

Automated Event Classification in Synoptic Sky Surveys

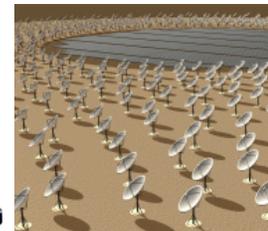
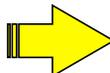
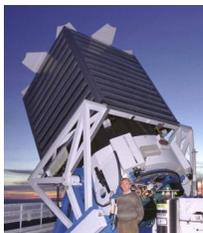
S. G. Djorgovski (*Caltech*)

With: A. Mahabal, C. Donalek, M. Graham, A. Drake, B. Moghaddam, M. Turmon, and many students and collaborators

IVOA InterOp, Naples, May 2011

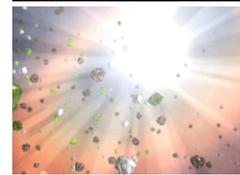
The Evolving Data-Rich Astronomy

- Digital sky surveys have brought us into the Terascale regime, and stimulated:
 - Extensive use of databases in astronomy
 - The Virtual Observatory concept
 - Incipient data-mining-based astronomy
- Synoptic digital sky surveys – i.e., panoramic cosmic cinematography – are moving us into the Petascale regime
 - The same old challenges, only more so
 - New challenges: real time response, event classification, data mining in the time domain...



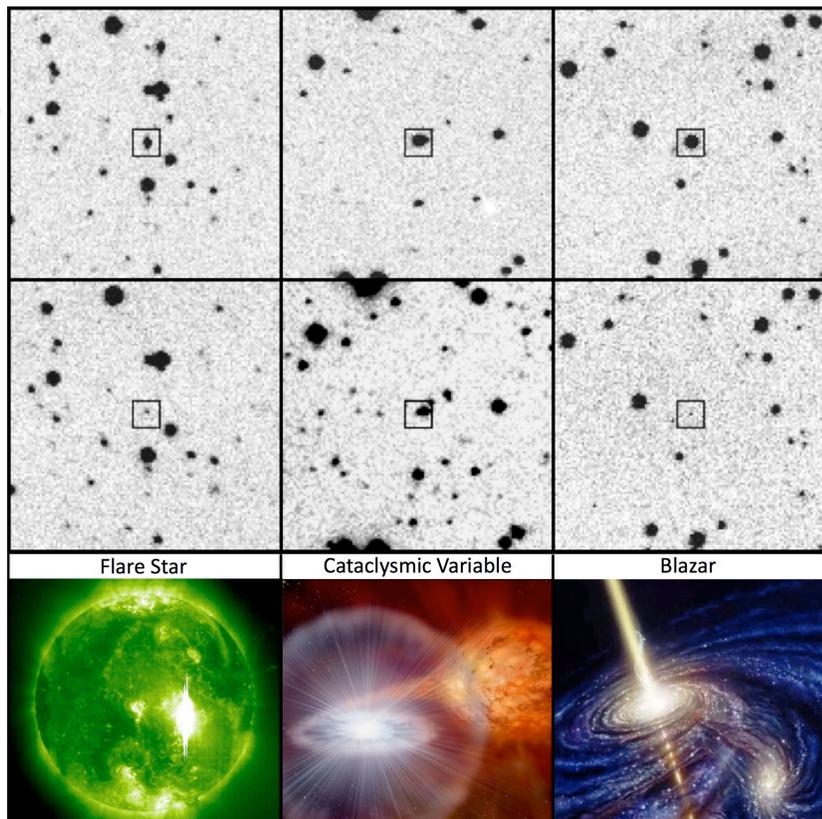
Astronomy in the Time Domain

- A major new growth area of astrophysics
- Driven by the new generation of large digital synoptic sky surveys, leading to LSST, SKA, etc.
- Rich phenomenology, from the Solar system to cosmology and extreme relativistic physics
 - For some phenomena, time domain information is a key to the physical understanding
- Transformational in many ways:
 - Static → Dynamic sky
 - Sources → Events
- Real-time discovery in massive data streams poses new challenges in automated classification, anomaly detection, decision making, etc.

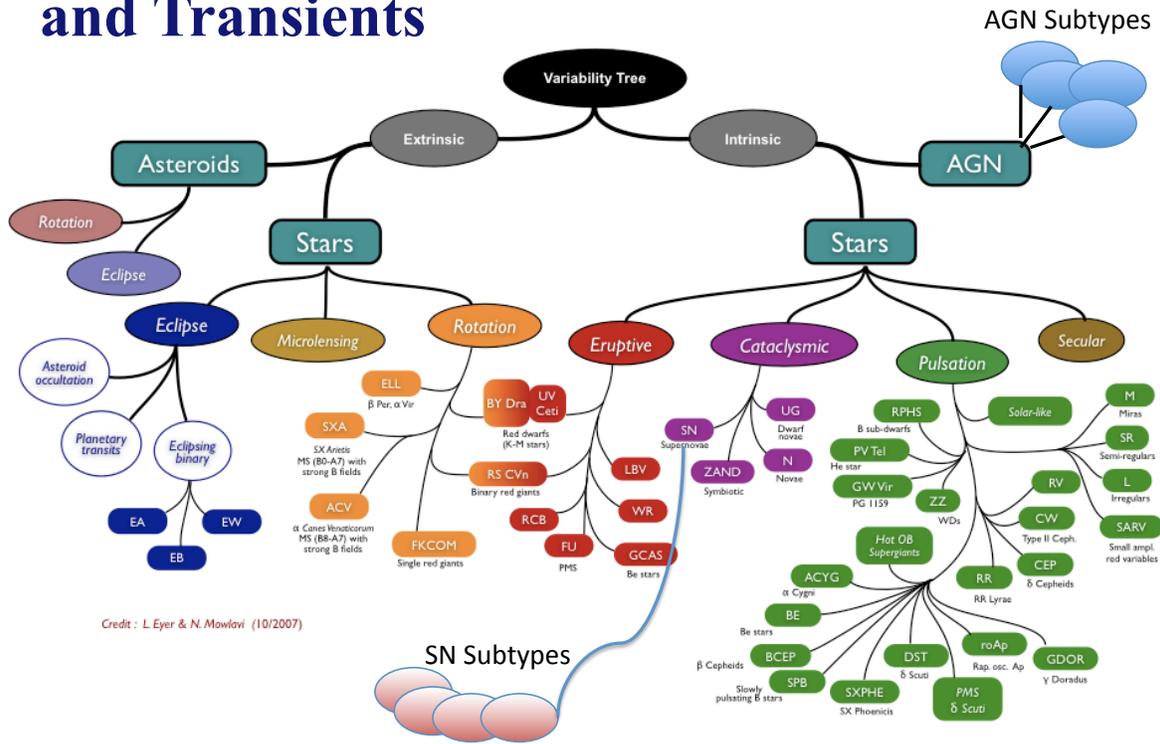


**All transients
look the same,
but can
represent
vastly
different
physical
phenomena**

Which ones are the most interesting and worthy of follow-up efforts?



Semantic Tree of Astronomical Variables and Transients



The Tsunami Wave of the Future



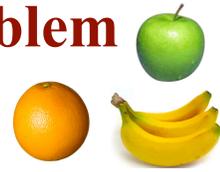
- Now: data streams of **~ 0.1 TB / night**, **~ 10² transients / night** (CRTS, PQ, PTF, various SN surveys, asteroid surveys)
- Forthcoming on a time scale **~ 1 - 5 years**: **~ 1 TB / night**, **~10⁴ transients / night** (PanSTARRS, Skymapper, VISTA, VST...)
- Forthcoming in **~ 8 - 10 years**: LSST, **~ 30 TB / night**, **~ 10⁵ - 10⁶ transients / night**
- Observational follow-up needs:
 - Rapid photometric/positional monitoring
 - Rapid spectroscopy
 - Information/computation infrastructure

A major, qualitative change!

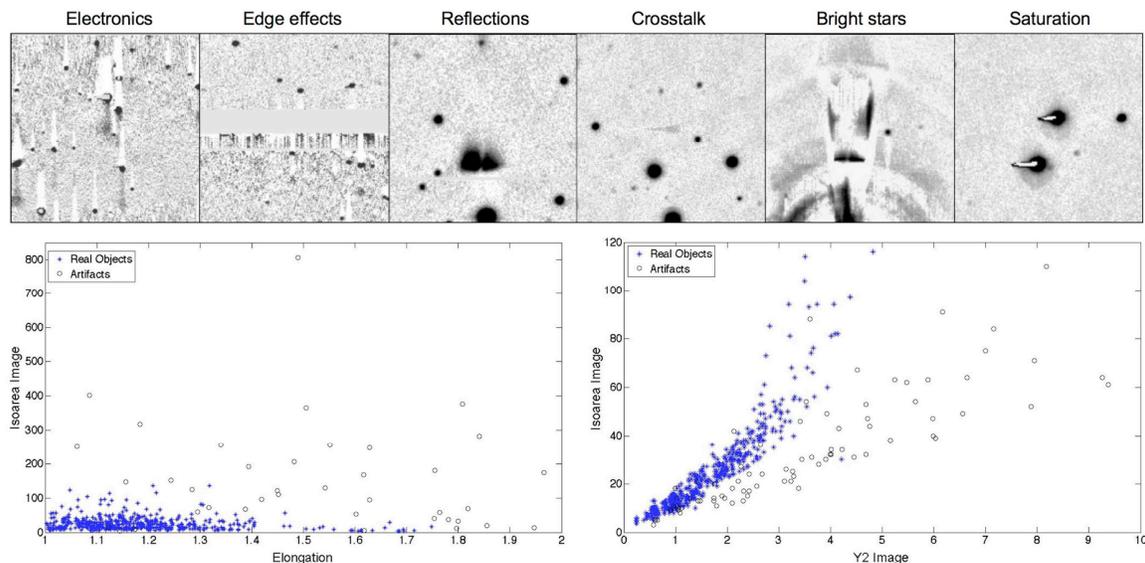
Transient classification technologies are essential

Event Classification is a **Hard Problem**

- Classification of transient events is essential for their astrophysical interpretation and uses
 - Must be done in real time and iterated dynamically
- Human classification is already unsustainable, and will not scale to the future Petascale data streams
- This is hard:
 - Data are sparse and heterogeneous: feature vector approaches do not work; using Bayesian approach
 - Completeness vs. contamination ☹️
 - Follow-up resources are expensive and/or limited: only the most interesting events
 - Iterate classifications dynamically as new data come in
- Traditional DP pipelines do not capture a lot of the relevant contextual information, prior/expert knowledge, etc.



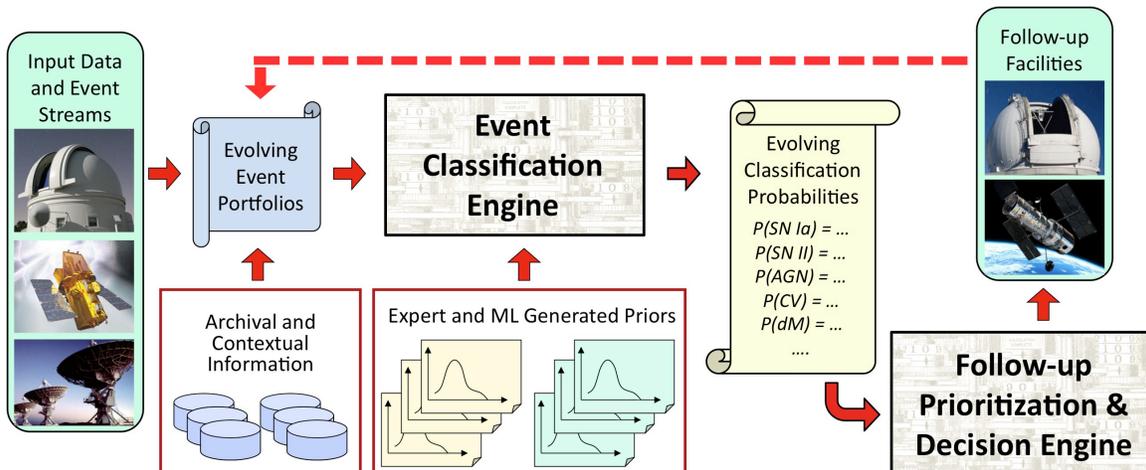
Automated Detection of Artifacts



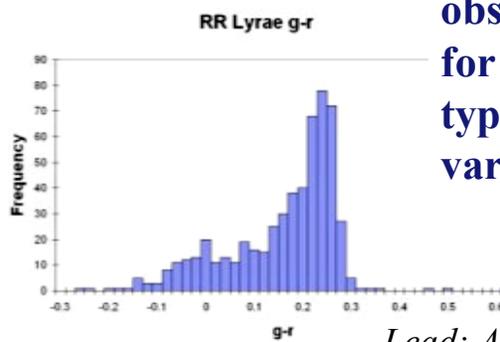
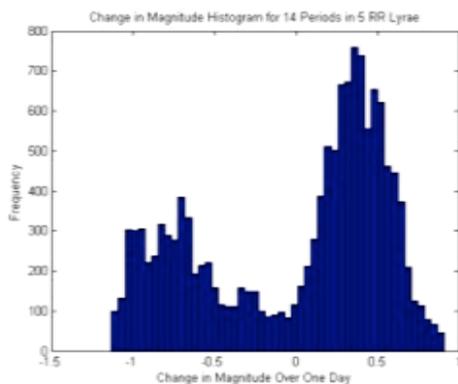
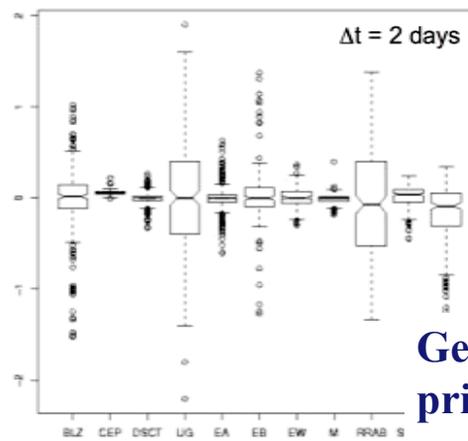
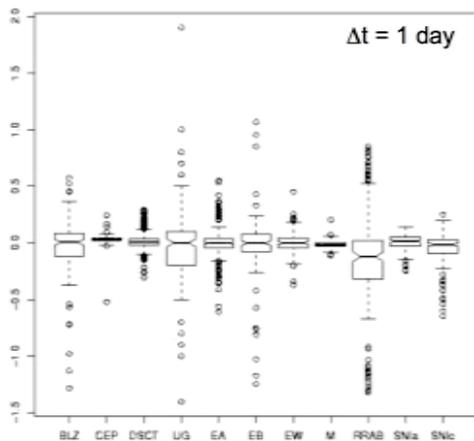
Automated classification and rejection of artifacts masquerading as transient events in the PQ survey pipeline, using a Multi-Layer Perceptron ANN

Lead:
C. Donalek

Towards the Automated Event Classification



- Incorporation of the contextual information (archival, and from the data themselves) is essential
- Automated prioritization of follow-up observations, given the available resources and their cost
- A dynamical, iterative system



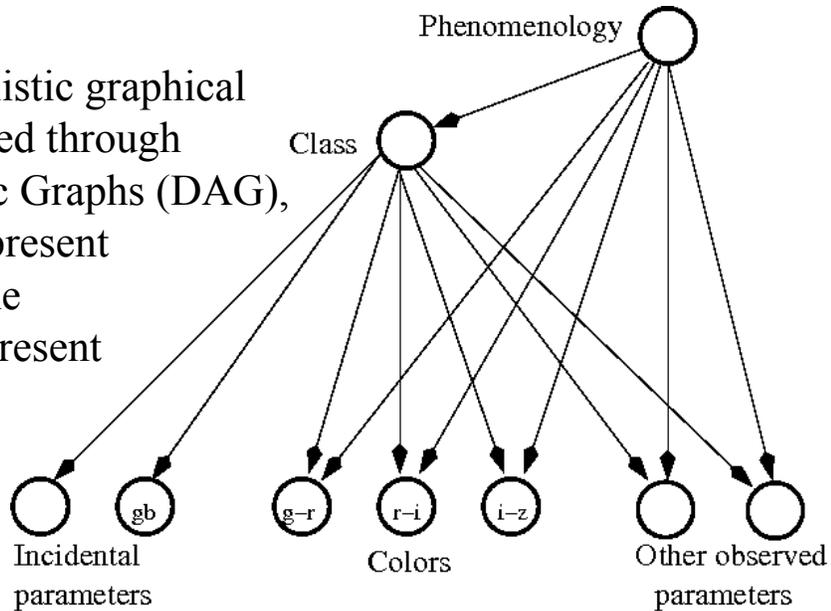
Generating priors for various observables for different types of variables

Lead: A. Mahabal

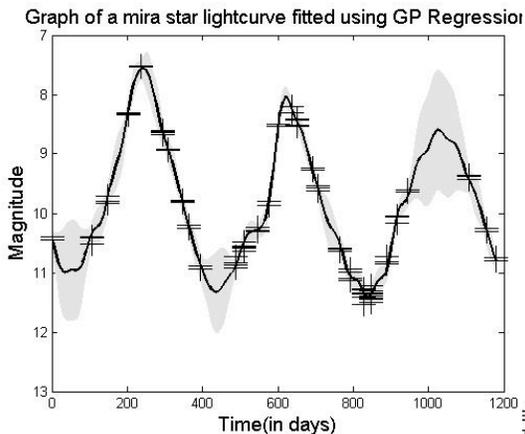
Bayesian Networks (BN)

Bayesian methodology is desirable and attractive for this task, since it can deal with missing or heterogeneous data

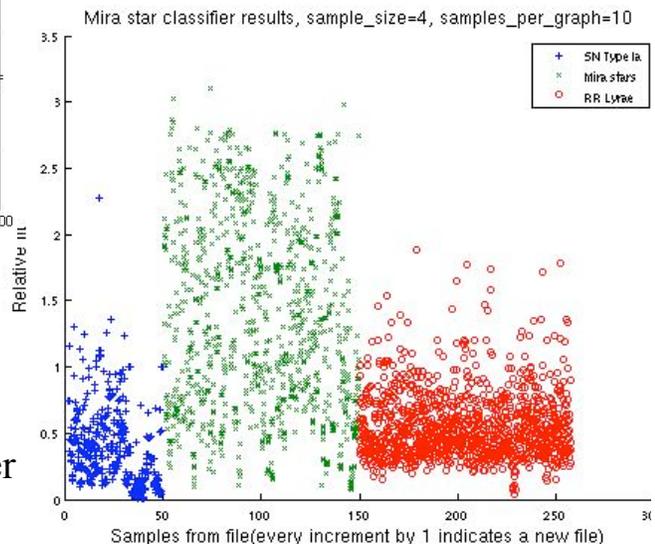
BN is a probabilistic graphical model represented through Directed Acyclic Graphs (DAG), whose nodes represent variables, and the missing arcs represent conditional independence assumptions



Gaussian Process Regression (GPR)



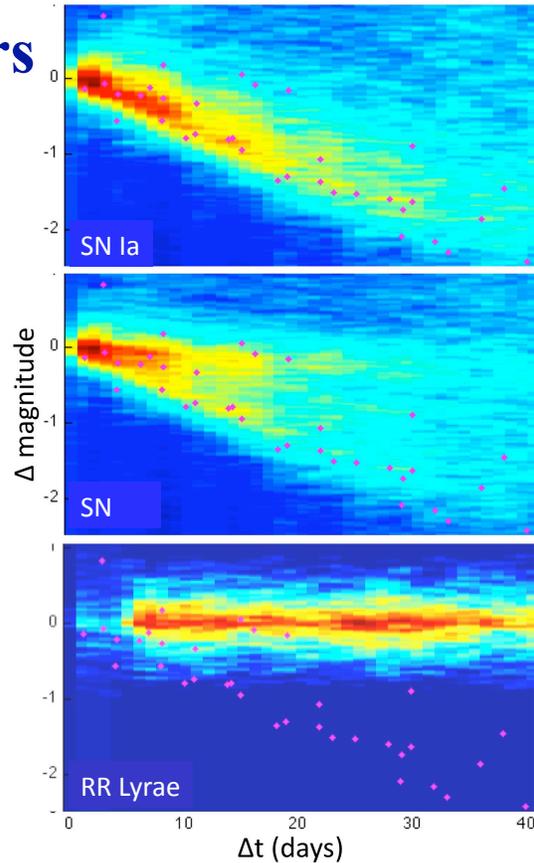
A Mira variable star light curve fitted using GPR



Given 4 random points from the light curve of a Mira variable, the probability of it being a Mira variable is higher than, say, a SN

2D Light Curve Priors

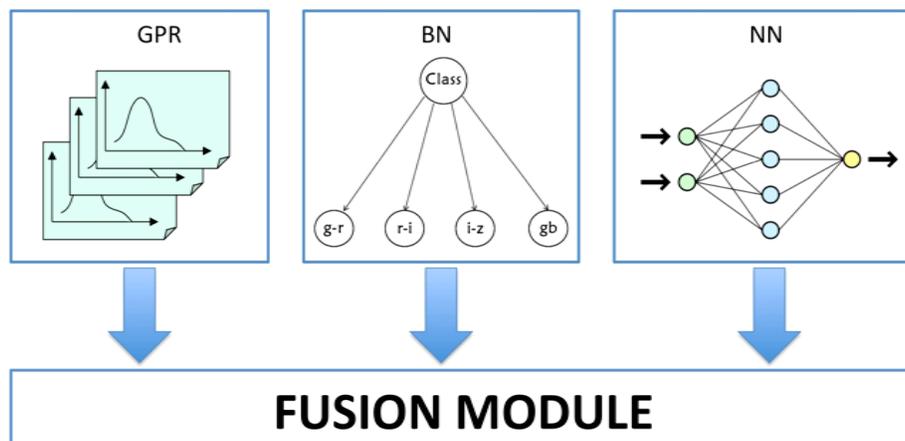
- For any pair of light curve measurements, compute the Δt and Δm , make a 2D histogram
 - Note: N independent measurements generate N^2 correlated data points
- Compare with the priors for different types of transients
- Repeat as more measurements are obtained, for an evolving, constantly improving classification.



Lead: B. Moghaddam

Fusion Module

Colors and light curve information can be combined in one network. This "fusion module" combines the probabilistic results from each constituent classifier



Exploring a variety of techniques for optimal classification fusion:

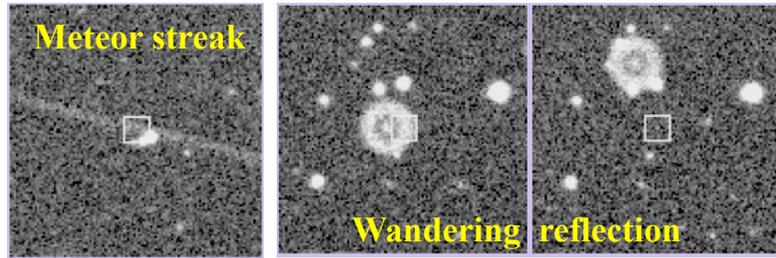
$\mathcal{P}_{\text{class}}$

Markov Logic Networks, Diffusion Maps, Multi-Arm Bandit, Sleeping Expert...

Harvesting the Human Pattern Recognition

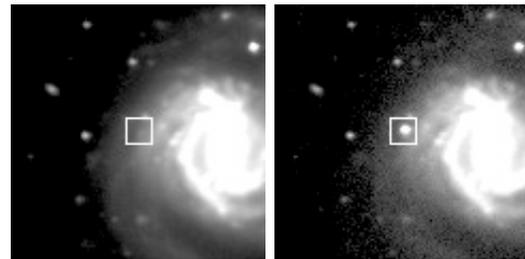
Recognizing the artifacts (false transients)

Contextual information is essential



A more sophisticated case uses a **prior (expert) knowledge:**

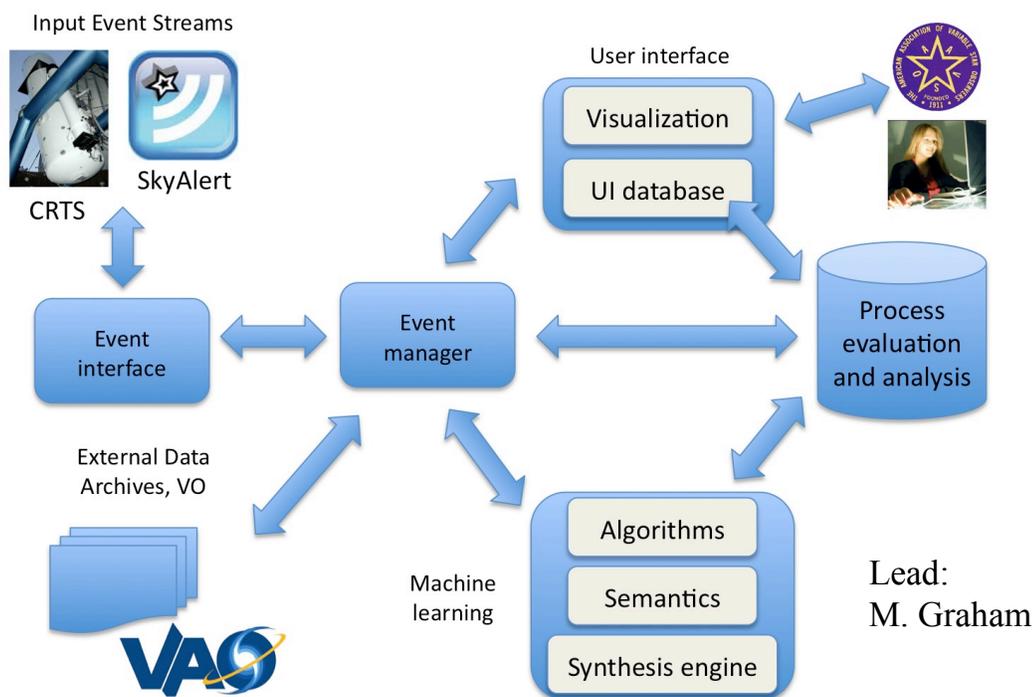
Star-like transient apparently associated with a non-coincident galaxy a likely Supernova

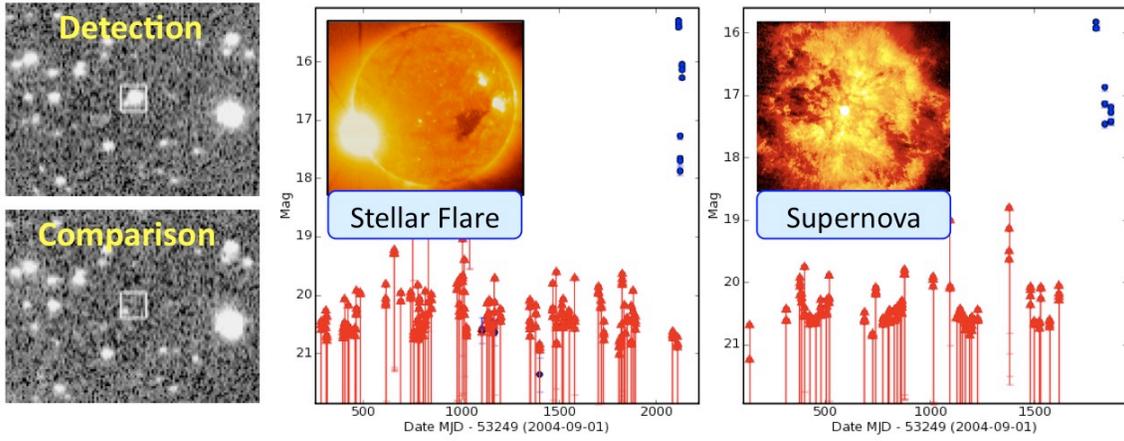


Spiral host galaxy
a possible Type II

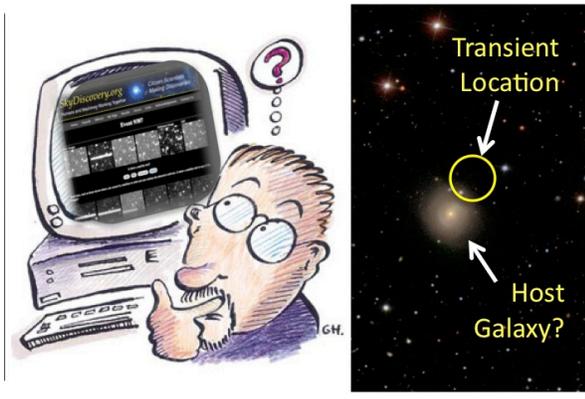
How to capture this and teach a machine to do the same thing?

AstroCollation: Towards Harvesting Human Pattern Recognition and Domain Expertise

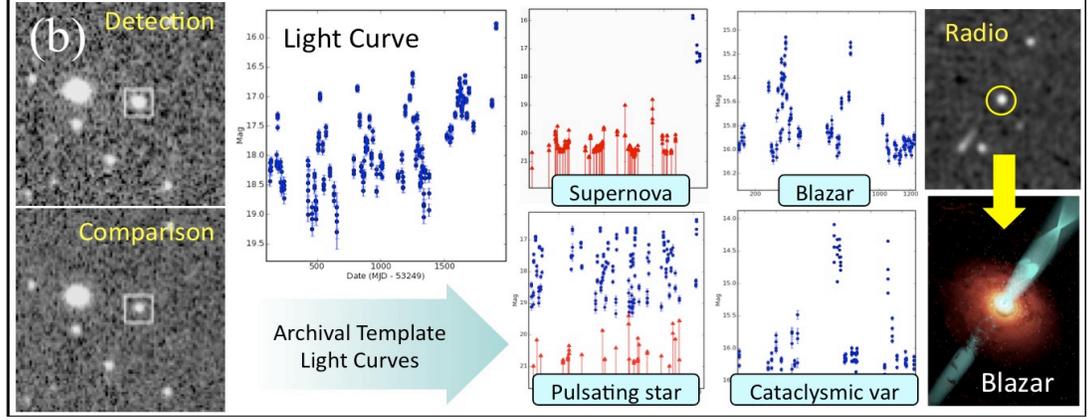
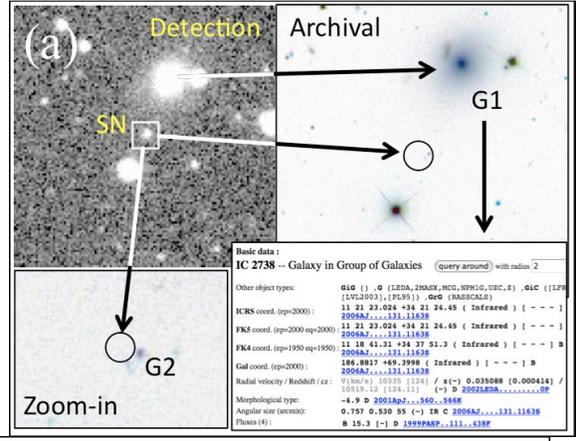




Use Case Scenario:
 Light curves are ambiguous, but the presence of a possible host galaxy suggest that it is a Supernova



Use Case Scenarios:
 (a) Archival data on potential host galaxies provides the more likely choice,
 (b) Presence of a radio source discriminates between a CV and a blazar



SkyDiscovery.org

Humans and Machines Working Together

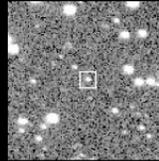
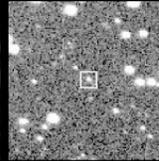
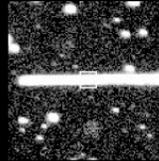
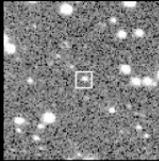


Citizen Scientists
Making Discoveries

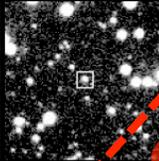
- Home
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Event 9387

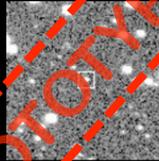
New Images



Reference Image



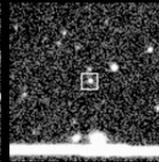
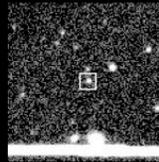
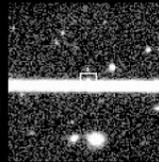
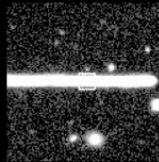
GIF



Is there a satellite trail?

- Yes
- No
- Unsure
- Help

Bold lines, such as those shown below, are caused by satellites in orbit and can confuse the detection software. Is there a satellite trail in any of the images?



Citizen Science Supernova Hunt



An Open Optical Transient Survey

- Home
- Download New
- Download Diff
- Download Ref
- Contact

See the celestial context in the WorldWide Telescope

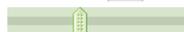


- Images
- Parameters

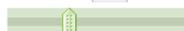
Images of ESO145-16 RA= 327.29583 Dec= -59.03694

Image Scaling

Brightness: -30



Contrast: 0.3



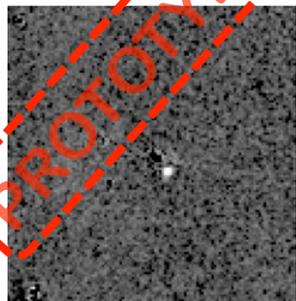
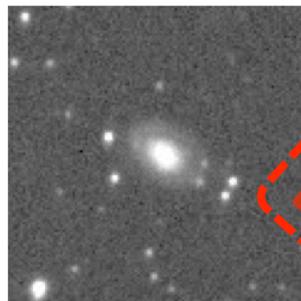
Legacy: Invert:

New:

Reference:

Difference:

Adjust B&C Reset



- Back
- Next

RA Dec

RA 327.2849 Dec -59.0628

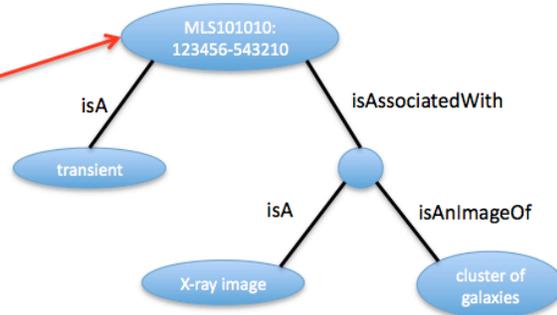
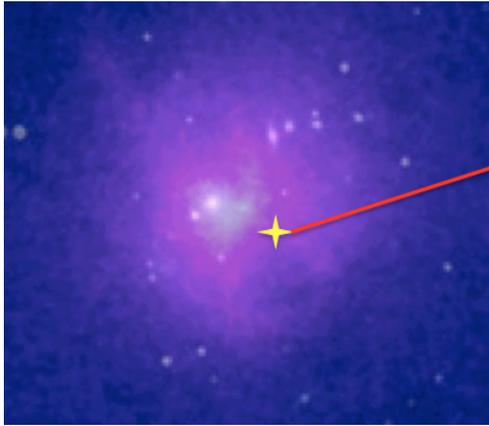
Lead: A. Drake

New Image

Reference Image

Difference Image

Developing an Interface Between Carbon-Based and Silicon-Based Minds

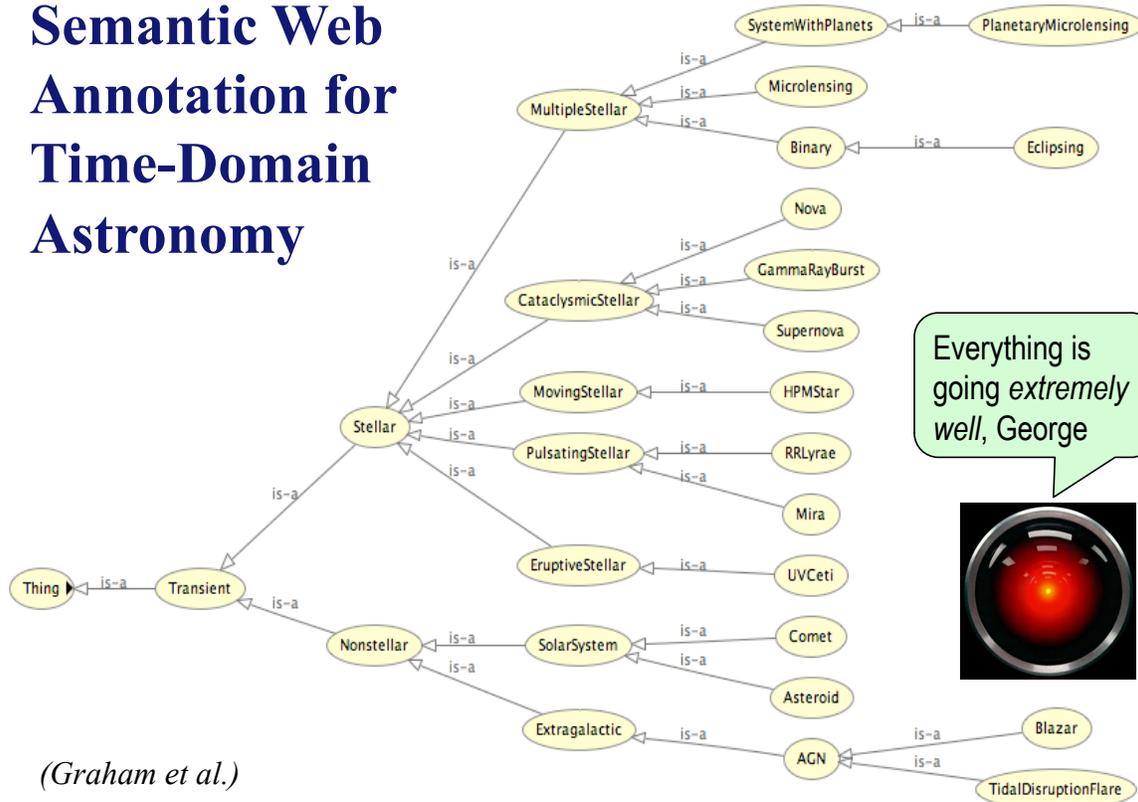


Human-annotated images (via *SkyDiscovery.org*)

- ⇒ Semantic descriptors
- ⇒ Machine processing
- ⇒ Novel algorithms



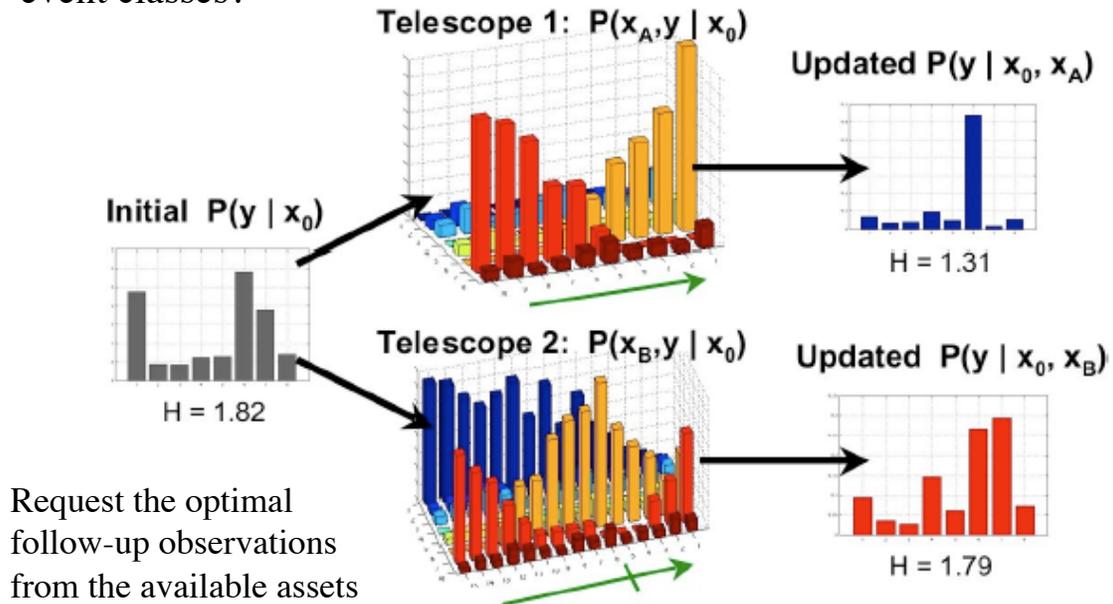
Semantic Web Annotation for Time-Domain Astronomy



(Graham et al.)

Automating the Optimal Follow-Up

For the potentially most interesting events, what type of follow-up data has the greatest potential to discriminate among the competing event classes?



Summary

- Real-time mining of massive data streams offers great opportunities and challenges
 - Synoptic sky surveys and real-time astronomy are an excellent science & technology testbed
- We are making progress on real-time, automated, iterated event classification
 - *Not your grandma's classification problem!*
 - Sparse and heterogeneous data, real time, dynamically iterated, resource-limited
 - Next: an automated decision making for optimal follow-up observations
- Harvesting human pattern recognition skills and expertise using citizen science
- A broader relevance for a real-time mining of massive data streams

