



JWST is NASA's flagship mission

- Endorsed by the National Academy as a top priority for space-based astronomy in the 2000 Decadal Survey, science priorities reaffirmed in 2010 Decadal Survey
- Launch in 2018 out to L2, 5 year baseline mission with extension to 10 years
- Operates at cryogenic temperatures
- New generation of complex instrumentation to ensure diverse modes of operation without servicing:
 - >40 imaging filters
 - 7 coronagraphs for high-resolution imaging
 - 8 different types of spectroscopic modes

4 instruments for wide range of science all delivered to Goddard Space Flight Center



Near Infrared Camera

Near Infrared Spectrograph

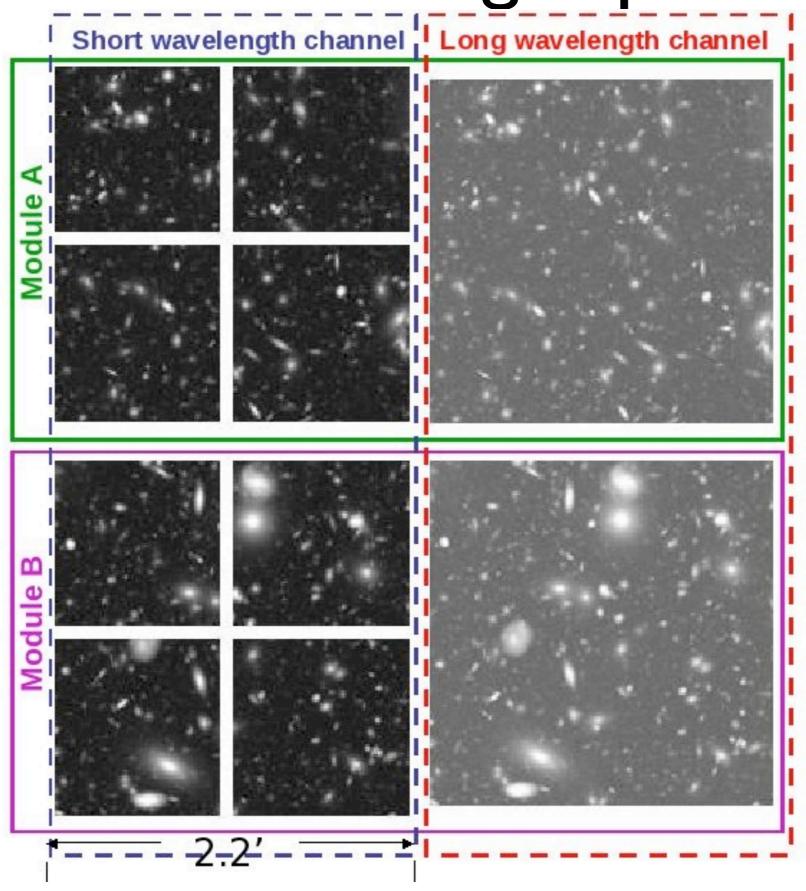


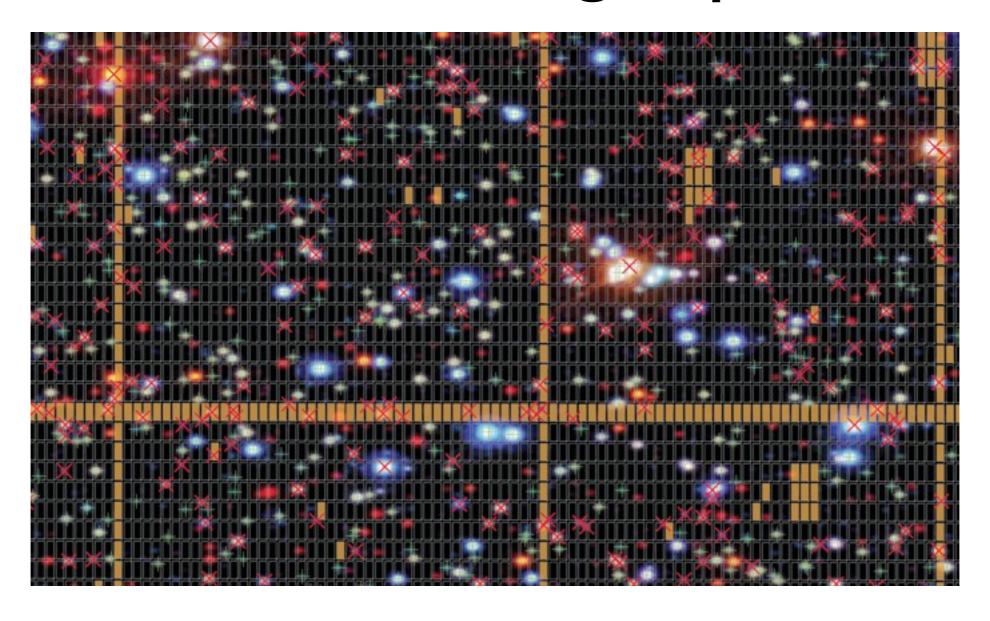


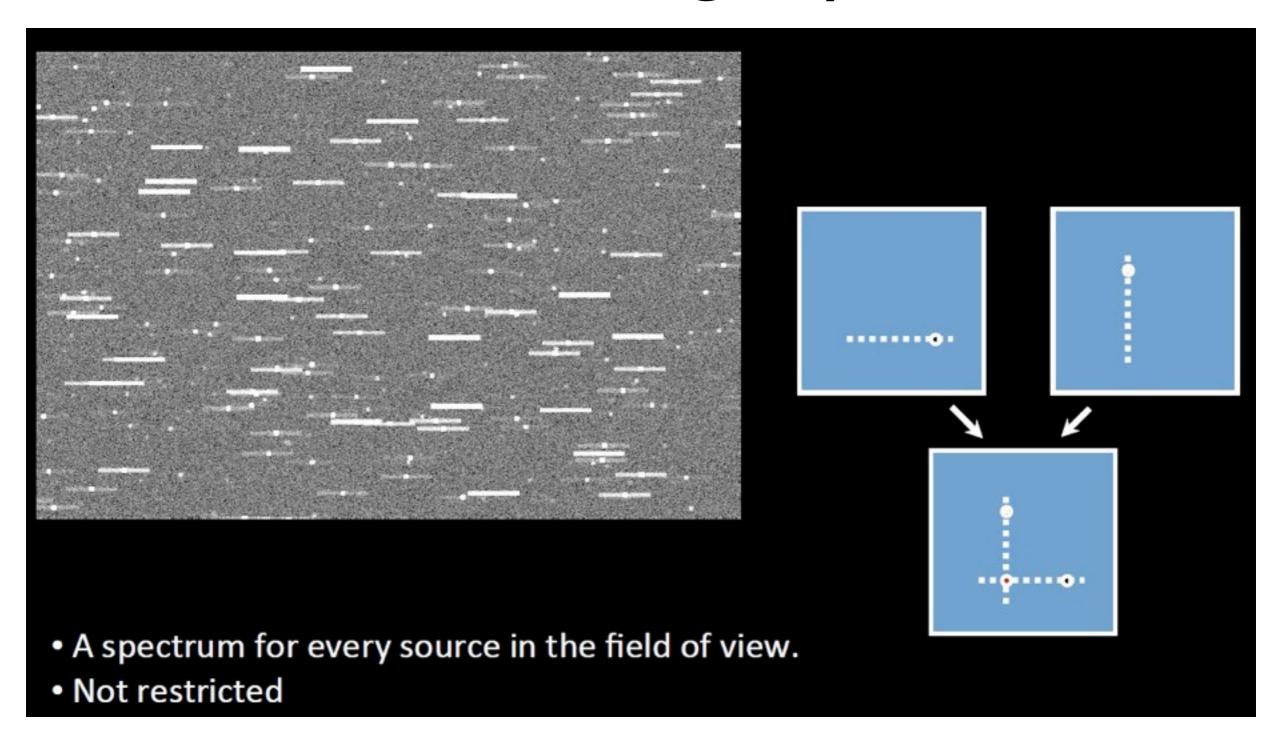
Mid Infrared Instrument

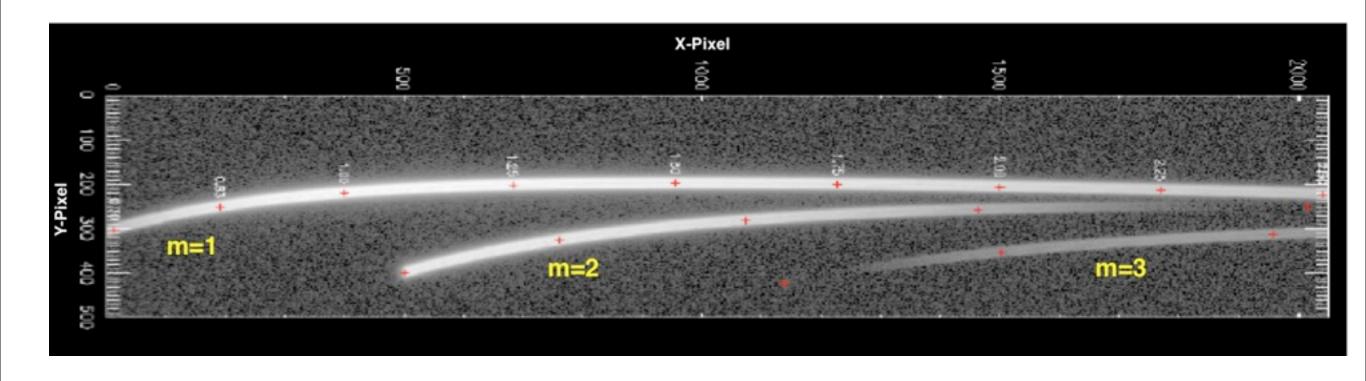
Near Infrared Imager & Slitless Spectrograph

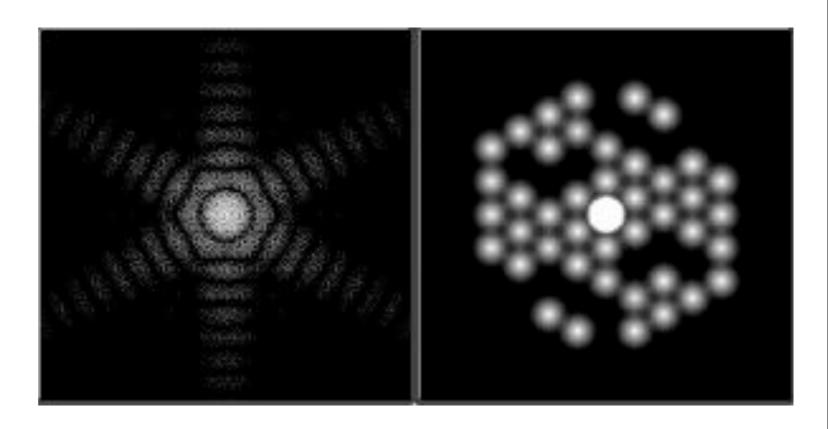


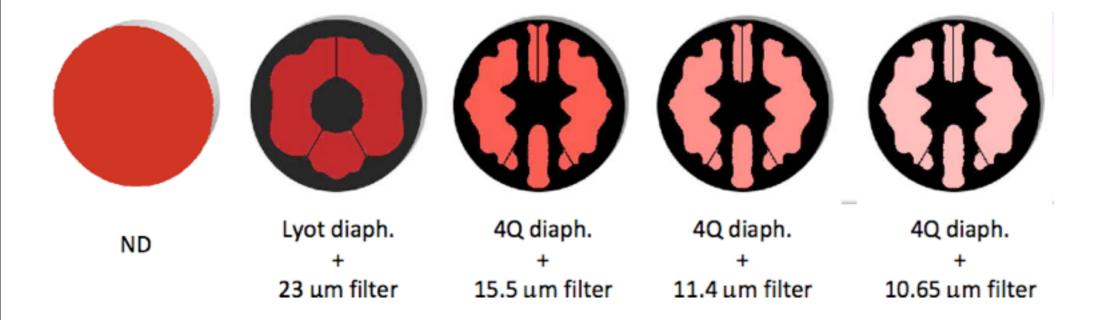


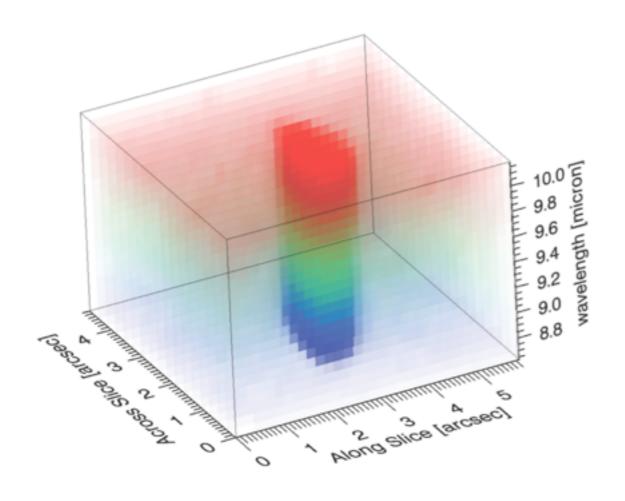












JWST Imaging Modes

Mode	Instrument	Wavelength (micron)	Pixel Scale (arcsec)	Field of View
	NIRCam	0.6-2.3	0.032	2.2×4.4'
Imaging	NIRCam	2.4-5.0	0.065	2.2×4.4'
	NIRISS	0.9-5.0	0.065	2.2×2.2'
	MIRI	5.0-28	0.11	1.23×1.88'
Aperture Mask Interferometry	NIRISS	3.8-4.8	0.065	2.2×2.2'
	NIRCam	0.6-2.3	0.032	20×20"
Coronography	NIRCam	2.4-5.0	0.065	20×20''
	MIRI	10.65	0.11	24×24"
	MIRI	11.4	0.11	24×24"
	MIRI	15.5	0.11	24×24"
	MIRI	23	0.11	30×30"

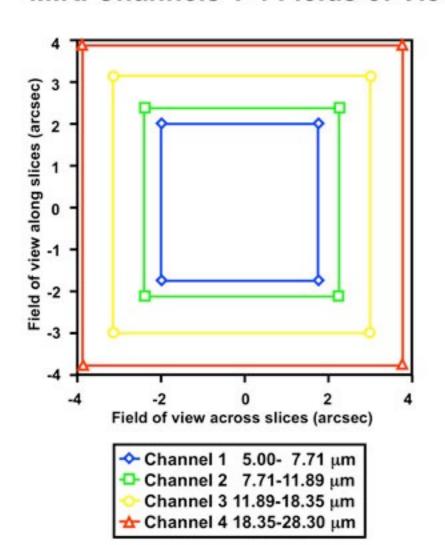
Friday, June 20, 14

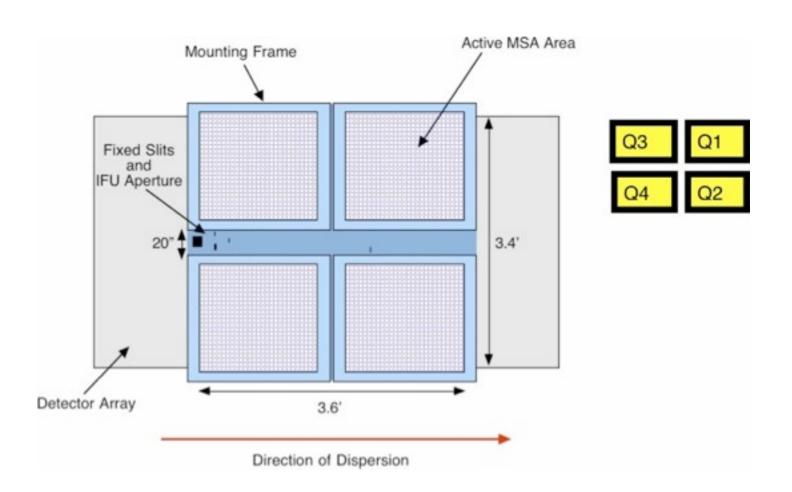
JWST Spectroscopic Modes

	<i>J</i>			
Mode	Instrument	Wavelength (micron)	Resolving Power $(R=\lambda/\Delta\lambda)$	Field of View
Slitless	NIRISS	1.0-2.5	150	2.2×2.2'
Spectroscopy	NIRISS	0.6-2.5	700	single object
	NIRCam	2.4-5.0	2000	2.2×2.2'
Multi-Object Spectroscopy	NIRSpec	0.6-5.0	100 1000 2700	3.4x3.4' with 250k 0.2x0.5" microshutters
Single Slit Spectroscopy	NIRSpec	0.6-5.0	100 1000 2700	slits with 0.4×3.8" 0.2×3.3"
	MIRI	5.0- ~ I 4.0	~100 at 7.5 microns	0.6×5.5" slit
	NIRSpec	0.6-5.0	100 1000 2700	3.0×3.0"
IFU	MIRI	5.0-7.7	3500	3.0×3.9"
Spectroscopy	MIRI	7.7-11.9	2800	3.5×4.4"
	MIRI	11.9-18.3	2700	5.2×6.2"
, , ,	MIRI	18.3-28.8	2200	6.7×7.7"

IFUs have small fields of view, obtain low- and medium- resolution spectra

MIRI Channels 1-4 Fields of View





NIRSpec IFU FOV is 3"x3"

Requirements on Data Management System (DMS) Tools I.

- utilize FITS data format
- utilize VO data access protocols
- provide a capability for users to run individual Calibration Pipeline steps on JWST data
- distribute Data Analysis Software to users on request
- provide a capability to display and analyze Archived Science Data and Science Data Products
- provide a capability for JWST users to perform data analysis functions on Science Data Products electronically from remote locations
- provide a capability for visualizing JWST data

Requirements on Data Management System (DMS) Tools II.

- provide a capability for arithmetic operations on collections of JWST datasets.
- provide the capability of comparing JWST Calibrated Science Data with models that can be cast into arrays with the same dimensionality as the Calibrated Science Data
- provide the capability to combine JWST datasets including tools to determine dataset registration and optimal combination of I-d spectra, 2-d images and 3-d spectral data cubes.
- provide tools to support JWST coronagraphic science
- provide tools to enable analysis of JWST IFU data
- provide facilities for analyzing JWST spectral data
- provide tools for modeling JWST instrumental effects.

Current Data Analysis Efforts

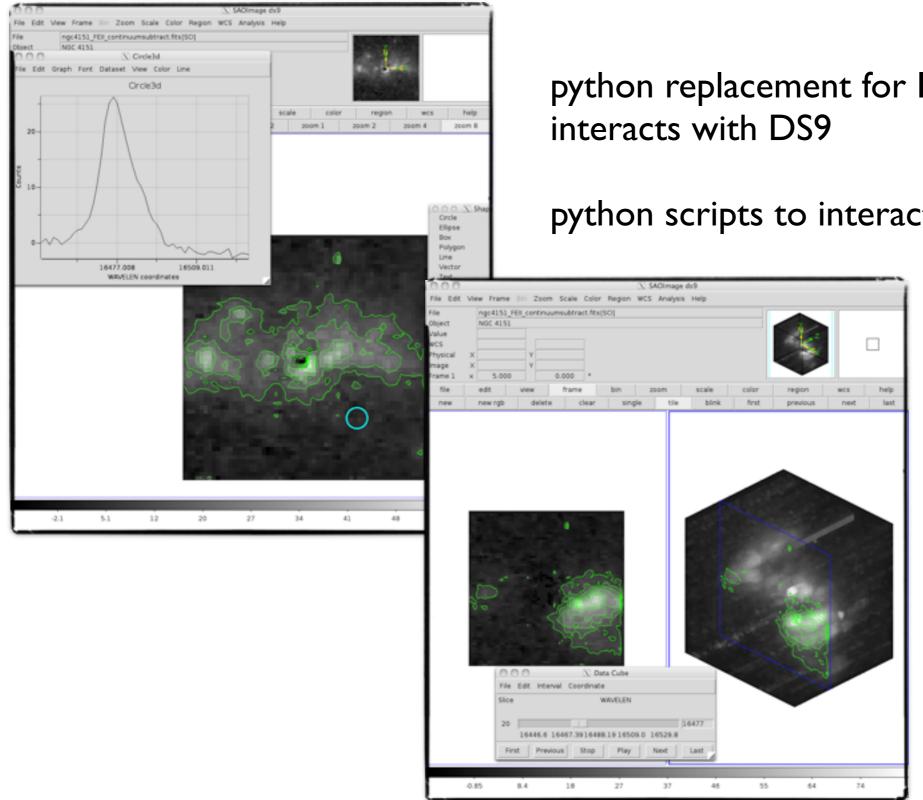
- python based, modular
 - compatible with Python 2.x and 3.x
 - PEP8 coding style standard
 - use standard core libraries for
 - FITS IO
 - ASCII tables
 - VOTables
 - Physical Units/Quantities
 - WCS
 - Coordinates
 - Fitting/Models
 - Science Data Objects
 - use Travis/Shining Panda Continuous Integration tools
- designed to be generic, extensible: AstroPy collaboration

IFU Visualization

Work on next generation visualization tools

DS9 3D module

multiple dimensions: I-D spectra, 2-D images/spectra, 3-D IFU datacubes



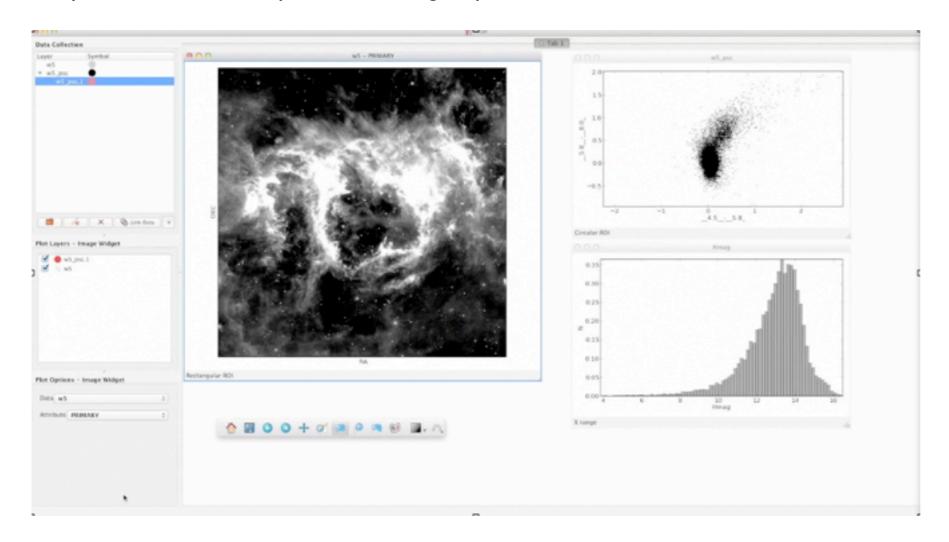
python replacement for IRAF imexamine task

python scripts to interact with visualization tool

IFU Visualization

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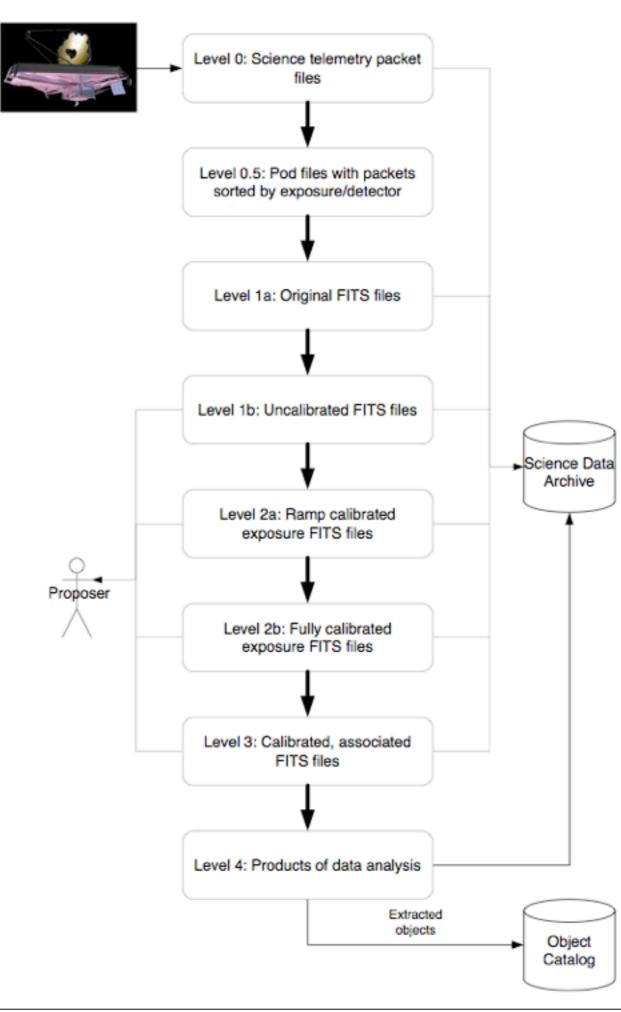


interactive layer on python matplotlib offers linked view functionality implementing cube visualization, scripting ability

GLUE

Questions to consider

- What data products will be delivered?
- Is there a plan to do this in a VOcompatible manner?
- For data cubes, is there an intent to provide users with a capability to modify projectprovided data before downloading it?



NIRCam Associations

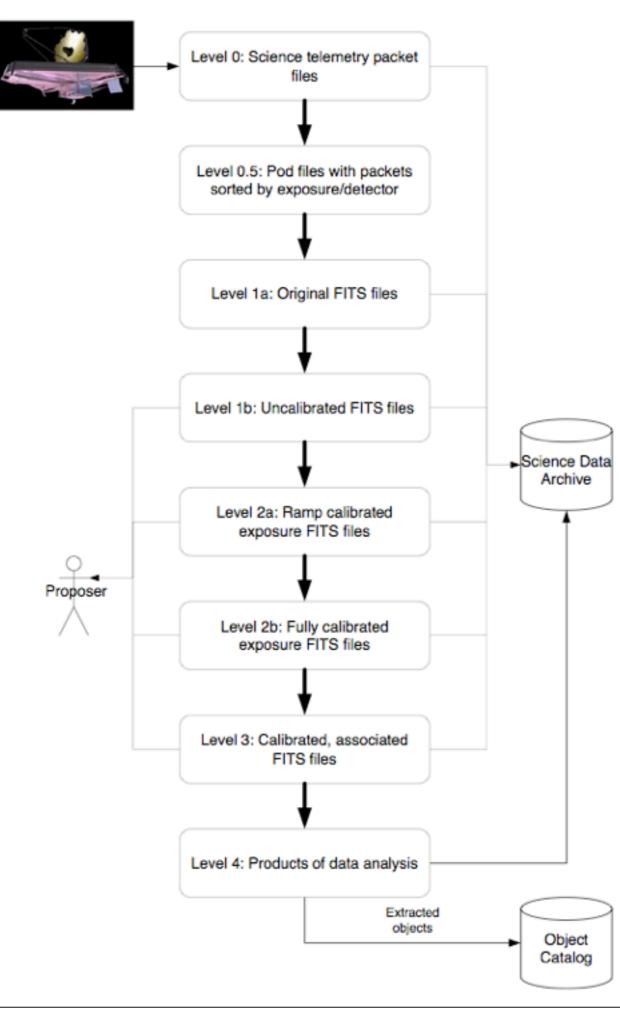
Data products

- mosaic combined image (multiple detectors)
- dither combined image (if no detector)

Related files

- individual exposures (single detector; levels Ia, Ib, 2a, 2b)
- coronagraphic target acquisition
- jitter data

- flats (taken with science exposure)
- persistence darks (in the distant future)



NIRSpec Associations

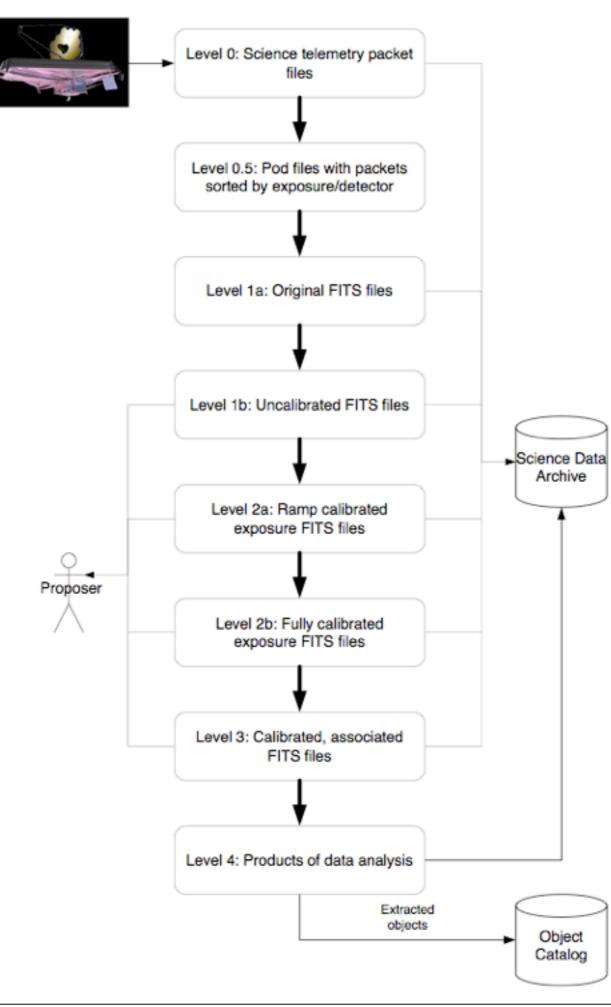
Data products

- extracted I-d and 2-d spectra over both detectors
- MSA Source Catalog (level 4)

Related files

- individual exposures (single detector; levels Ia, Ib, 2a, 2b)
- target acquisition images
- confirmation images
- jitter data
- MSA Pre-image
- MSA Metadata file
- reference star file
- background exposures (may be shared among exposures)

- auto-flats
- wave-cals



MIRI Associations

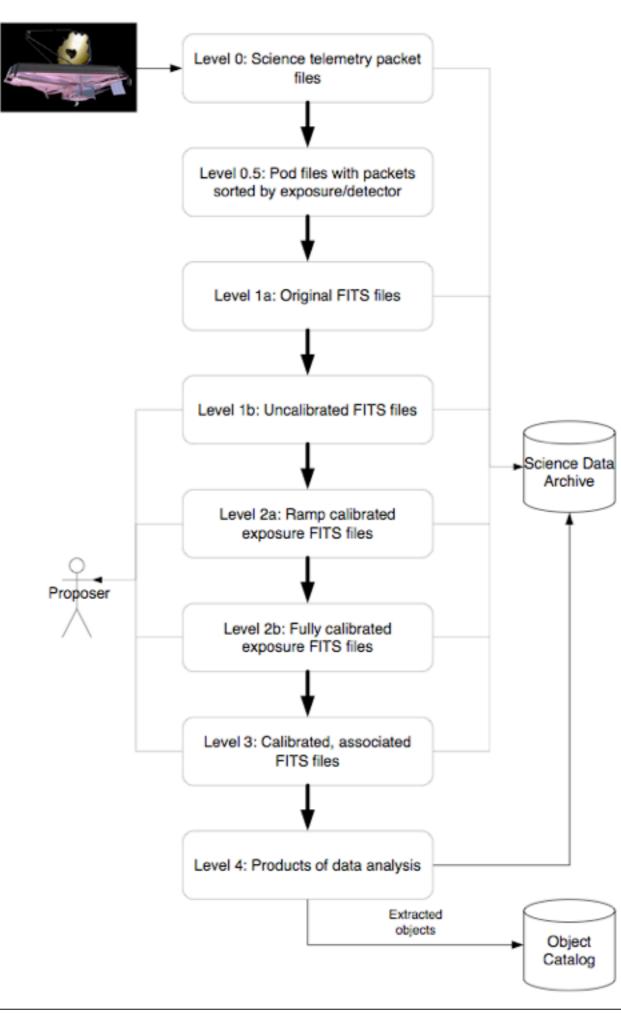
Data products

- mosaic combined image
- dither combined image (if no mosaic)
- extracted I-d and 2-d spectra
- extracted I-d and 2-d spectra over both IFU detectors

Related files

- individual exposures (single detector; levels Ia, Ib, 2a, 2b)
- coronagraphic target acquisition images
- spectroscopic target acquisition images
- jitter data

- imager internal flats
- MRS internal flats
- chopped background image
- persistence darks (in the distant future)



NIRISS Associations

Data products

- mosaic combined image
- dither combined image (if no mosaic)

Related files

- individual exposures (single detector; levels Ia, Ib, 2a, 2b)
- jitter data

- point source image for PSF characterization
- persistence darks (in the distant future)