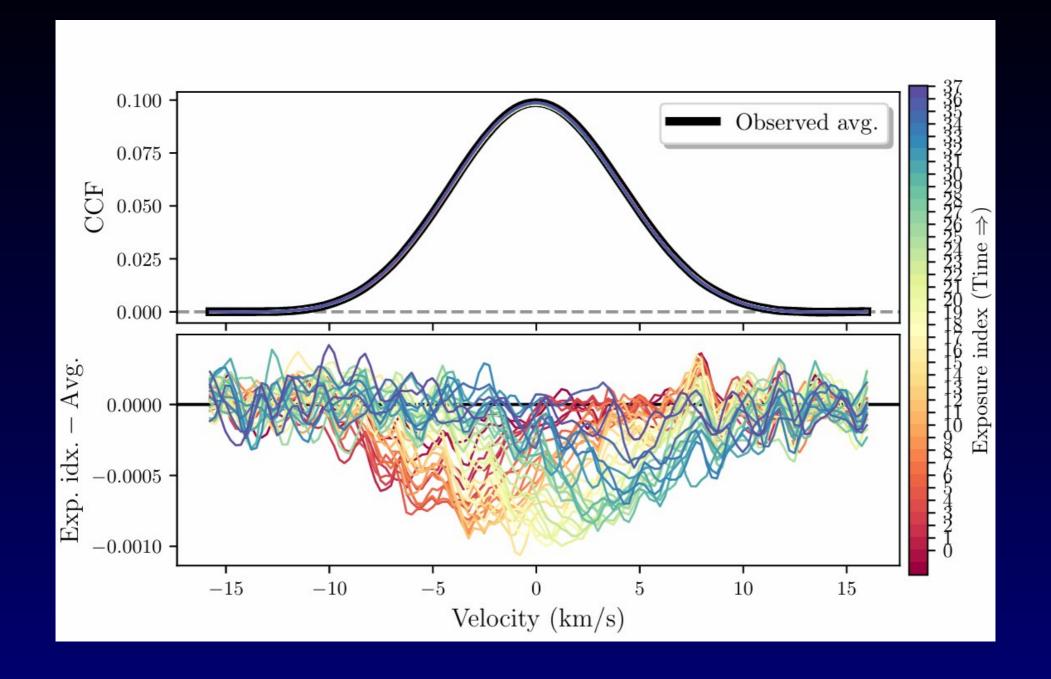


IRAF fxcor task

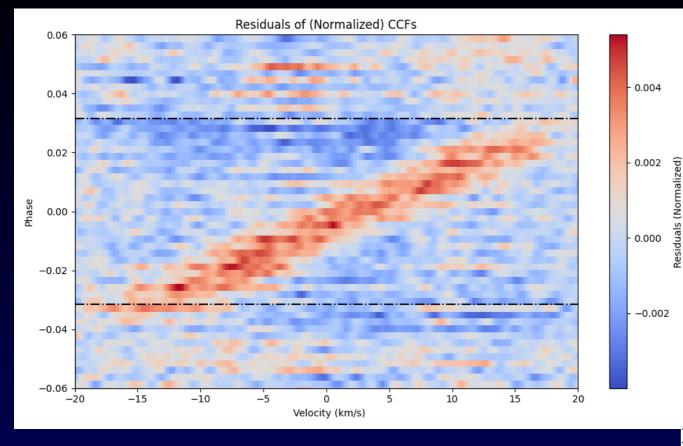
Cross Correlation Function - CCF



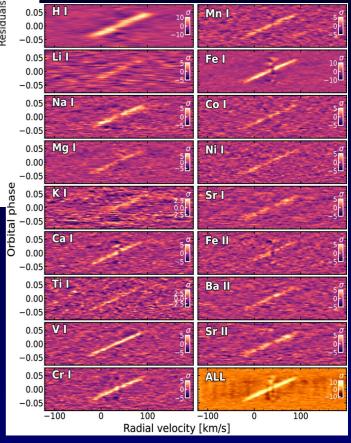
Residual CCF



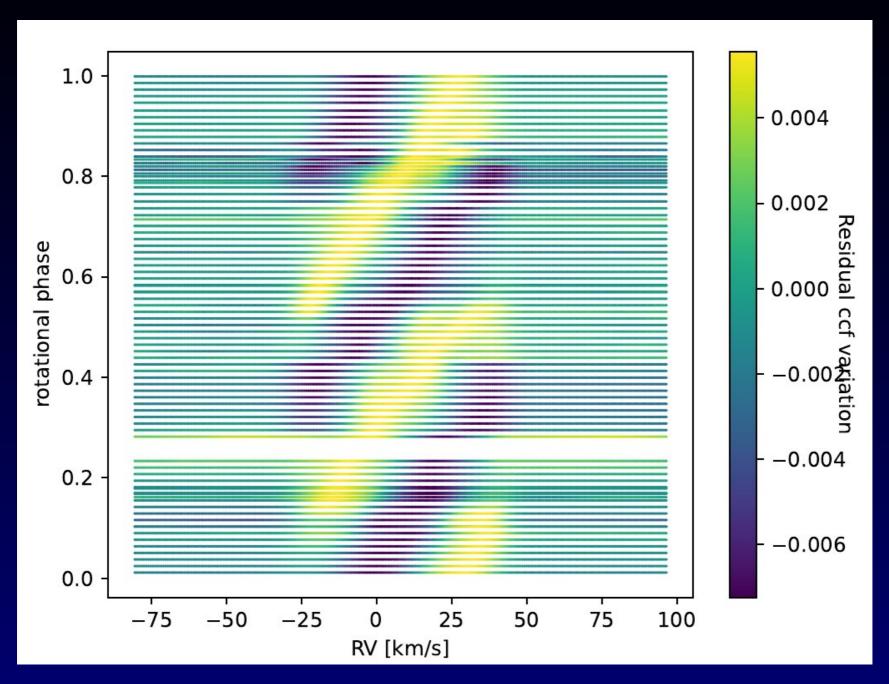
Doppler Shadow



HARPS-N archival CCF of HAT-P-41A. Balkoová 2025



Dynamic spectra – CCF



Lipták 2025, private comm.

EPRV Standard

- About 15 spectrographs (~ 5 passively following)
 e.g. CARMENES)
- Translators from native format
- Started 2024
- NASA JPL (EPRV project WG)
- Bruce Berriman at IVOA 2025 June (after my advice to contact IVOA)

https://github.com/EPRV-RCN/RVData

https://eprv-data-standard.readthedocs.io/en/latest

EPRV L2 – Order by Order Unblazed

HDU	Name	DataType	MinBetDepth	Multiplicity	Required	Description
0	PRIMARY	PrimaryHDU		False	True	EPRV Standard FITS HEADER (no data)
1	INSTRUMENT_HEADER	ImageHDU		False	True	Inherited instrument header (no data)
2	RECEIPT	BinTableHDU		False	True	Table of operations that have been performed on this file
3	DRP_CONFIG	BinTableHDU		False	True	Pipeline details (settings etc) to go from native data to L2
4	EXT_DESCRIPT	BinTableHDU		False	True	Table describing contents of each extension
5	ORDER_TABLE	BinTableHDU		TRUE	TRUE	Table capturing the wavelength extent of each order
6	TRACE1_FLUX	ImageHDU		True	True	Flux in trace 1
7	TRACE1_WAVE	ImageHDU	64	True	True	wavelength solution for trace 1
8	TRACE1_VAR	ImageHDU		True	True	variance for trace 1
9	TRACE1_BLAZE	ImageHDU		True	True	blaze for trace 1
10	BARYCORR_KMS	ImageHDU		False	True	barycentric correction in km/s
11	BARYCORR_Z	ImageHDU		False	True	barycentric correction in redshift
12	BJD_TDB	ImageHDU	64	False	True	Photon weighted midpoint
13	DRIFT	ImageHDU		False	False	Drift measurement map in delta lambda
14	TRACE1_QUALITY	ImageHDU	UINT8	True	False	Quality of each pixel in trace 1
15	EXPMETER	BinTableHDU		False	False	Table of exposure meter counts timeseries over the exposure
16	TELEMETRY	BinTableHDU		False	False	Table of telemetry collected during the exposure
17	TRACE1_TELLURIC	ImageHDU		True	False	Telluric model for trace 1
18	TRACE1_SKYMODEL	ImageHDU		True	False	Sky model for trace 1
19	ANCILLARY_SPECTRUM	ImageHDU		True	False	Extension(s) that store ancillary spectra
20	IMAGE	ImageHDU		True	False	Extension(s) that store useful support images
21	CUSTOM1_TRACE1_FLUX	ImageHDU		True	False	Additionally corrected flux in trace 1
22	CUSTOM1_TRACE1_WAVE	ImageHDU		True	False	Additionally corrected wavelength solution for trace 1
23	CUSTOM1_TRACE1_VAR	ImageHDU		True	False	Additionally corrected variance for the flux in trace 1

EPRV L3 – Merged orders normalized

This is what we already know – SSAP, "distorted data low precision"

HDU	Name	DataType	Multiplicity	Required	Description
0	PRIMARY	PrimaryHDU	False	True	EPRV Standard FITS HEADER (no data)
1	INSTRUMENT_HEADER	ImageHDU	False	True	Inherited instrument header (no data)
2	RECEIPT	BinTableHDU	False	True	Table of operations that have been performed on this file
3	DRP_CONFIG	BinTableHDU	False	True	Pipeline details (settings etc) to go from native data to L2
4	ORDER_TABLE	BinTableHDU	TRUE	TRUE	Table capturing the wavelength extent of each order
5	STITCHED_CORR_TRACE1_FLUX	ImageHDU	True	True	Order stitched blaze-corrected flux in trace 1
6	STITCHED_CORR_TRACE1_WAVE	ImageHDU	True	True	Order stitched BC- and drift-corrected wavelength solution for trace 1
7	STITCHED_CORR_TRACE1_VAR	ImageHDU	True	True	Order stitched variance for the flux in STITCHED_CORR_TRACE1_FLUX
8	COMBINED_STITCHED_CORR_FLUX	ImageHDU	True	True	Order stitched and blaze-corrected flux co-added across all traces
9	COMBINED_STITCHED_CORR_WAVE	ImageHDU	True	True	Order stitched BC- and drift-corrected wavelength solution
10	COMBINED_STITCHED_CORR_VAR	ImageHDU	True	True	Order stitched variance for the combined flux in COMBINED_STITCHED_COR
11	STITCHED_CUSTOMCORR1_TRACE1_FLUX	ImageHDU	True	False	Additional corrections made to STITCHED_CORR_TRACE1_FLUX
12	STITCHED_CUSTOMCORR1_TRACE1_WAVE	ImageHDU	True	False	Wavelength solution corresponding to STITCHED_CUSTOMCORR1_TRACE1_
13	STITCHED_CUSTOMCORR1_TRACE1_VAR	ImageHDU	True	False	Variance corresponding to STITCHED_CUSTOMCORR1_TRACE1_FLUX
14	COMBINED_STITCHED_CUSTOMCORR1_FLUX	ImageHDU	True	False	Additional corrections made to COMBINED_STITCHED_CORR_FLUX
15	COMBINED_STITCHED_CUSTOMCORR1_WAVE	ImageHDU	True	False	Wavelength solution corresponding to COMBINED_STITCHED_CORR_WAVE
16	COMBINED_STITCHED_CUSTOMCORR1_VAR	ImageHDU	True	False	Variance corresponding to COMBINED_STITCHED_CORR_VAR

EPRV L4 – Derived core exo science

Radial Velocity and Cross Correlation Function (various masks)

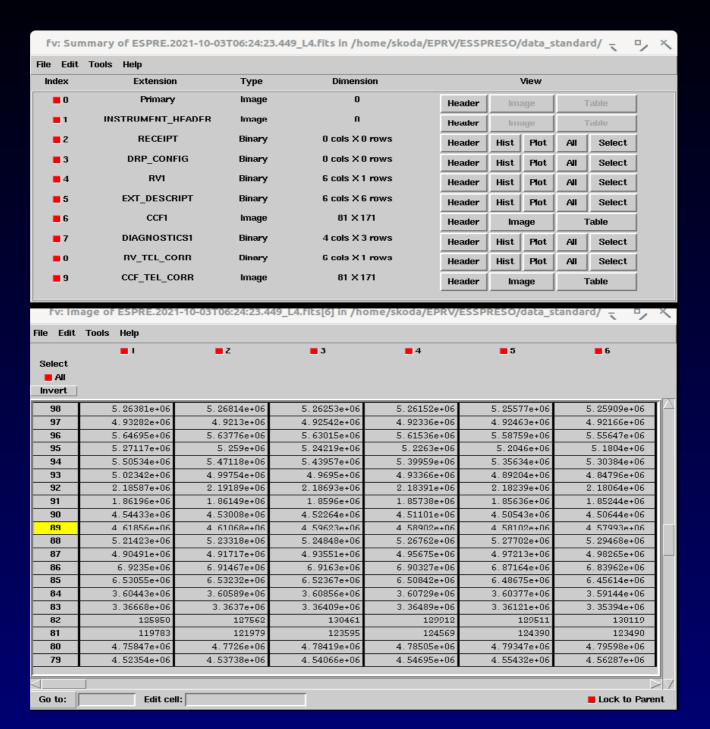
It is the most important output of pipeline

But exoscientists need Time Series (RV curves, Doppler shadows)

VO can do it on the fly (select all observation, apply SODA extract, create time series)

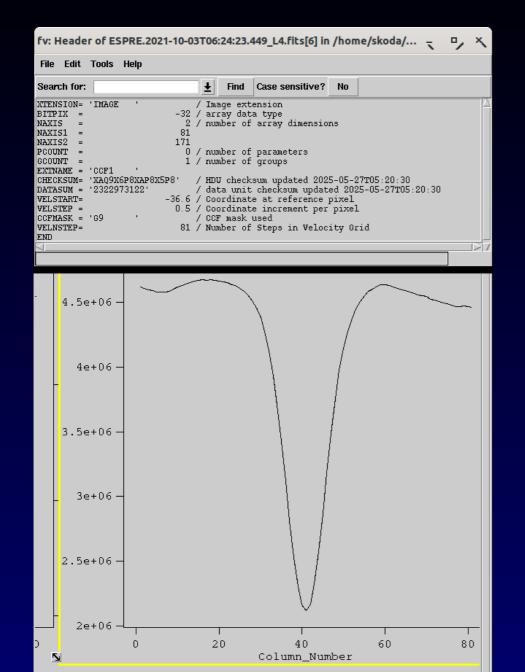
HDU	Name	DataType	Multiplicity	Required	Description
0	PRIMARY	PrimaryHDU	FALSE	TRUE	EPRV Standard FITS HEADER (no data)
1	INSTRUMENT_HEADER	ImageHDU	FALSE	TRUE	Inherited instrument header (no data)
2	RECEIPT	BinTableHDU	FALSE	TRUE	Table of operations that have been performed on this file
3	DRP_CONFIG	BinTableHDU	FALSE	TRUE	Pipeline details (settings etc) to go from native data to L2
4	RV1N	BinTableHDU	TRUE	TRUE	Derived Radial Velocity Measurement
5	CCF1N	ImageHDU	TRUE	FALSE	Array with same dimensionality of RV1N, that contains the CCF that produced each RV1N
6	DIAGNOSTICS1J	BinTableHDU	TRUE	FALSE	Activity indicators, CCF metrics, etc
7	CUSTOM_CCF1N	ImageHDU	TRUE	FALSE	Additional CCFs from (e.g.) different masks
8	CUSTOM_RV1N	BinTableHDU	TRUE	FALSE	Derived Radial Velocity Measurement from CUSTOM_CCF1N

EPRV L4 – ESPRESSO Tau Cet



Thanks: E.Fontanet 2025 priv.com.

EPRV L4 CCF in fv



VLSTART=-36.6 km/s VELSTEP=0.5 km/s

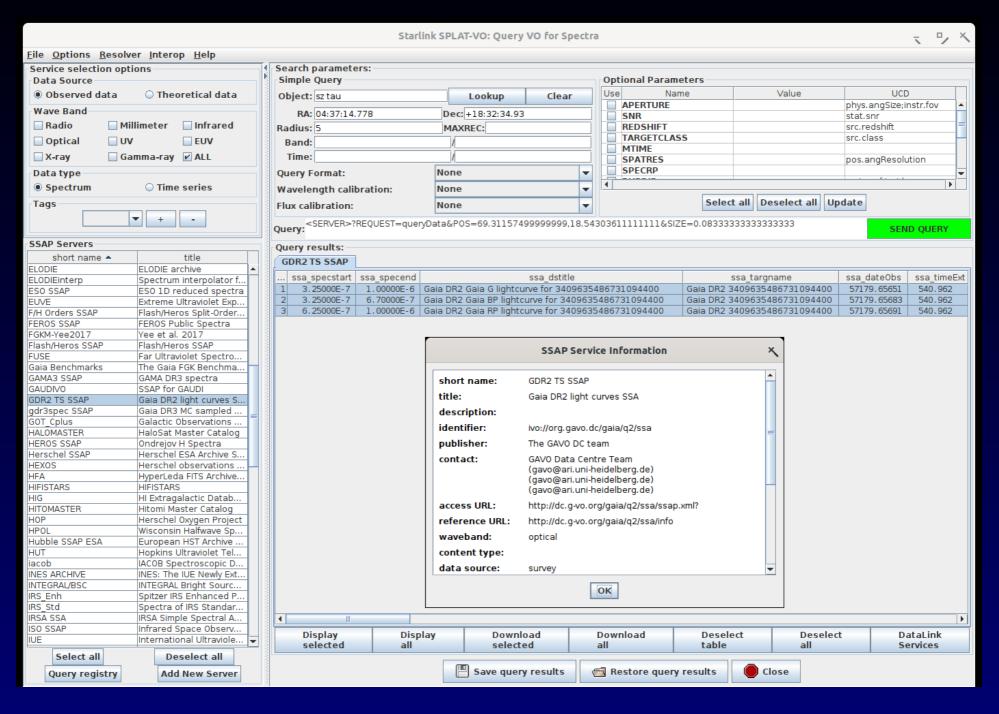
Center~41

-36.6+40*0.5=-16.6

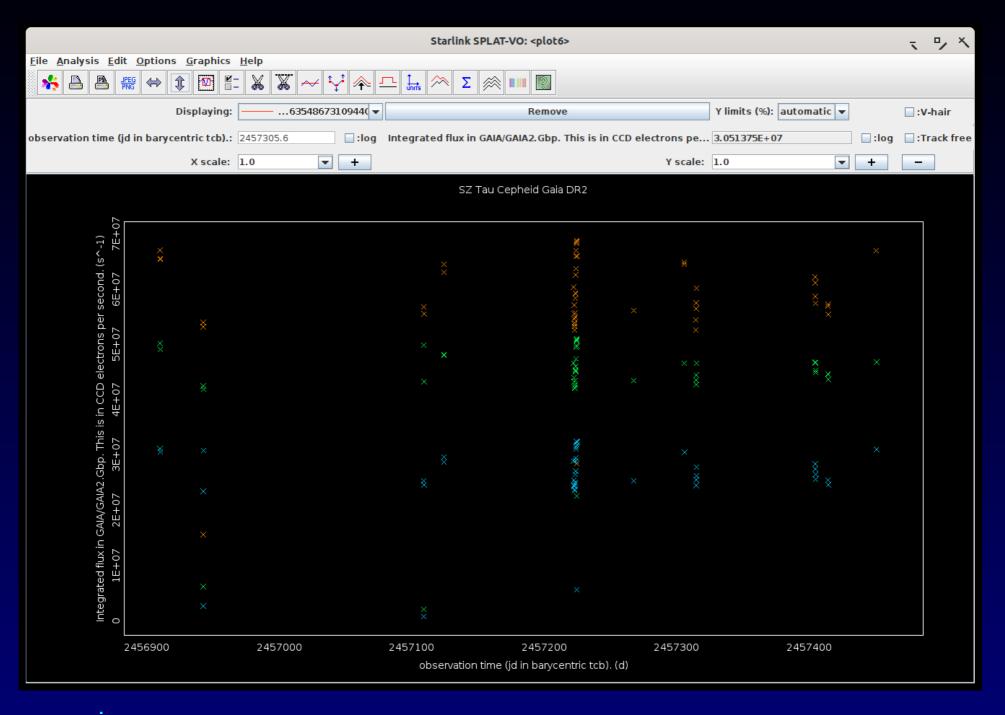
Astrometry				
Radial velocity (R _v)	$-16.68 \pm 0.05^{[5]}$ km/s			
Proper motion (μ)	RA: -1721.728 mas/yr ^[1]			
	Dec.: +854.963 mas/yr ^[1]			
Parallax (π)	273.8097 ± 0.1701 mas ^[1]			

Thanks: E.Fontanet 2025 priv.com.

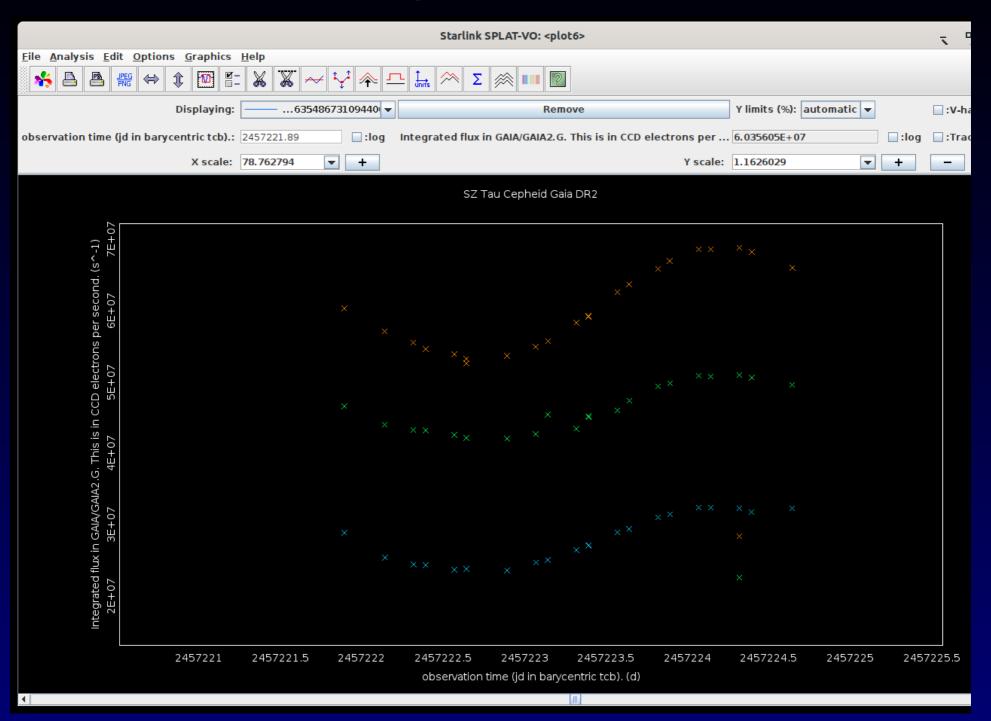
Light Curves Gaia DR2 - Query



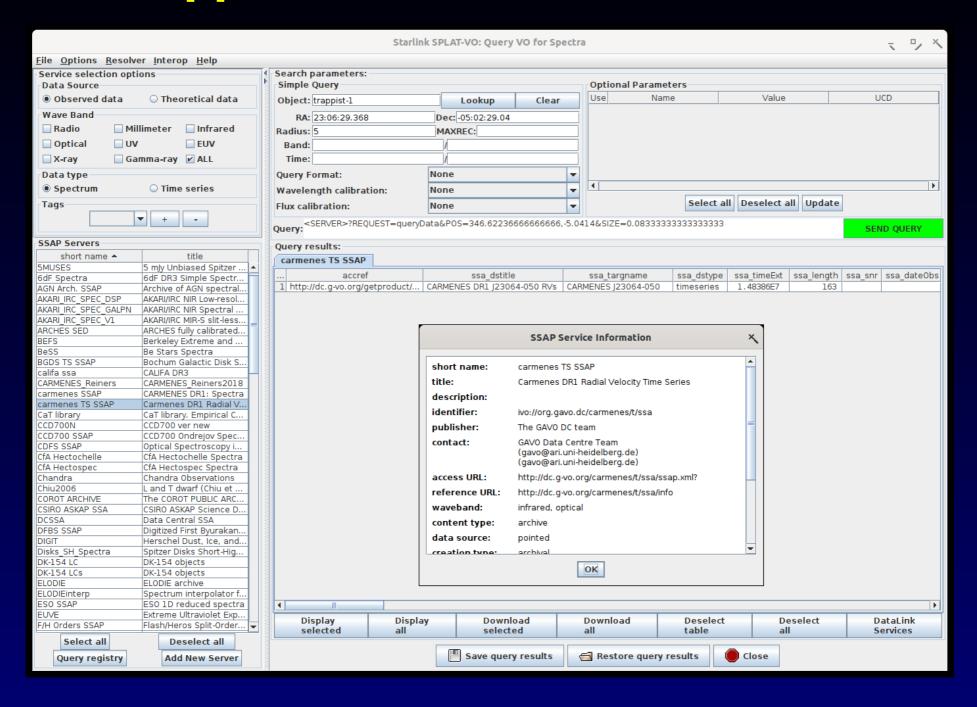
Gaia DR2 Light Curves



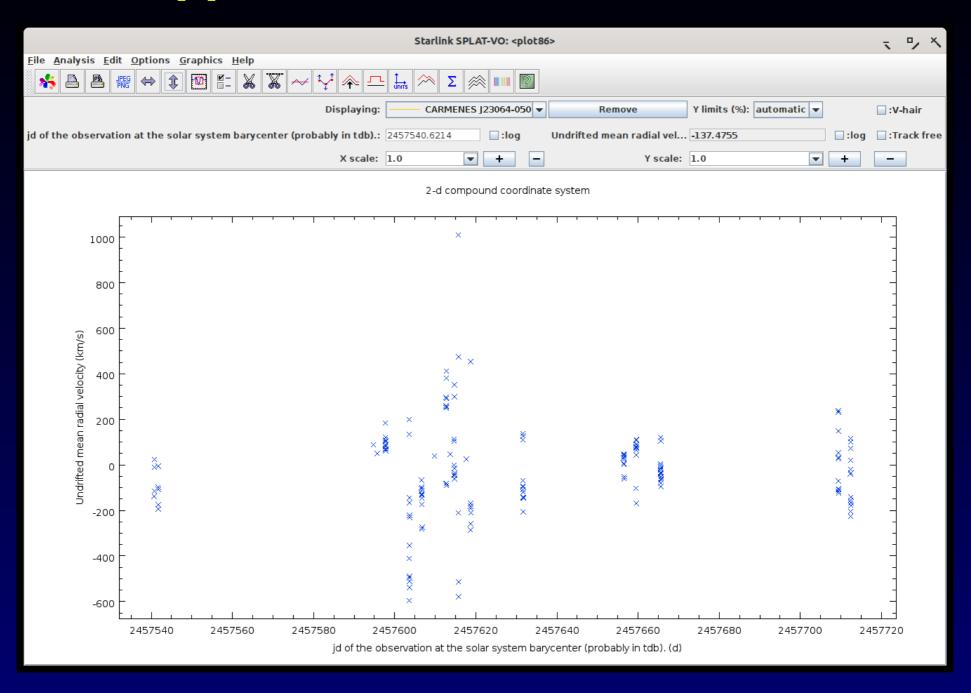
Gaia DR2 Light Curves - Zoom



Trappist-1 RV Curve - Carmenes



Trappist-1 RV Curve – Carmenes



AST Library – SSAP "flux"

Nov 12, 2025 9:21:20 PM uk.ac.starlink.splat.ast.ASTJ makeSpectral INFO: Data units unknown: km/s (AST: Error at line 420 in file AstObject.c. astGetSystem(FluxFrame): The current units (km/s) cannot be used with any of the supported flux systems.

Nov 12, 2025 9:48:08 PM uk.ac.starlink.splat.ast.ASTJ makeSpectral

INFO: Data units unknown: s^-1 (AST: Error at line 420 in file AstObject.c.

astGetSystem(FluxFrame): The current units (s^-1) cannot be used with any of the supported flux systems.

4.2.5.4 General Dataset Metadata

General dataset metadata describes the overall dataset.

UTYPE Description		Req	Default	
Dataset.DataModel	Datamodel name and version	MAN	Spectrum-1.0	
Dataset.Type	Type of dataset	OPT	Spectrum	
Dataset.Length	Number of points in spectrum	MAN		
Dataset.Deleted	Set to deletion time, if dataset is deleted	OPT		
Dataset.TimeSI	SI factor and dimensions	OPT		
Dataset.SpectralSI	SI factor and dimensions	OPT		
Dataset.FluxSI	SI factor and dimensions	OPT		
Dataset.SpectralAxis	SpectralAxis column name (native data)	OPT		
Dataset.FluxAxis	FluxAxis column name (native data)	OPT		

SSAP

4.2.5.9 Dataset Characterization Axis Metadata

The Characterization axis metadata specifies the type of physical quantity on each physical measurement axis as well as the observable.

UTYPE	Description	Req	Default	
Char.FluxAxis.Ucd	ucd for flux	REC		
Char.SpectralAxis.Ucd	ucd for spectral coord	REC	V:	

Conclusions – Are we Alone?

To discover The Second Earth we need:

- Satellites for precise photometry
- Extremely precise spectrographs (on ELTs)
- EPRV like format as an unification standard
- VO for federation across world
- Some protocol based on SSAP approach
- Light curves (multi-color)
- •RV curves (in different spectral lines, ranges)
- Graphic tool with SPLAT-VO capabilities
- Enthusiastic VO-educated experts!

Thank you